

# Bootstrapping Robust EPSI with coarsely sampled data

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University of British Columbia

## Outline

1. Intro/review of Robust EPSI algorithm
2. “Bootstrapping” by deliberate data decimation
3. Application to subsampled data
4. ... bonus slides? (if time permits)

# Robust EPSI primer

## From SRME to Robust EPSI

Based on **Estimation of Primaries by Sparse Inversion** (van Groenestijn and Verschuur, 2009)

true primary wavefield

SRME-produced primary

$$\mathbf{P}_o = \mathbf{P} - A(f)\mathbf{P}_o\mathbf{P}$$

$\mathbf{P}$  total up-going wavefield

$\mathbf{P}_o$  primary wavefield

$A(f)$  “matching” operator



# From SRME to Robust EPSI

Based on **Estimation of Primaries by Sparse Inversion** (van Groenestijn and Verschuur, 2009)

true primary wavefield

SRME-produced primary

$$\mathbf{P}_o \approx \mathbf{P} - A(f) \mathbf{P} \mathbf{P}$$

SRMP

$\mathbf{P}$  total up-going wavefield

$\mathbf{P}_o$  primary wavefield

$A(f)$  “matching” operator

## From SRME to Robust EPSI

Based on **Estimation of Primaries by Sparse Inversion** (van Groenestijn and Verschuur, 2009)

**adaptive subtraction**

$$\min_A \sum_f \|\mathbf{P} - A(f) \mathbf{P}\|$$

**SRMP**

$\mathbf{P}$  total up-going wavefield

$\mathbf{P}_o$  primary wavefield

$A(f)$  “matching” operator

## From SRME to Robust EPSI

Based on **Estimation of Primaries by Sparse Inversion** (van Groenestijn and Verschuur, 2009)

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$$\mathbf{P}_o = \mathbf{P} - A(f)\mathbf{P}_o\mathbf{P}$$

$\mathbf{P}$  total up-going wavefield

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$A(f)$  “matching” operator

## From SRME to Robust EPSI

Based on **Estimation of Primaries by Sparse Inversion** (van Groenestijn and Verschuur, 2009)

recorded data      predicted data from SRME

$$\mathbf{P} = \mathbf{P}_o + A(f)\mathbf{P}_o\mathbf{P}$$

$\mathbf{P}$       total up-going wavefield  
 $\mathbf{P}_o$       primary wavefield  
 $A(f)$       “matching” operator



## From SRME to Robust EPSI

Based on **Estimation of Primaries by Sparse Inversion** (van Groenestijn and Verschuur, 2009)

recorded data      predicted data from SRME

$$\mathbf{P} = \mathbf{P}_o + A(f)\mathbf{P}_o\mathbf{P}$$

$$\begin{aligned}\mathbf{P}_o &= \mathbf{Q}\mathbf{G} \\ A(f) &= -\mathbf{Q}^{-1}\end{aligned}$$

**P**      total up-going wavefield  
**Q**      down-going source signature  
**G**      primary impulse response

# From SRME to Robust EPSI

Based on **Estimation of Primaries by Sparse Inversion** (van Groenestijn and Verschuur, 2009)

recorded data      predicted data from SRME

$$\mathbf{P} = \mathbf{Q}\mathbf{G} - \mathbf{G}\mathbf{P}$$

- P**      total up-going wavefield
- Q**      down-going source signature
- G**      primary impulse response



## From SRME to Robust EPSI

Based on **Estimation of Primaries by Sparse Inversion** (van Groenestijn and Verschuur, 2009)

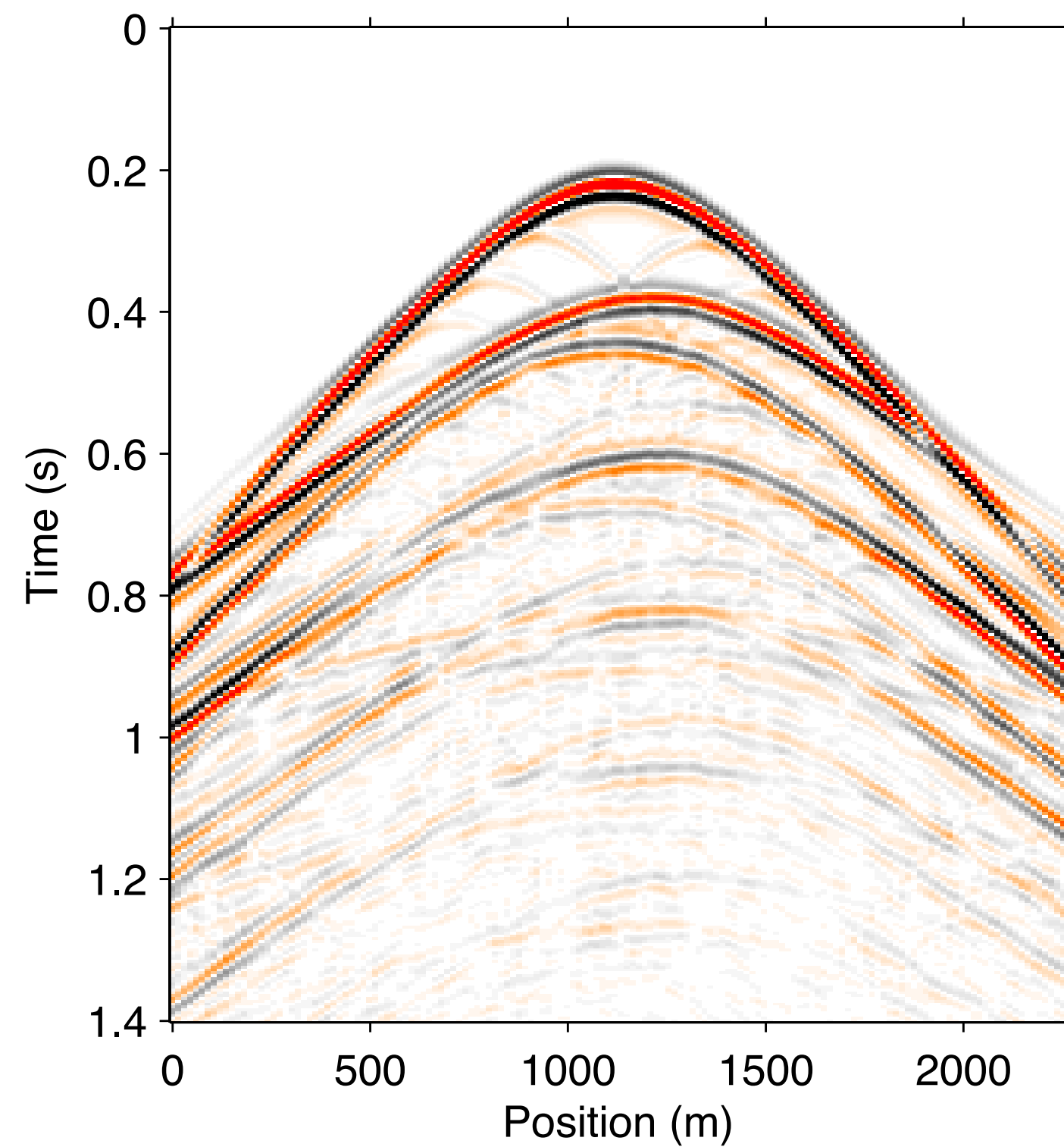
recorded data      predicted data from SRME

$$\mathbf{P} = \mathbf{Q}\mathbf{G} - \mathbf{G}\mathbf{P}$$

**Inversion objective:**

$$f(\mathbf{G}, \mathbf{Q}) = \frac{1}{2} \|\mathbf{P} - (\mathbf{Q}\mathbf{G} - \mathbf{G}\mathbf{P})\|_2^2$$

# From SRME to Robust EPSI



Primaries by Sparse Inversion (van Groenestijn and Verschuur, 2009)

observed data

predicted data from SRME

$$\mathbf{P} = \mathbf{Q}\mathbf{G} - \mathbf{G}\mathbf{P}$$

$$f(\mathbf{G}, \mathbf{Q}) = \frac{1}{2} \|\mathbf{P} - (\mathbf{Q}\mathbf{G} - \mathbf{G}\mathbf{P})\|_2^2$$

# From SRME to Robust EPSI

Based on **Estimation of Primaries by Sparse Inversion** (van

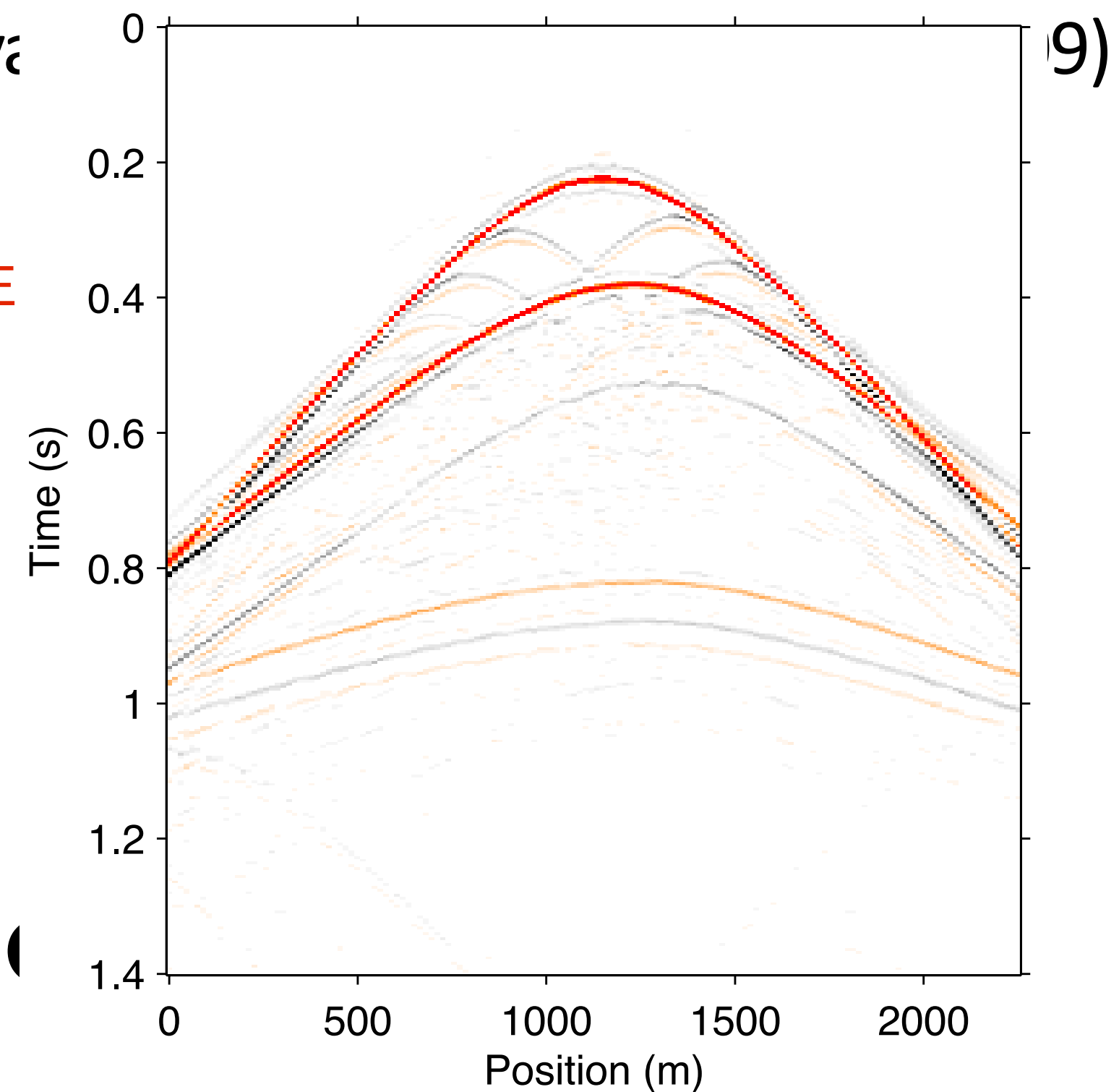
recorded data

predicted data from SRME

$$\mathbf{P} = \mathbf{Q}\mathbf{G} - \mathbf{G}\mathbf{P}$$

**Inversion objective:**

$$f(\mathbf{G}, \mathbf{Q}) = \frac{1}{2} \|\mathbf{P} - (\mathbf{Q}\mathbf{G} - \mathbf{G}\mathbf{P})\|$$



9)

# From SRME to Robust EPSI

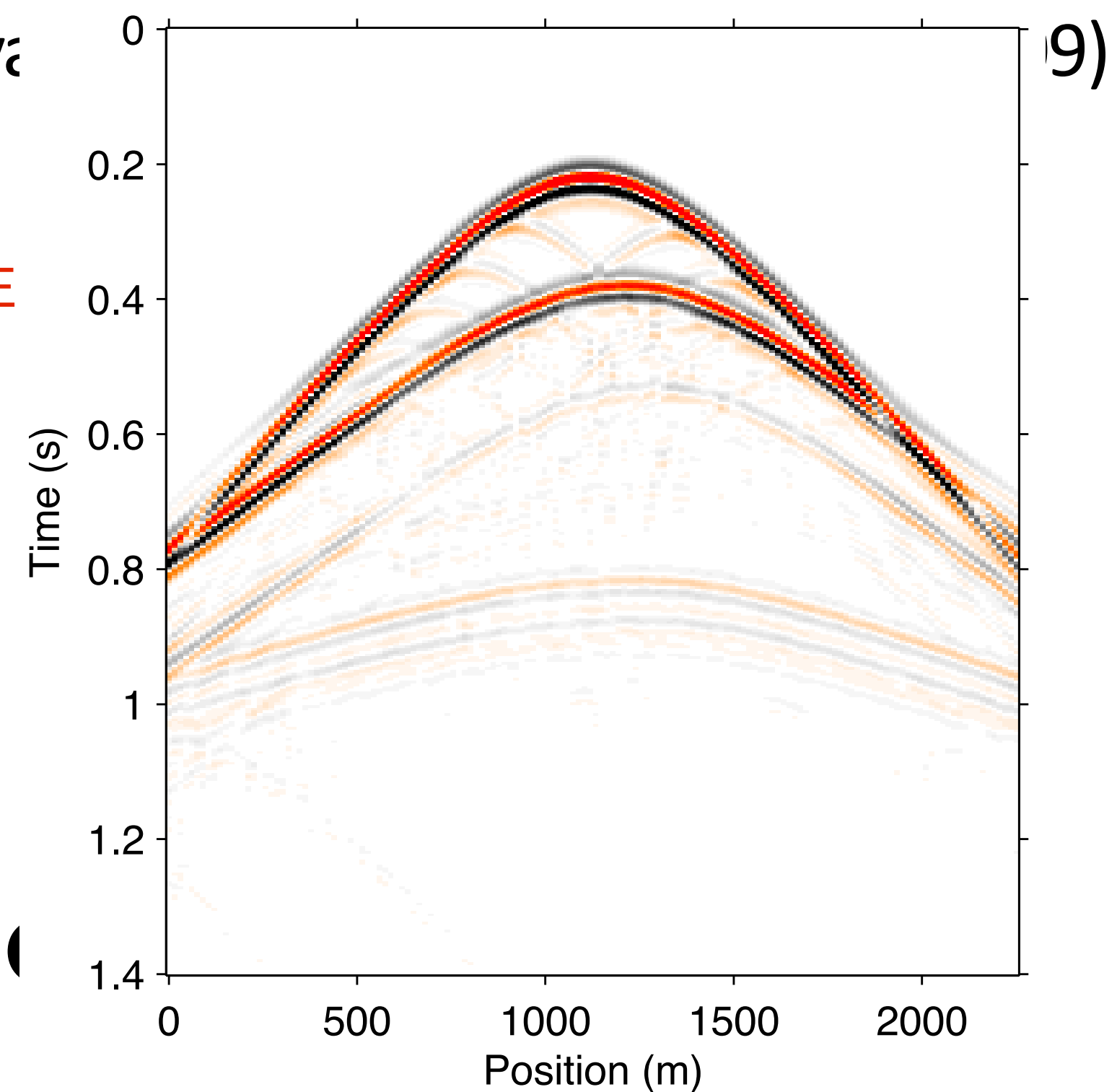
Based on **Estimation of Primaries by Sparse Inversion** (van

$$\mathbf{P} = \underbrace{\mathbf{Q}\mathbf{G}}_{\text{predicted data from SRME}} - \mathbf{G}\mathbf{P}$$

recorded data

**Inversion objective:**

$$f(\mathbf{G}, \mathbf{Q}) = \frac{1}{2} \|\mathbf{P} - (\mathbf{Q}\mathbf{G} - \mathbf{G}\mathbf{P})\|$$





# From SRME to Robust EPSI

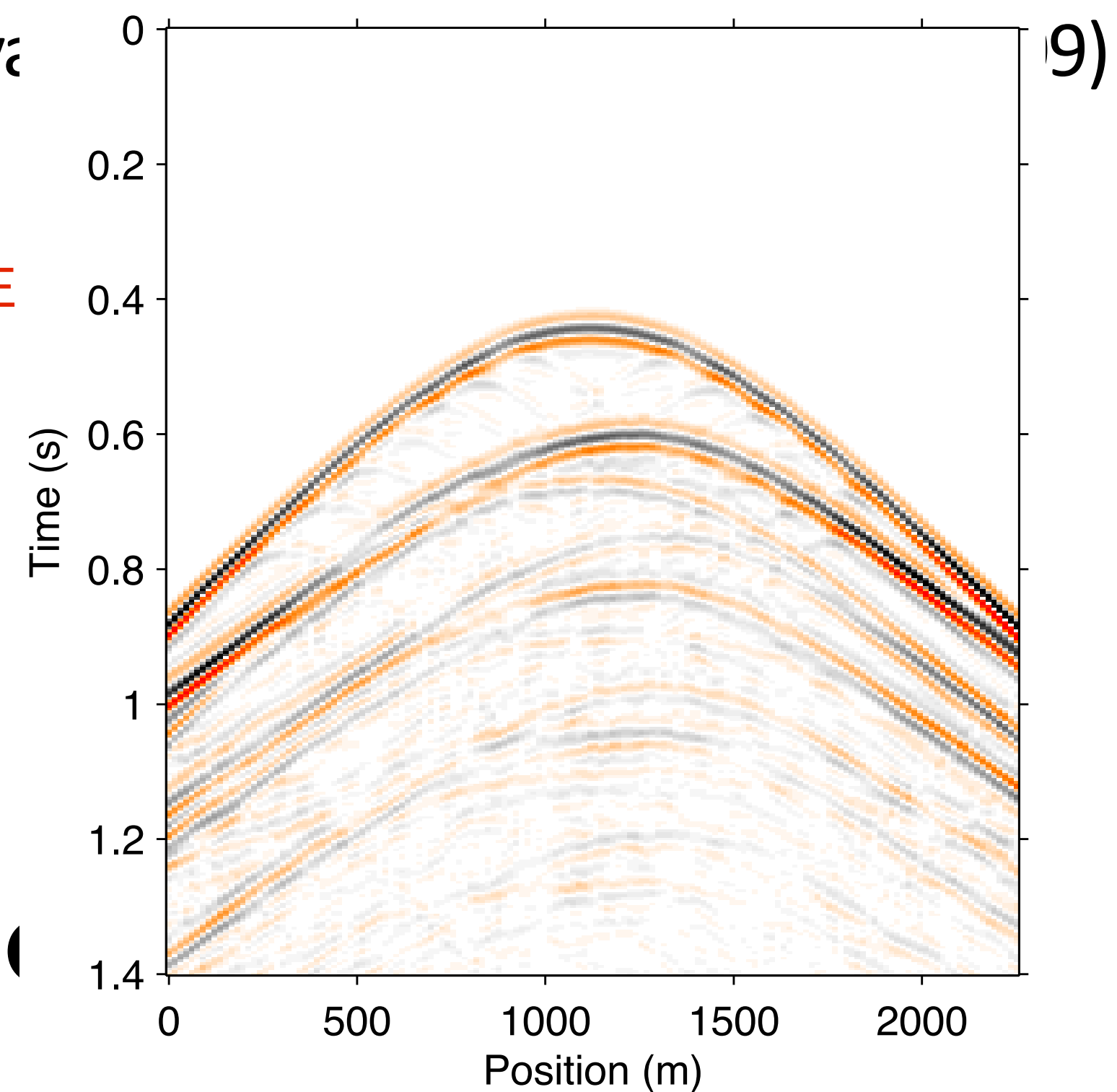
Based on **Estimation of Primaries by Sparse Inversion** (van

recorded data      predicted data from SRME

$$\mathbf{P} = \mathbf{Q}\mathbf{G} - \mathbf{G}\mathbf{P}$$

**Inversion objective:**

$$f(\mathbf{G}, \mathbf{Q}) = \frac{1}{2} \|\mathbf{P} - (\mathbf{Q}\mathbf{G} - \mathbf{G}\mathbf{P})\|$$



9)

# From SRME to Robust EPSI

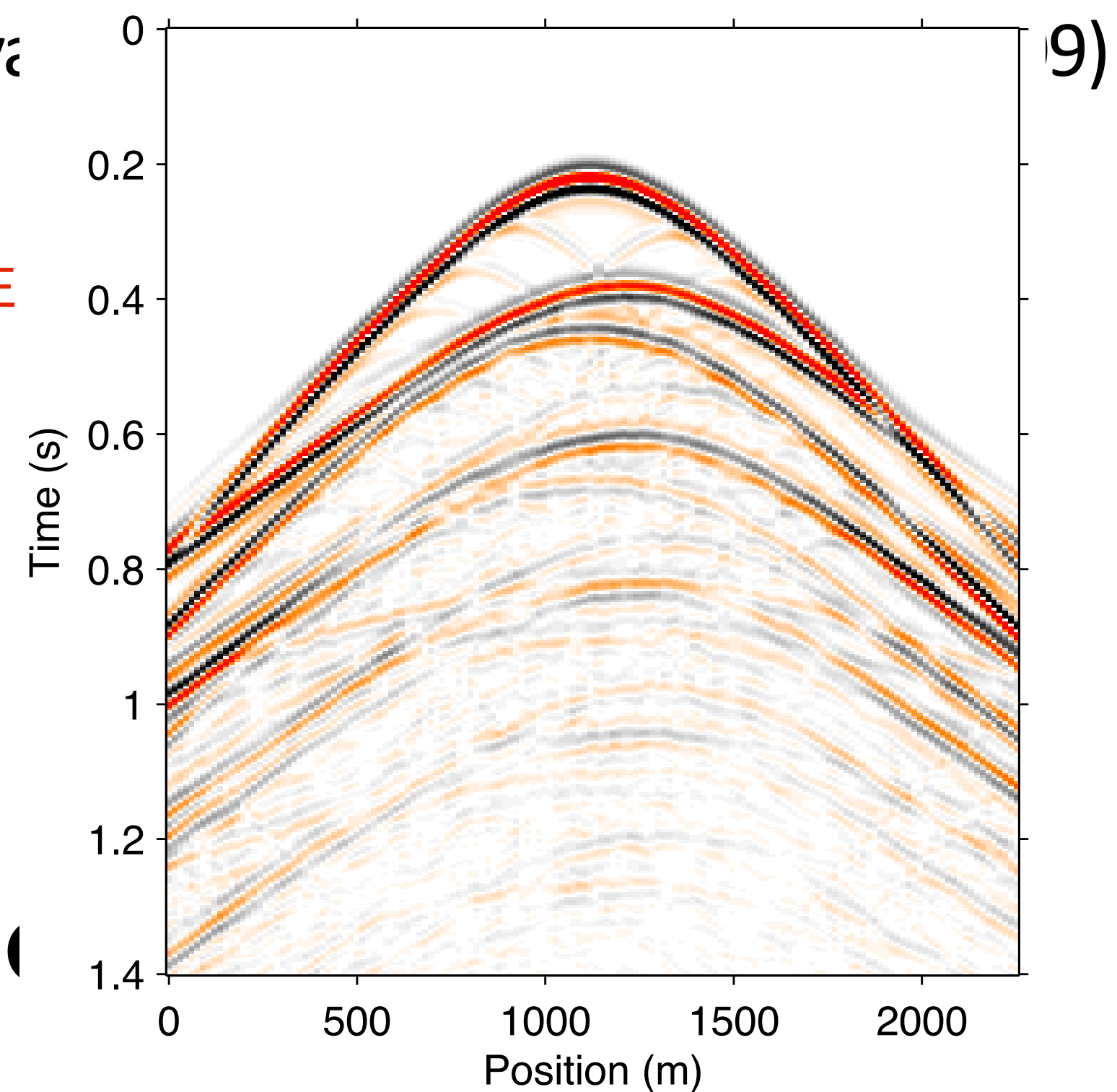
Based on **Estimation of Primaries by Sparse Inversion** (van

recorded data      predicted data from SRME

$$\mathbf{P} = \mathbf{Q}\mathbf{G} - \mathbf{G}\mathbf{P}$$

**Inversion objective:**

$$f(\mathbf{G}, \mathbf{Q}) = \frac{1}{2} \|\mathbf{P} - (\mathbf{Q}\mathbf{G} - \mathbf{G}\mathbf{P})\|$$





## From SRME to Robust EPSI

Based on **Estimation of Primaries by Sparse Inversion** (van Groenestijn and Verschuur, 2009)

recorded data      predicted data from SRME

$$\mathbf{P} = \mathbf{Q}\mathbf{G} - \mathbf{G}\mathbf{P}$$

**Inversion objective:**

$$f(\mathbf{G}, \mathbf{Q}) = \frac{1}{2} \|\mathbf{P} - (\mathbf{Q}\mathbf{G} - \mathbf{G}\mathbf{P})\|_2^2$$

## From SRME to Robust EPSI

Two ways to obtain the final primary wavefield

“Direct” Primary      “Conservative” Primary

$$\mathbf{QG} = \mathbf{P} + \mathbf{GP}$$

**Inversion objective:**

$$f(\mathbf{G}, \mathbf{Q}) = \frac{1}{2} \|\mathbf{P} - (\mathbf{QG} - \mathbf{GP})\|_2^2$$

# From SRME to Robust EPSI

**In time domain** (lower-case: whole dataset in time domain)

recorded data      predicted data from SRME

$$\mathbf{p} = \mathcal{M}(\mathbf{g}, \mathbf{q})$$

$$\mathcal{M}(\mathbf{g}, \mathbf{q}) := \mathcal{F}_t^\dagger \text{BlockDiag}_{\omega_1 \dots \omega_{n_f}} [(q(\omega)\mathbf{I} - \mathbf{P})^\dagger \otimes \mathbf{I}] \mathcal{F}_t \mathbf{g}$$

**Inversion objective:**

$$f(\mathbf{g}, \mathbf{q}) = \frac{1}{2} \|\mathbf{p} - \mathcal{M}(\mathbf{g}, \mathbf{q})\|_2^2$$

# Solving the EPSI problem

## Linearizations

$$\mathbf{p} = \mathcal{M}(\mathbf{g}, \mathbf{q})$$

$$\mathbf{M}_{\tilde{\mathbf{q}}} = \left( \frac{\partial \mathcal{M}}{\partial \mathbf{g}} \right)_{\tilde{\mathbf{q}}}$$

$$\mathbf{M}_{\tilde{\mathbf{g}}} = \left( \frac{\partial \mathcal{M}}{\partial \mathbf{q}} \right)_{\tilde{\mathbf{g}}}$$

In fact it is bilinear:

$$\mathbf{M}_{\tilde{\mathbf{q}}}\mathbf{g} = \mathcal{M}(\mathbf{g}, \tilde{\mathbf{q}}) \quad \mathbf{M}_{\tilde{\mathbf{g}}}\mathbf{q} = \mathcal{M}(\mathbf{q}, \tilde{\mathbf{g}})$$

# Solving the EPSI problem

## Linearizations

$$\mathbf{p} = \mathcal{M}(\mathbf{g}, \mathbf{q})$$

$$\mathbf{M}_{\tilde{q}} = \left( \frac{\partial \mathcal{M}}{\partial \mathbf{g}} \right)_{\tilde{q}}$$

$$\mathbf{M}_{\tilde{g}} = \left( \frac{\partial \mathcal{M}}{\partial \mathbf{q}} \right)_{\tilde{g}}$$

## Associated objectives:

$$f_{\tilde{q}}(\mathbf{g}) = \frac{1}{2} \|\mathbf{p} - \mathbf{M}_{\tilde{q}} \mathbf{g}\|_2^2$$

$$f_{\tilde{g}}(\mathbf{q}) = \frac{1}{2} \|\mathbf{p} - \mathbf{M}_{\tilde{g}} \mathbf{q}\|_2^2$$



# Solving the EPSI problem

**Do:**

$$\mathbf{g}_{k+1} = \mathbf{g}_k + \alpha \mathcal{S}(\nabla f_{q_k}(\mathbf{g}_k))$$

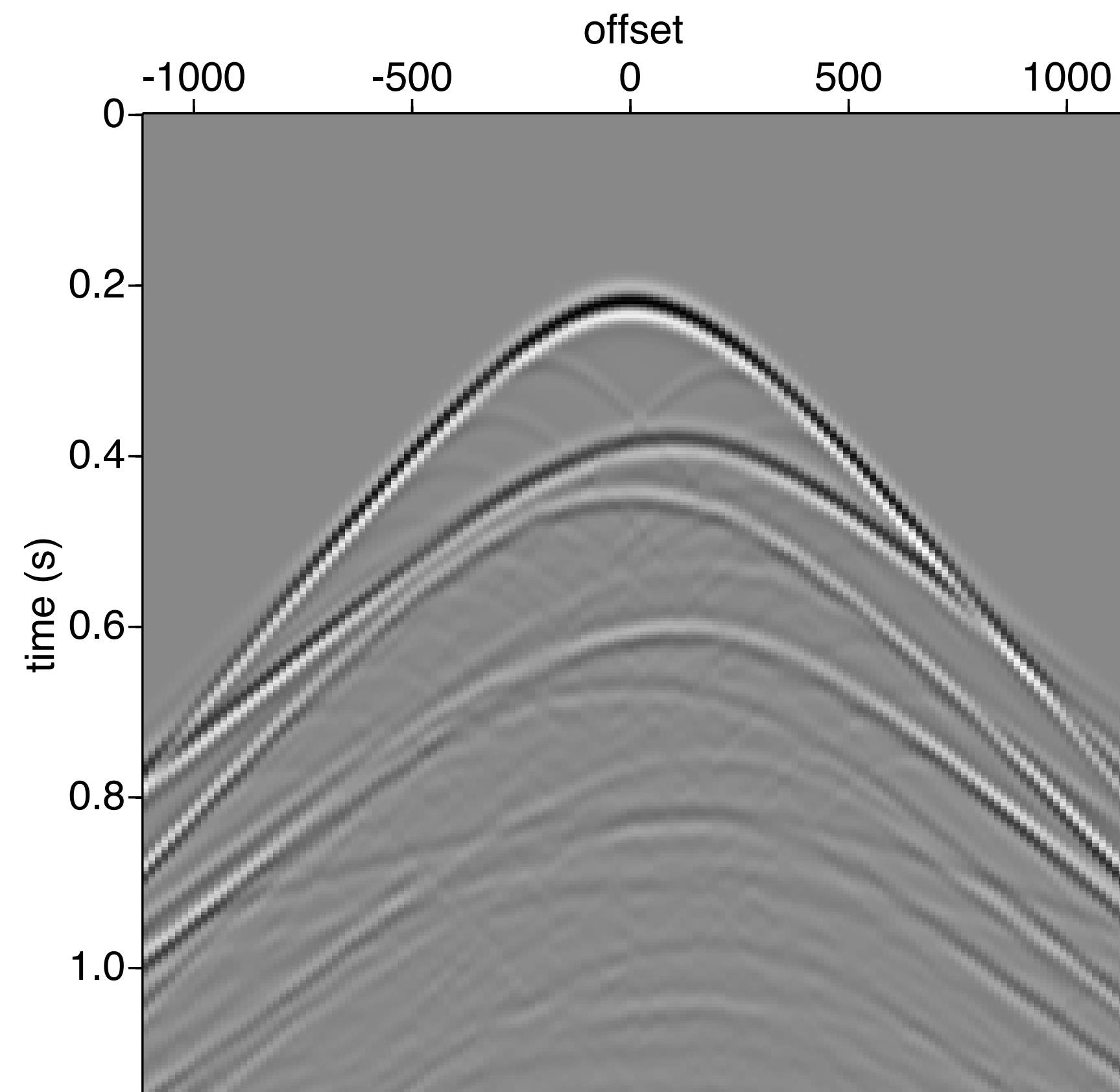
$$\mathbf{q}_{k+1} = \mathbf{q}_k + \beta \nabla f_{g_{k+1}}(\mathbf{q}_k)$$

**Gradient sparsity**

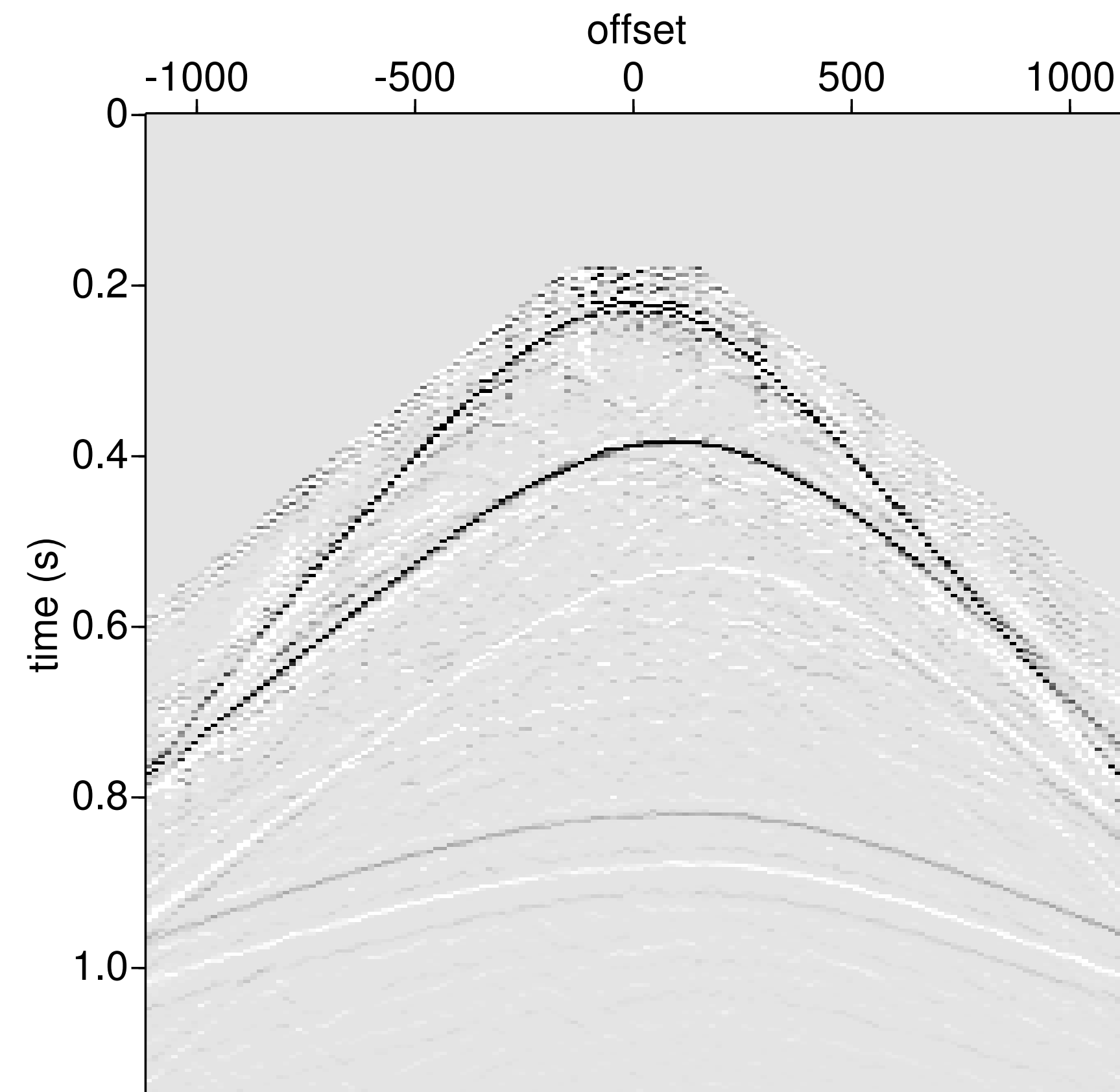
$\mathcal{S}$  : pick largest  $\rho$  elements per trace



# Solving the EPSI problem



Data



EPSI IR

# Robust EPSI

## L1-minimization approach to the EPSI problem

[Lin and Herrmann, 2013 *Geophysics*]

**While**  $\|\mathbf{p} - \mathcal{M}(\mathbf{g}_k, \mathbf{q}_k)\|_2 > \sigma$

determine new  $\tau_k$  from the Pareto curve

$$\mathbf{g}_{k+1} = \arg \min_{\mathbf{g}} \|\mathbf{p} - \mathbf{M}_{q_k} \mathbf{g}\|_2 \text{ s.t. } \|\mathbf{g}\|_1 \leq \tau_k$$

(Solve with SPG part of SPGL1 until Pareto curve reached)

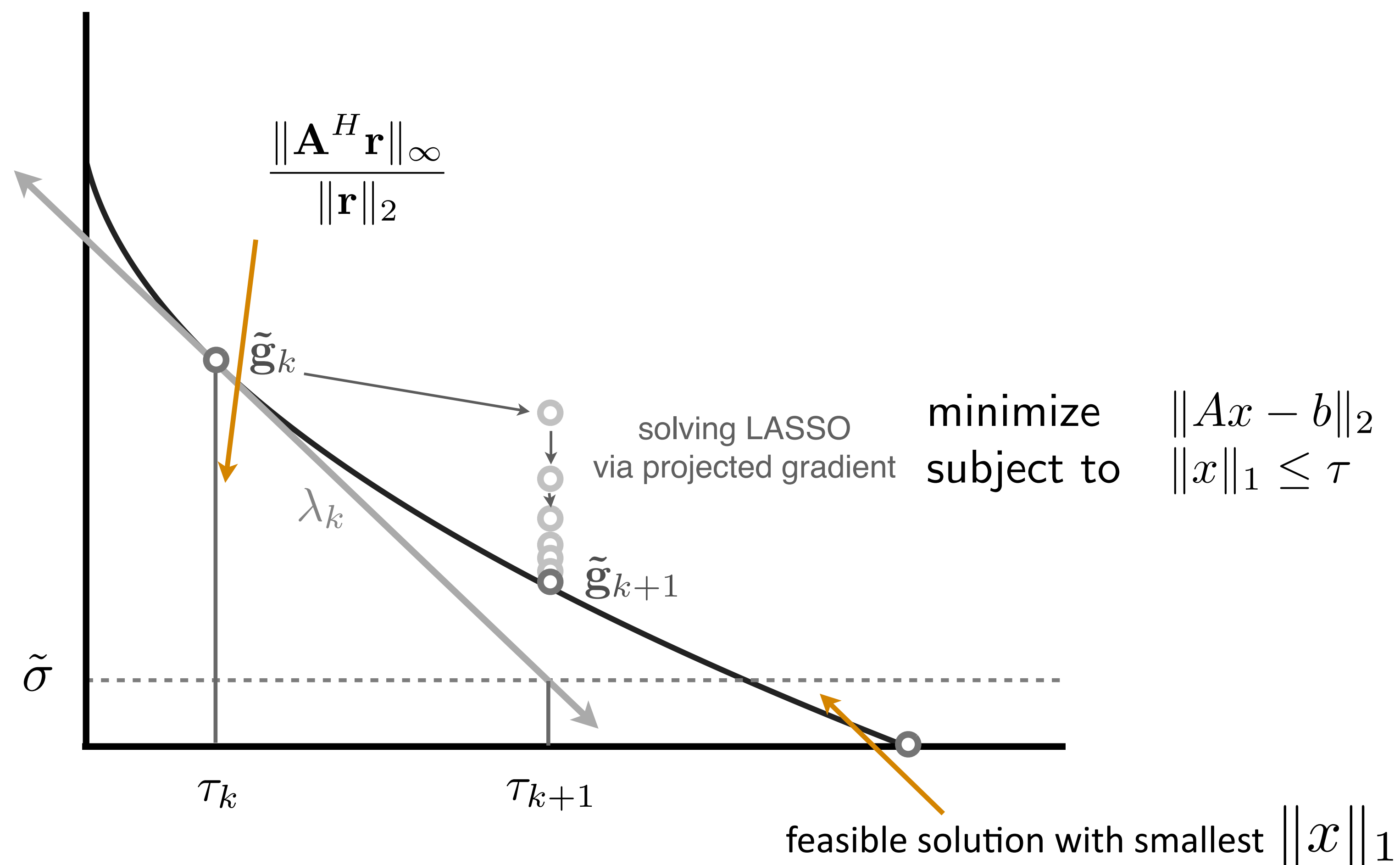
$$\mathbf{q}_{k+1} = \arg \min_{\mathbf{q}} \|\mathbf{p} - \mathbf{M}_{g_{k+1}} \mathbf{q}\|_2$$

(wavelet matching, solve with LSQR)

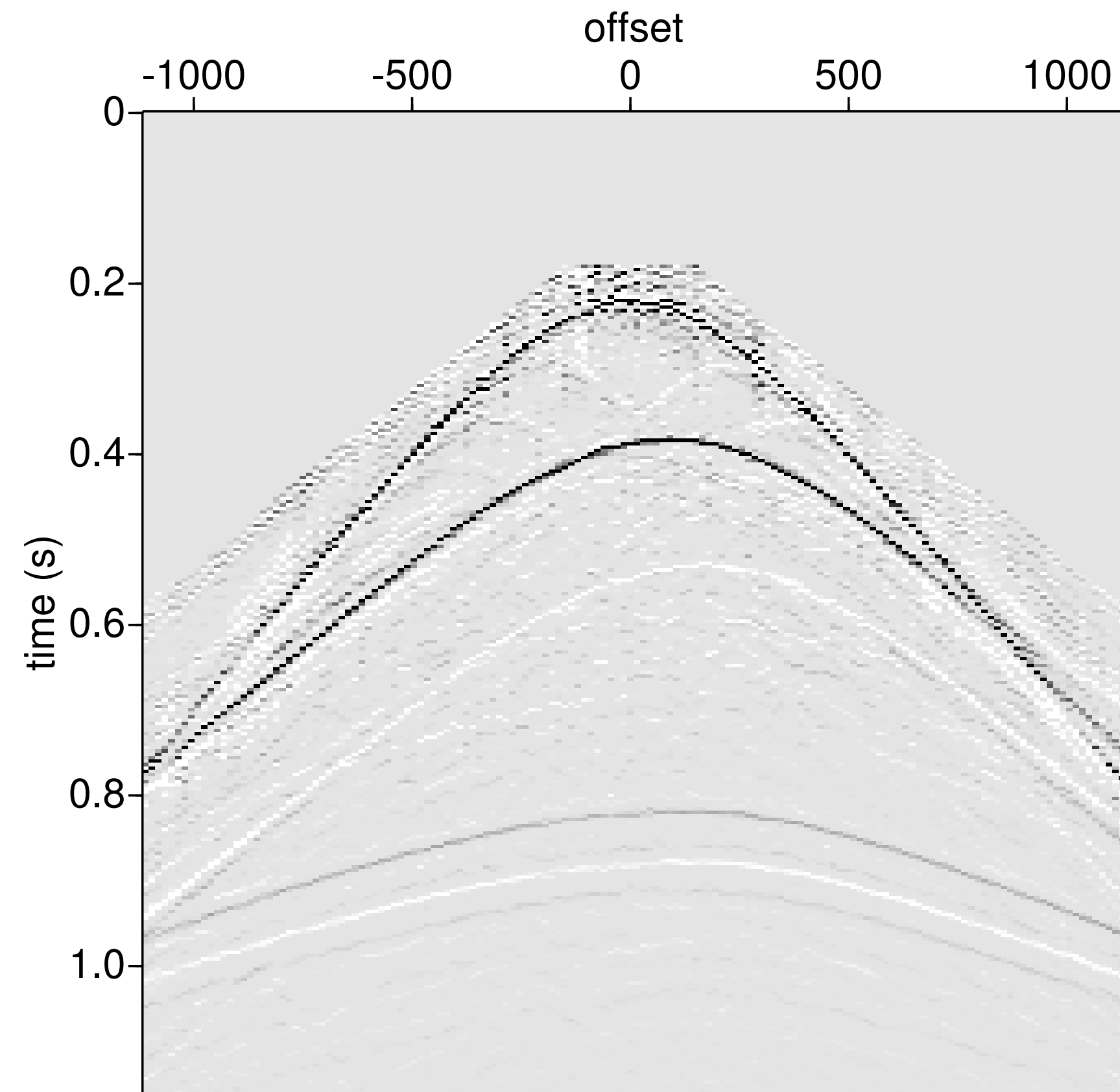
# Choosing Tau from the Pareto curve

$$\begin{aligned} &\text{minimize} && \|x\|_1 \\ &\text{subject to} && \|Ax - b\|_2 \leq \sigma \end{aligned}$$

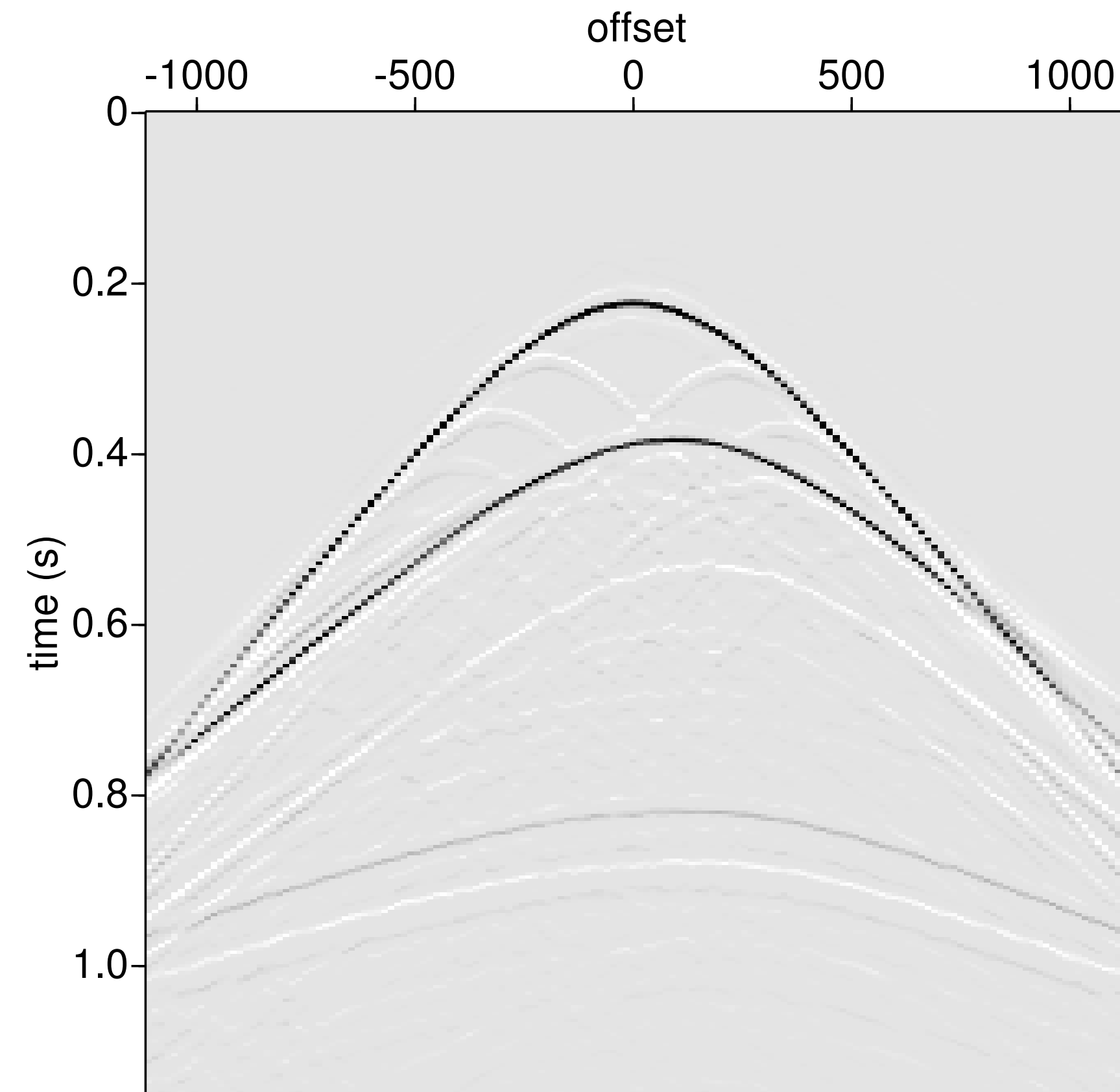
Look at the solution space and the line of optimal solutions (Pareto curve)



# Robust EPSI



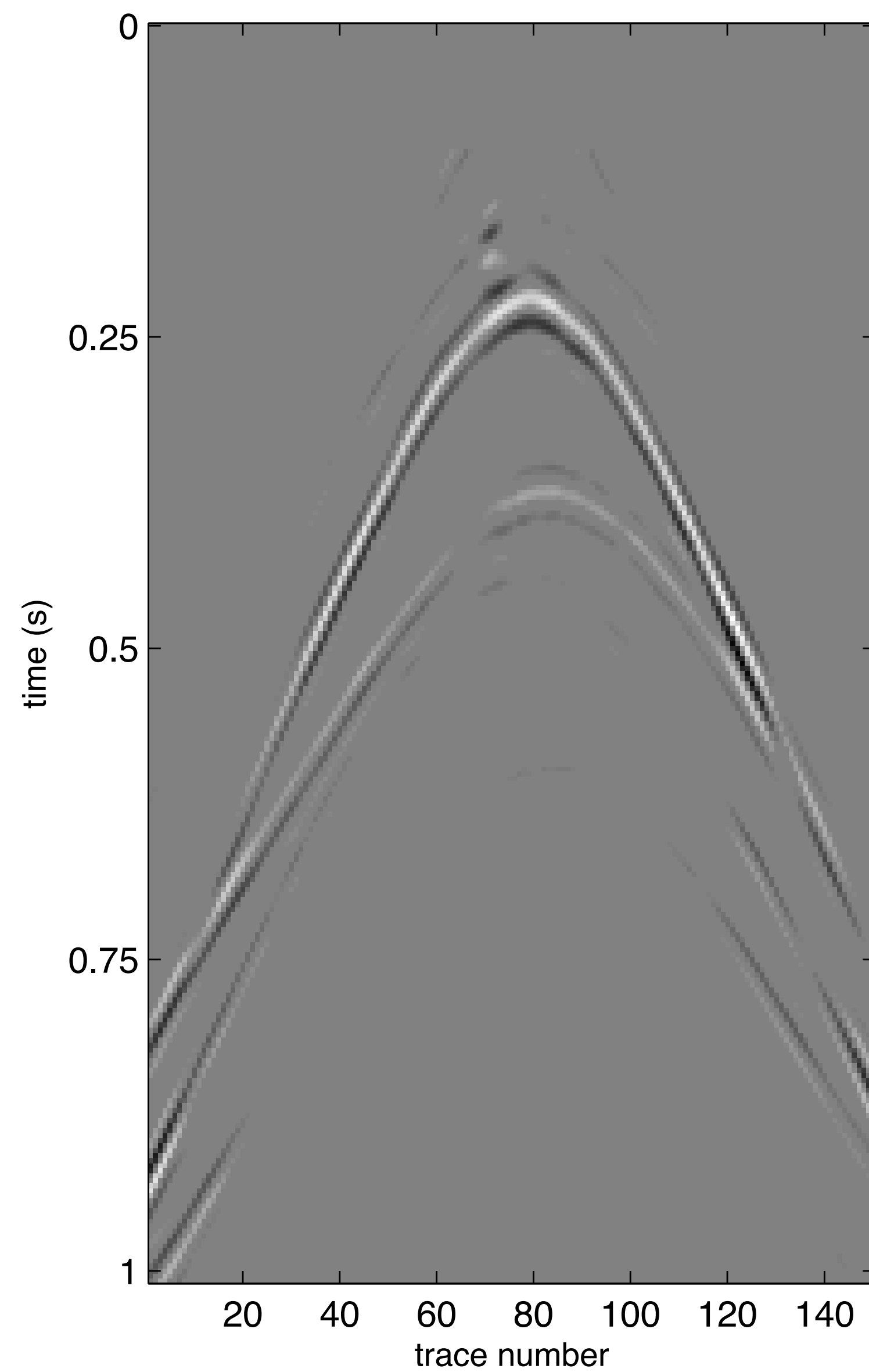
EPSI IR



Robust EPSI IR

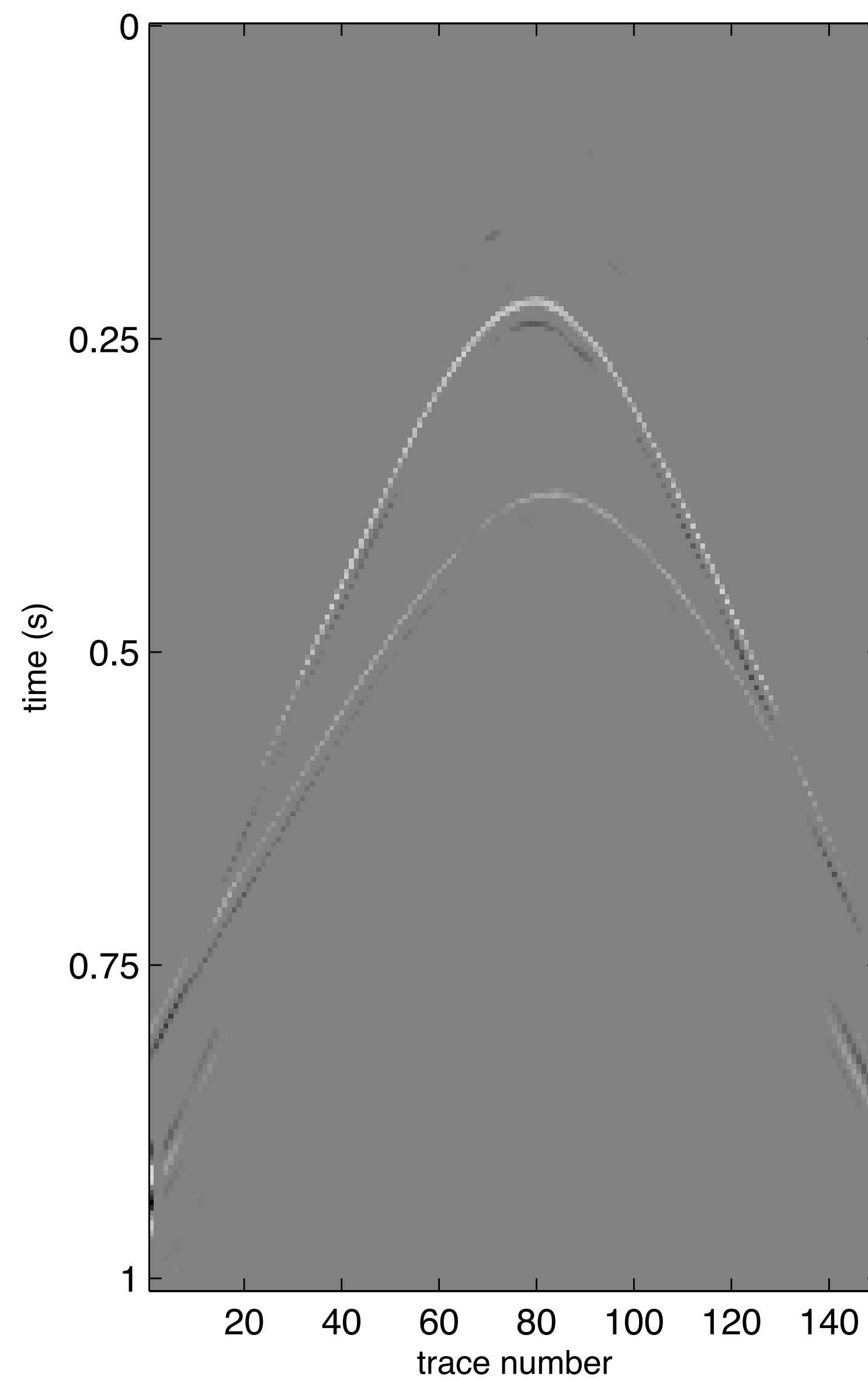


# L1 projection and sparsity



**Solution at beginning of SPG**

# L1 projection and sparsity

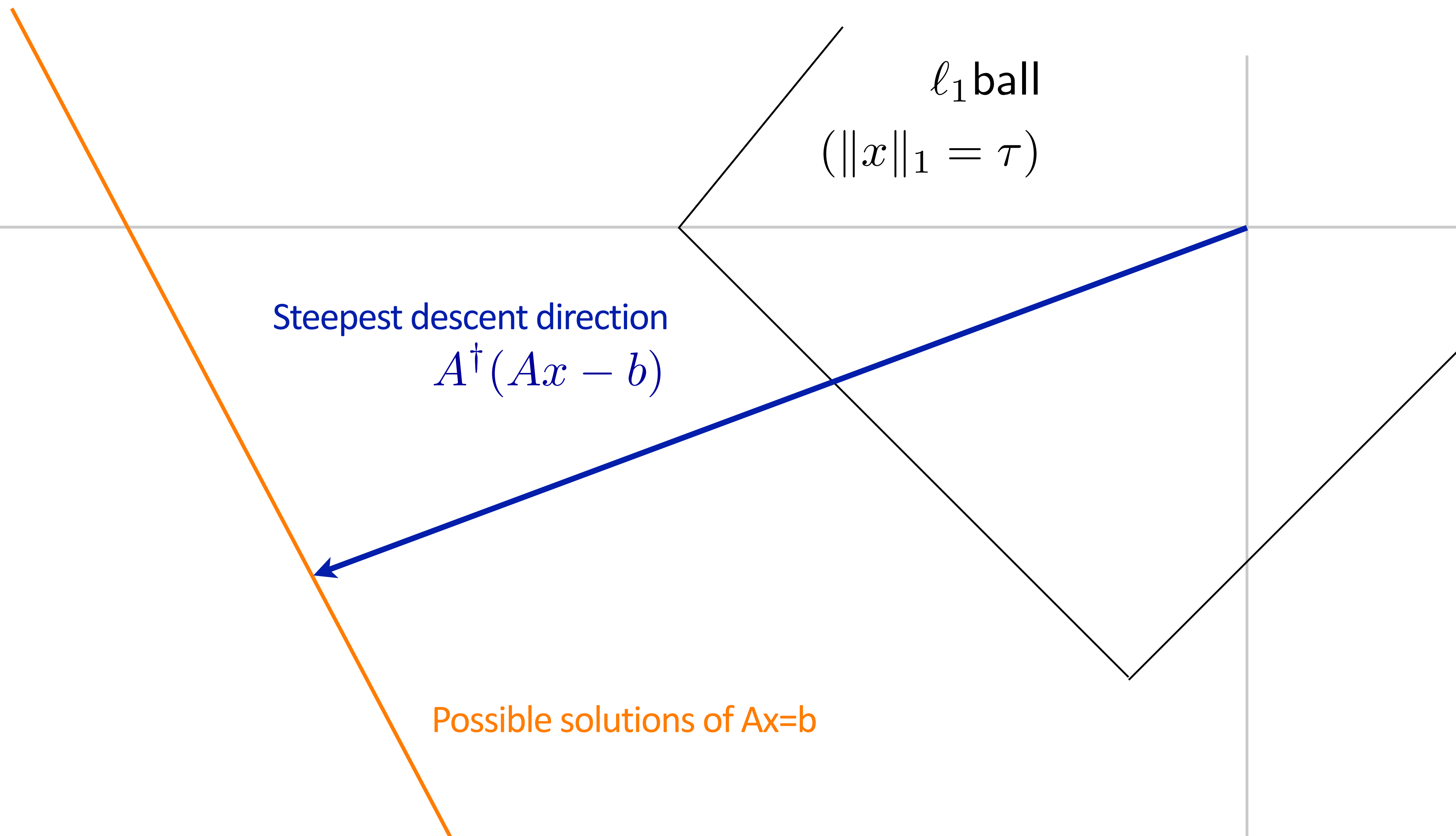


**Solution at end of SPG  
(pareto-optimal)**



# L1 projection and sparsity

$$\begin{aligned} &\text{minimize} && \|Ax - b\|_2 \\ &\text{subject to} && \|x\|_1 \leq \tau \end{aligned}$$

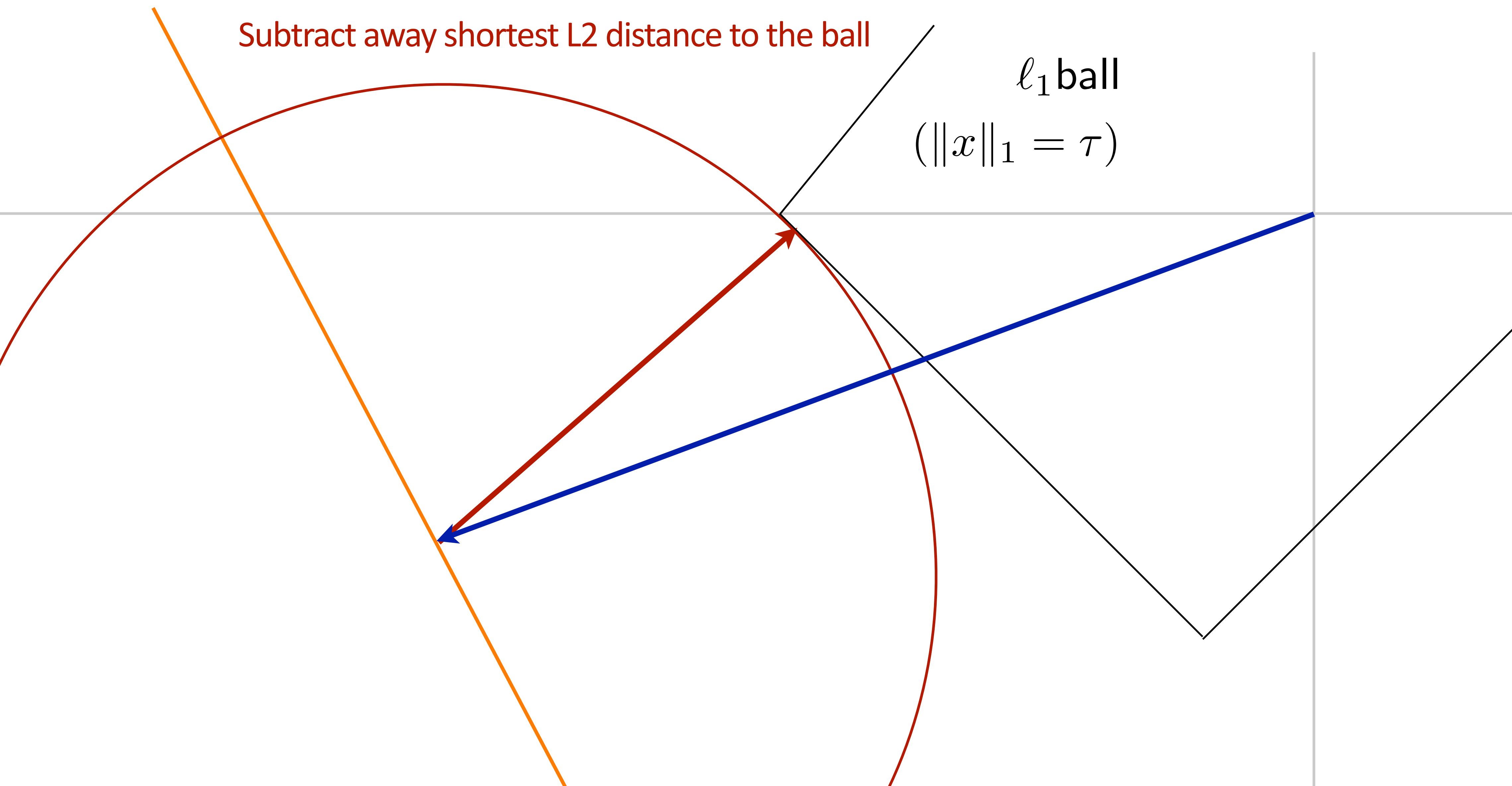


# L1 projection and sparsity

$$\begin{aligned} &\text{minimize} && \|Ax - b\|_2 \\ &\text{subject to} && \|x\|_1 \leq \tau \end{aligned}$$

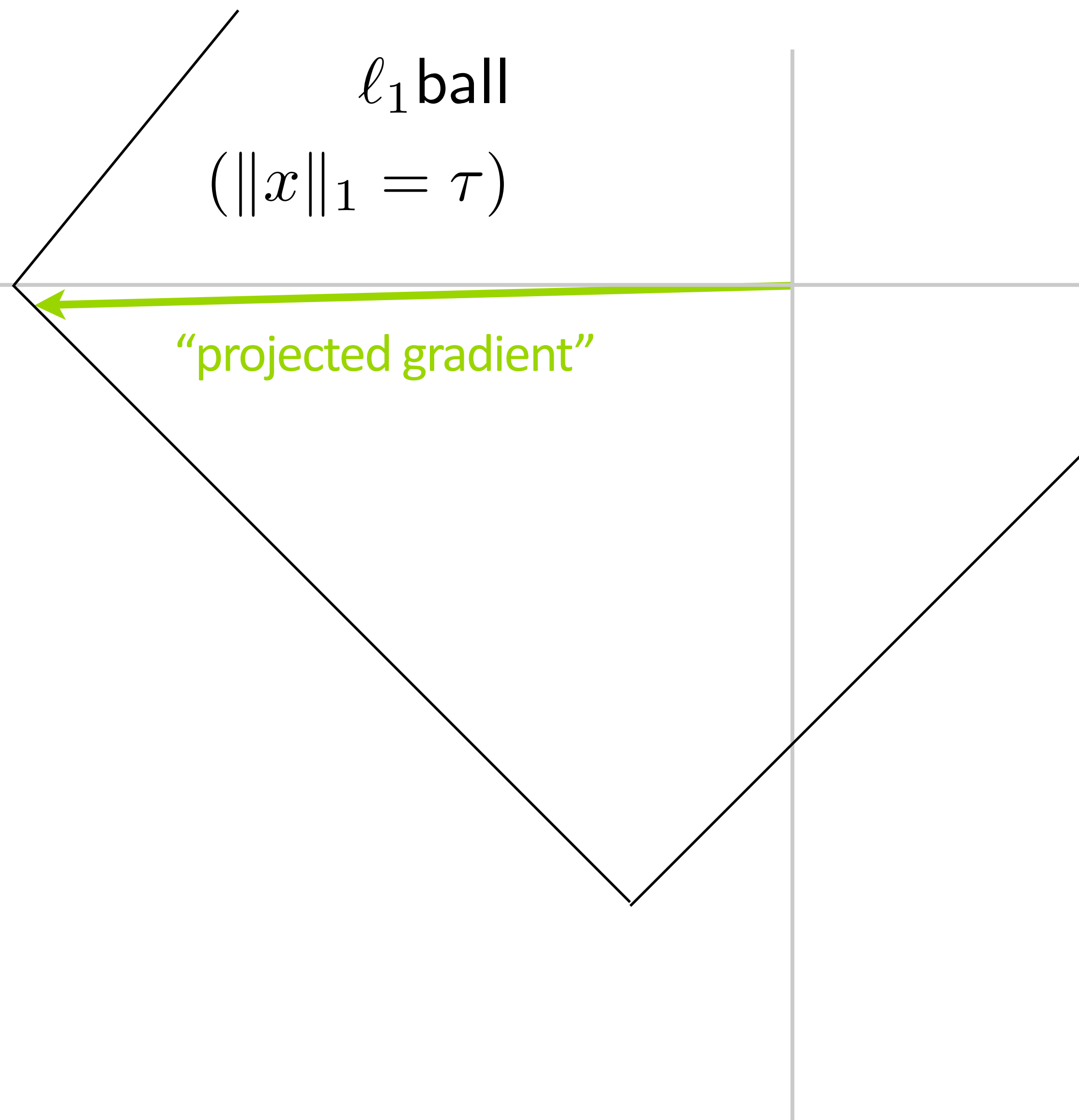
Subtract away shortest L2 distance to the ball

$\ell_1$  ball  
( $\|x\|_1 = \tau$ )



# L1 projection and sparsity

$$\begin{aligned} &\text{minimize} && \|Ax - b\|_2 \\ &\text{subject to} && \|x\|_1 \leq \tau \end{aligned}$$



# L1 projection and sparsity

$$\begin{aligned} &\text{minimize} && \|Ax - b\|_2 \\ &\text{subject to} && \|x\|_1 \leq \tau \end{aligned}$$

$\ell_1$  ball  
( $\|x\|_1 = \tau$ )

(proximal gradient if one-norm is a penalty)





# Robust EPSI summary

## Benefits

- Formulation is more physical than SRME w/adaptive subtraction
- Provide useful estimates of both primary and multiple wavefields
- Mostly hands-free
- Much more accurate estimate of the multiples than single-pass SRMP
- Gives sparse impulse response of primary wavefield

## Challenges

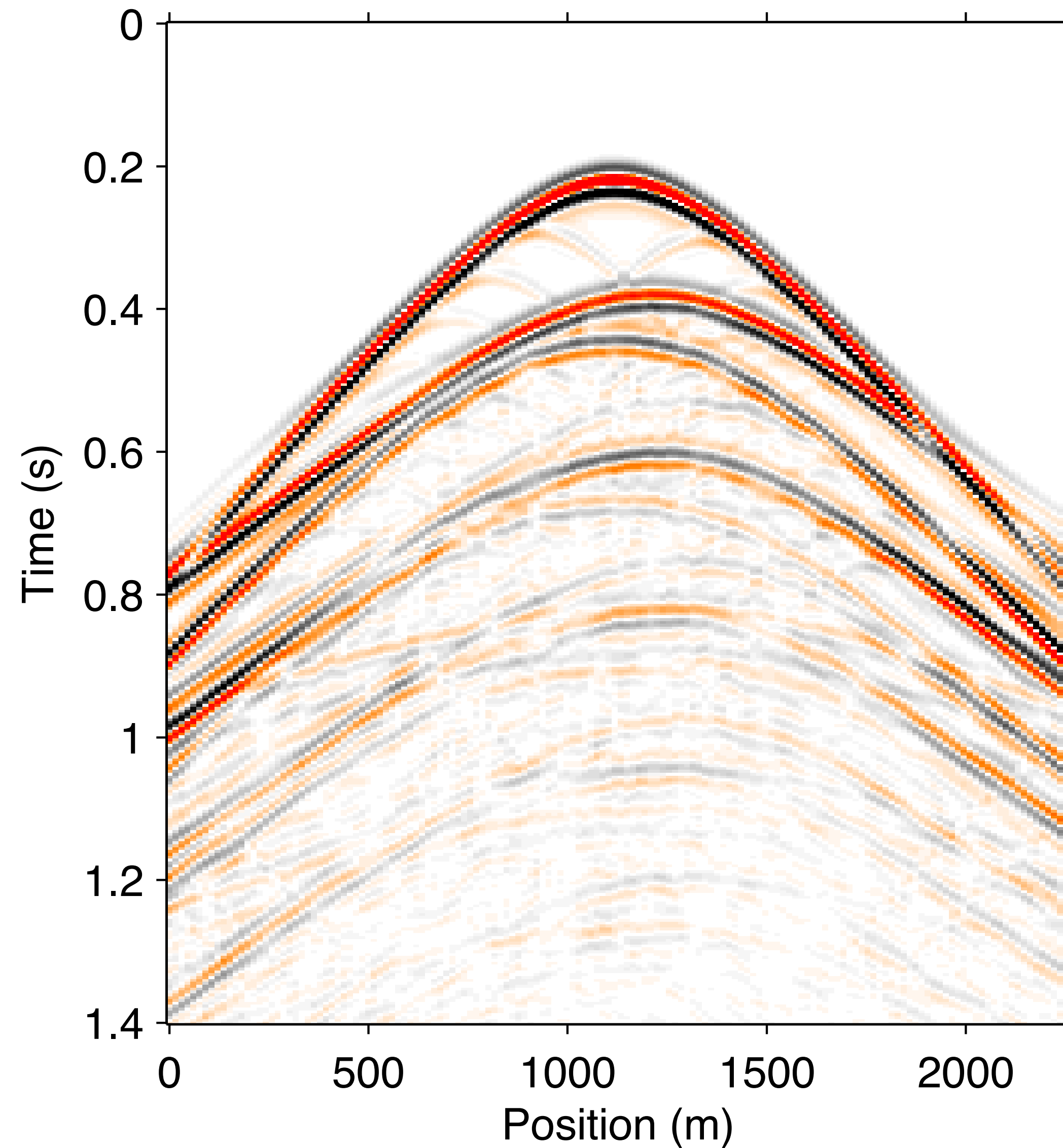
- Cost roughly 100 to 200 SRMP
- Inversion approach means hard to QC until too late



# **Bootstrapping**

(continuation strategy for spatial sampling)

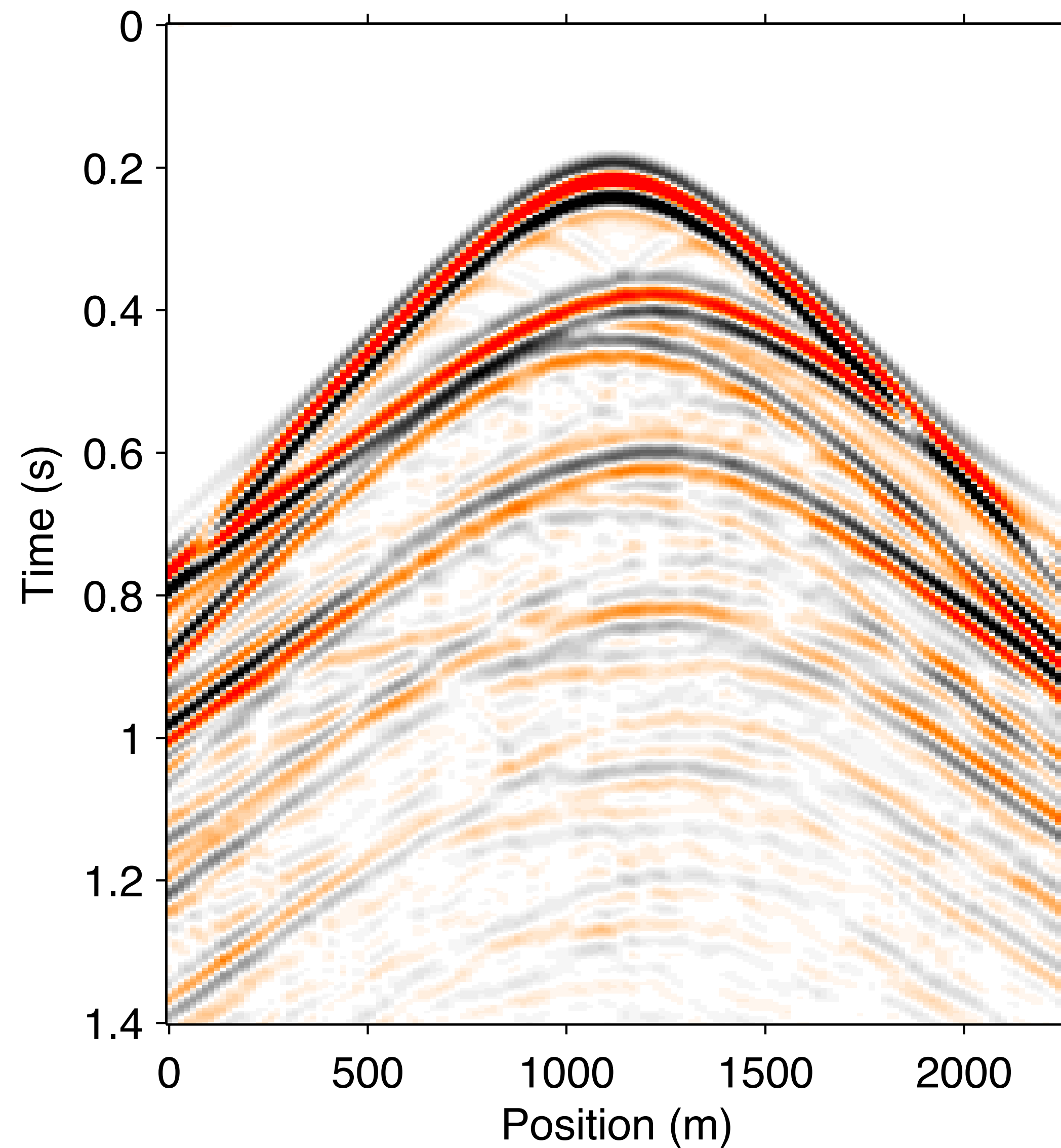
# Motivation: $G$ unchanged by global filters



## Data

modeled with Ricker 30Hz

# Motivation: $G$ unchanged by global filters



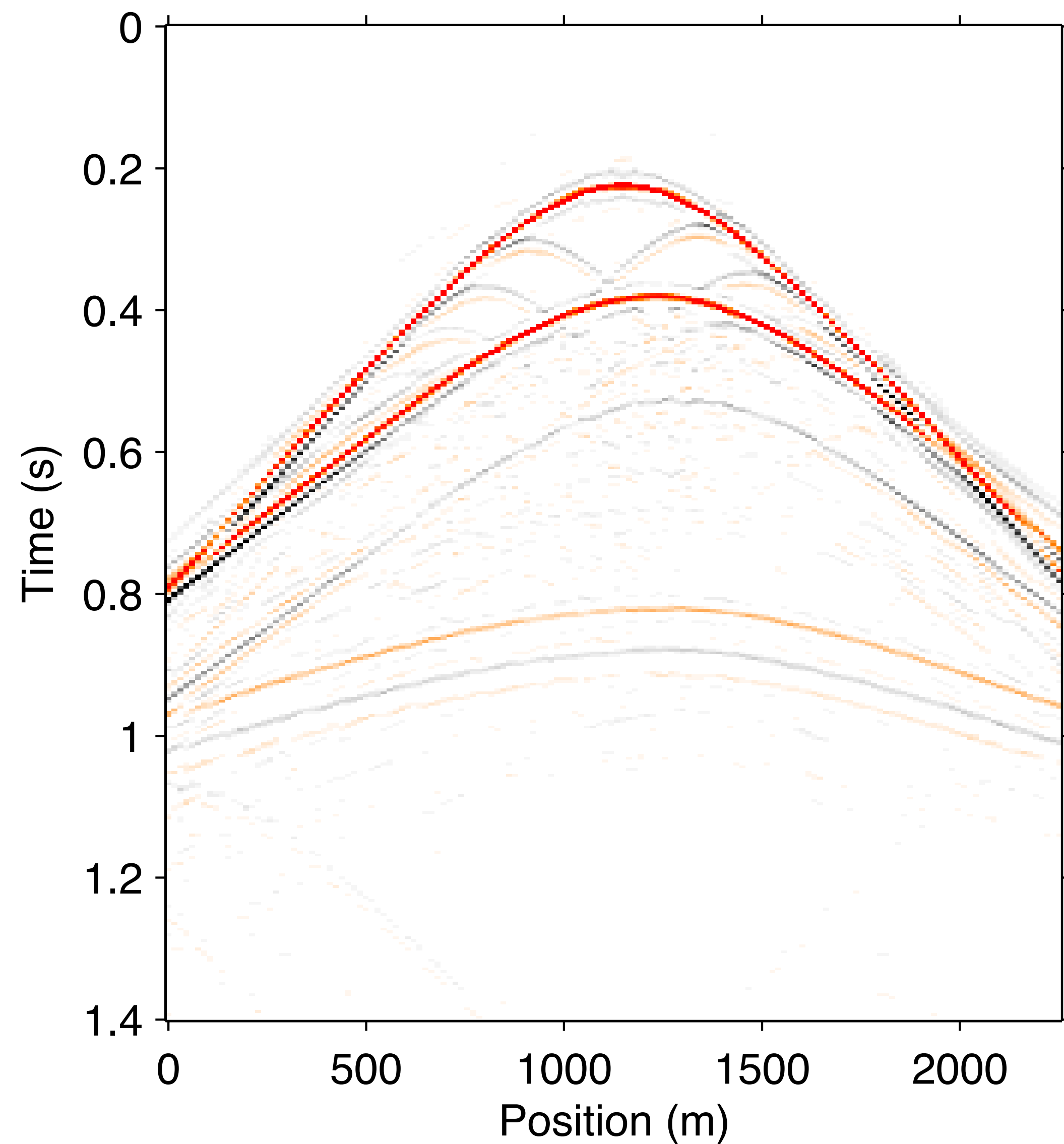
## Lowpassed Data

modeled with Ricker 30Hz

lowpass at 40Hz

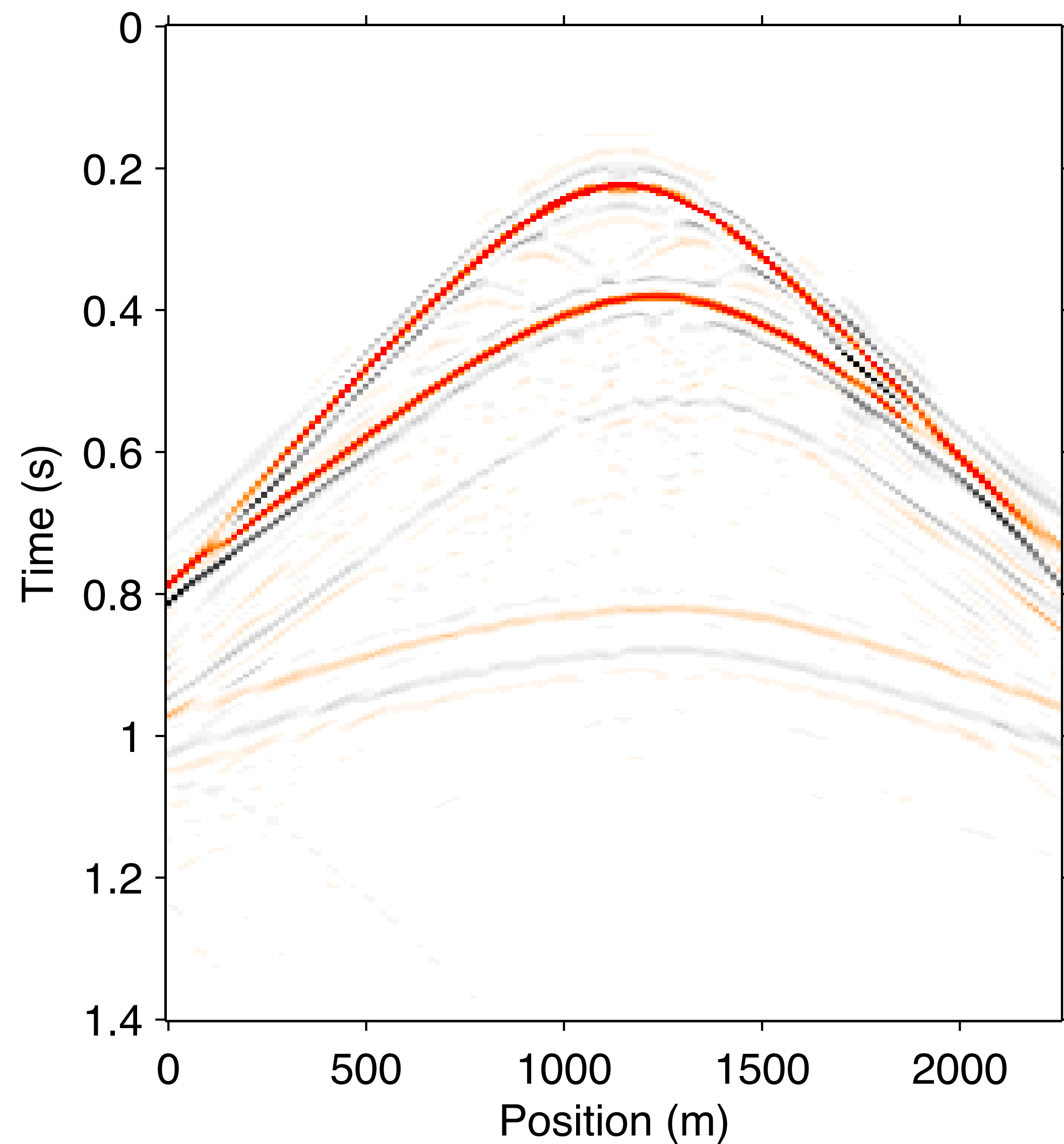
(zero-phase cosine window)

# Motivation: $G$ unchanged by global filters



**Reference REPSI primary IR**  
from original data

# Motivation: $G$ unchanged by global filters

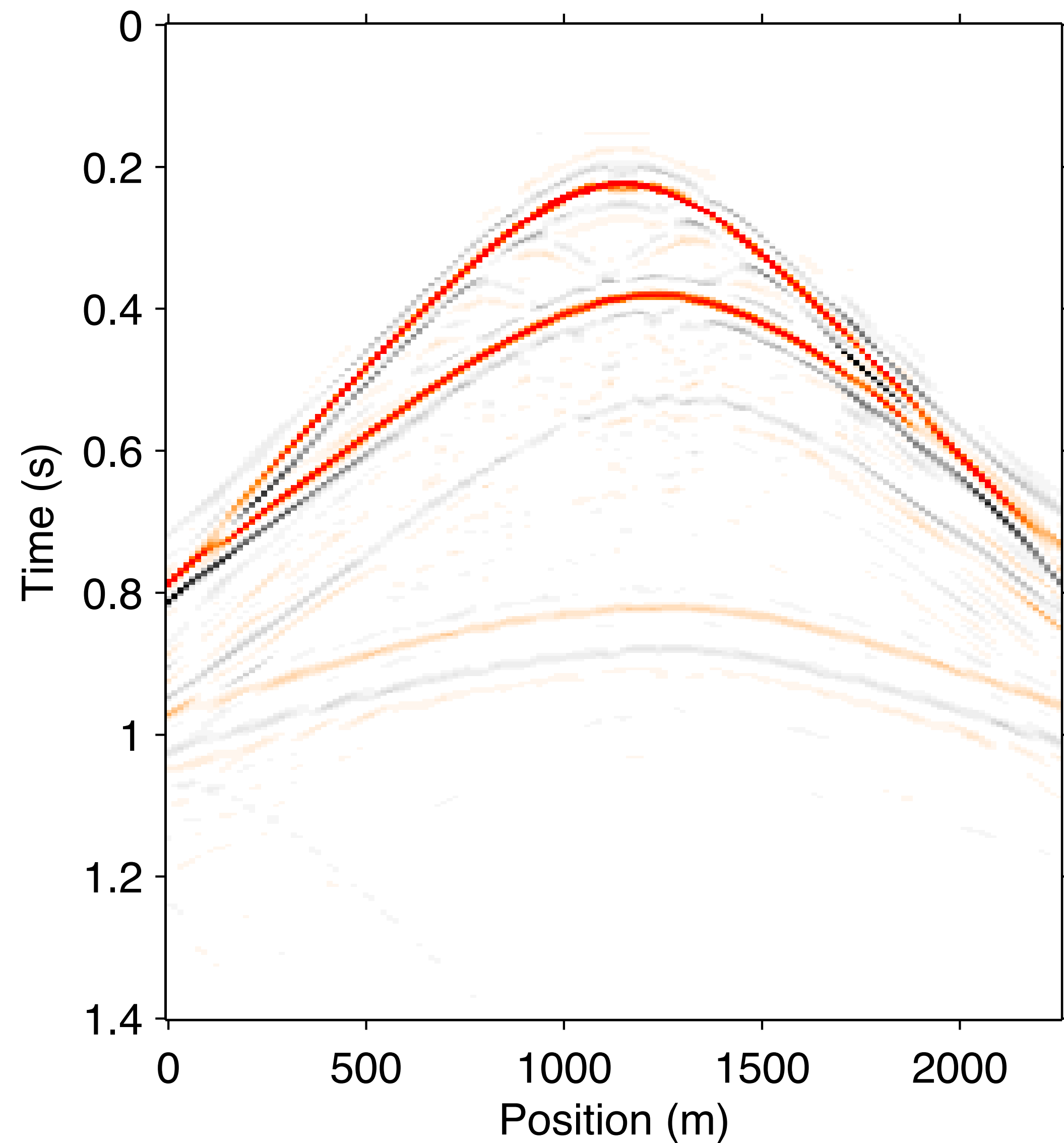


**REPSI primary IR**

from low-passed data @ 40Hz



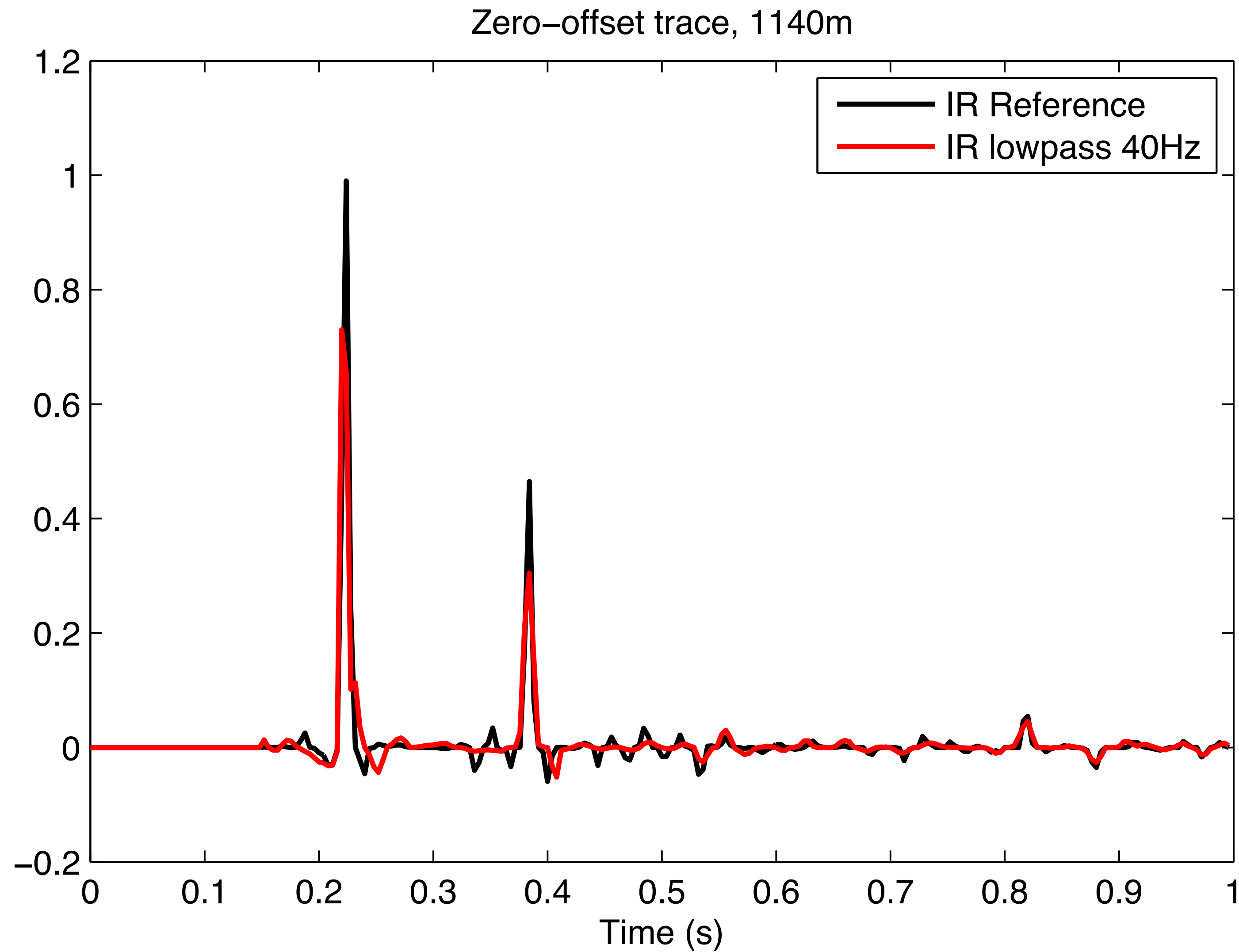
# Motivation: **G** tolerant to global filters



**REPSI primary IR**

from low-passed data @ 40Hz

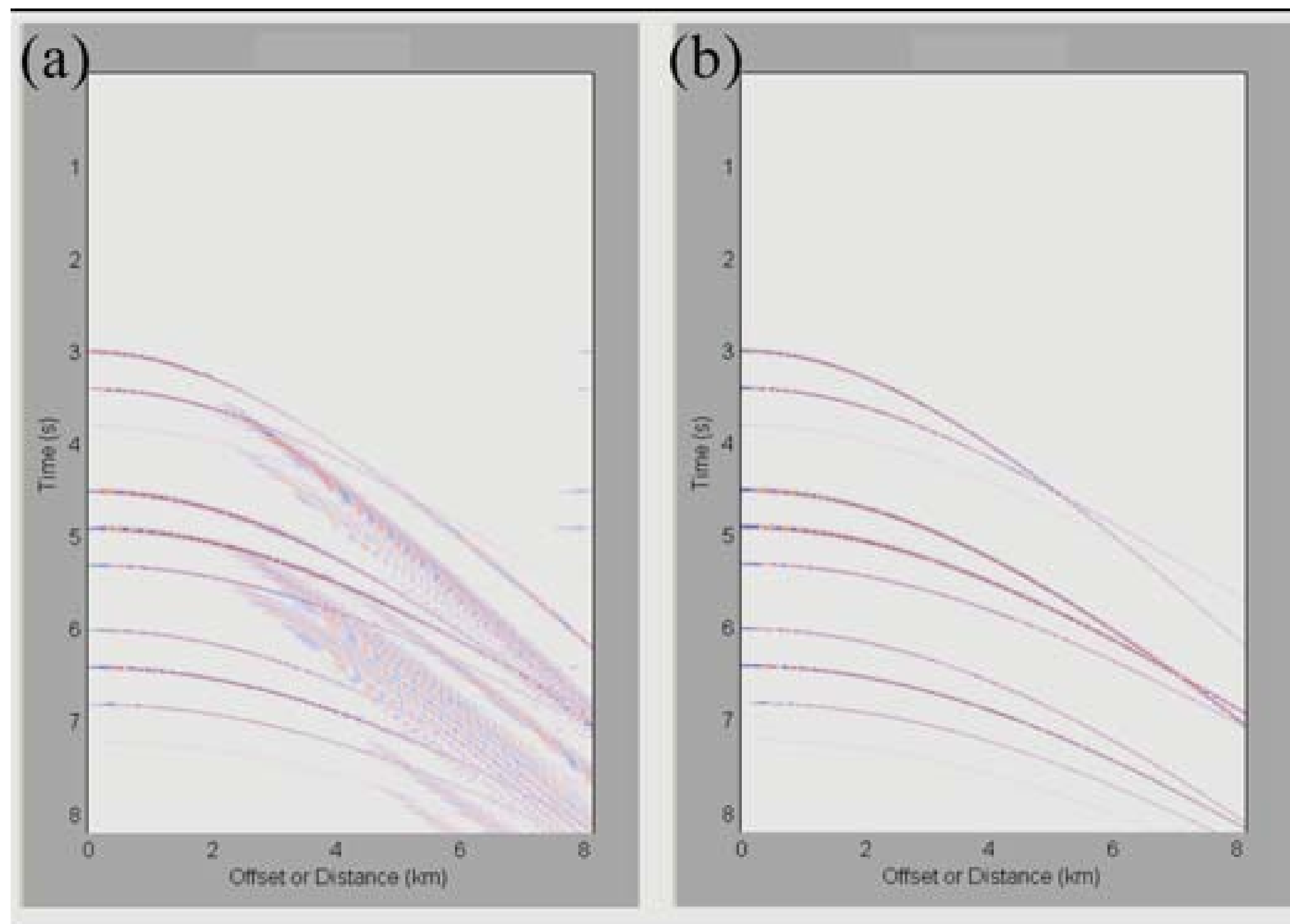
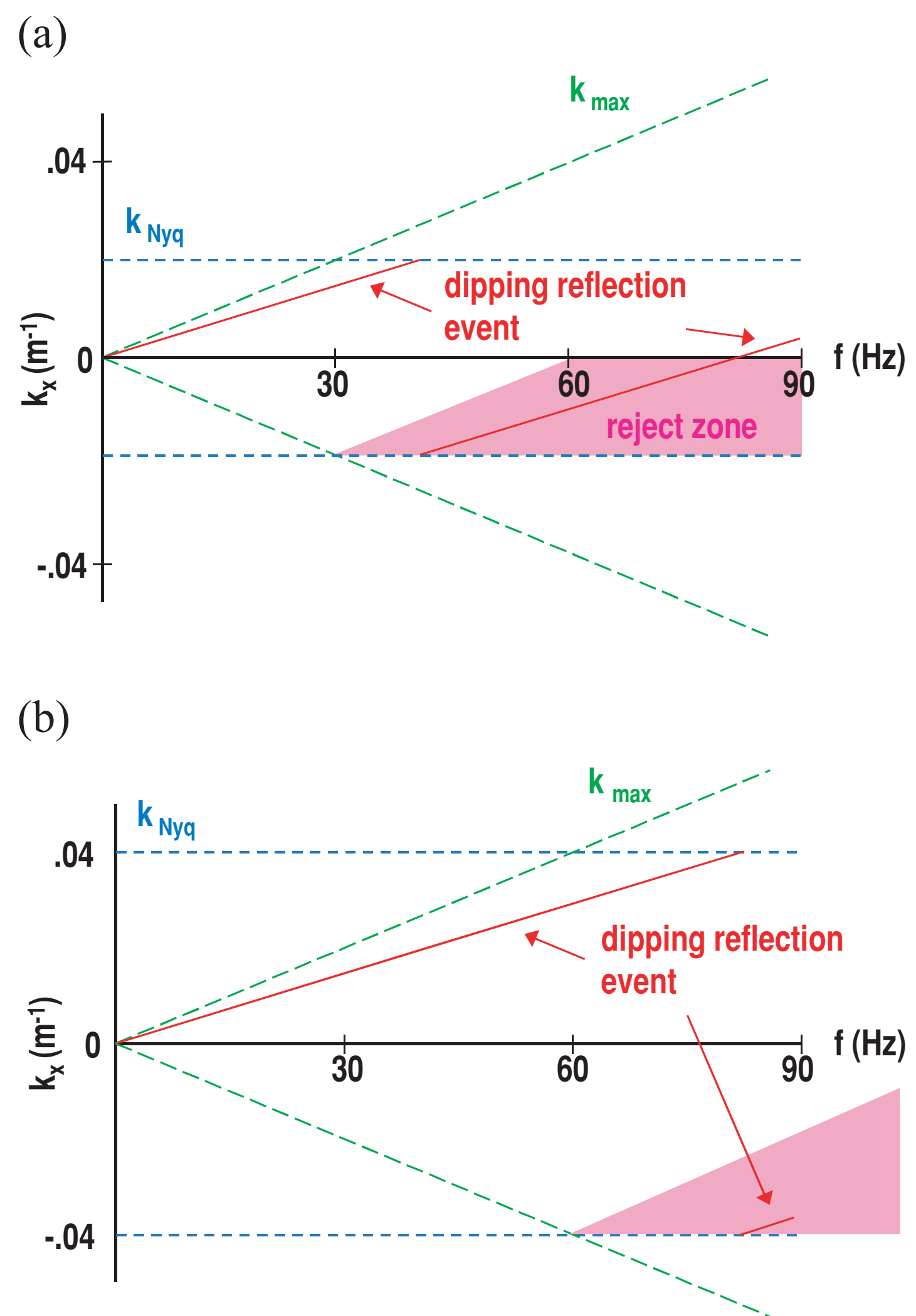
# Motivation: G tolerant to global filters



# Sampling issue in multiple prediction = alias

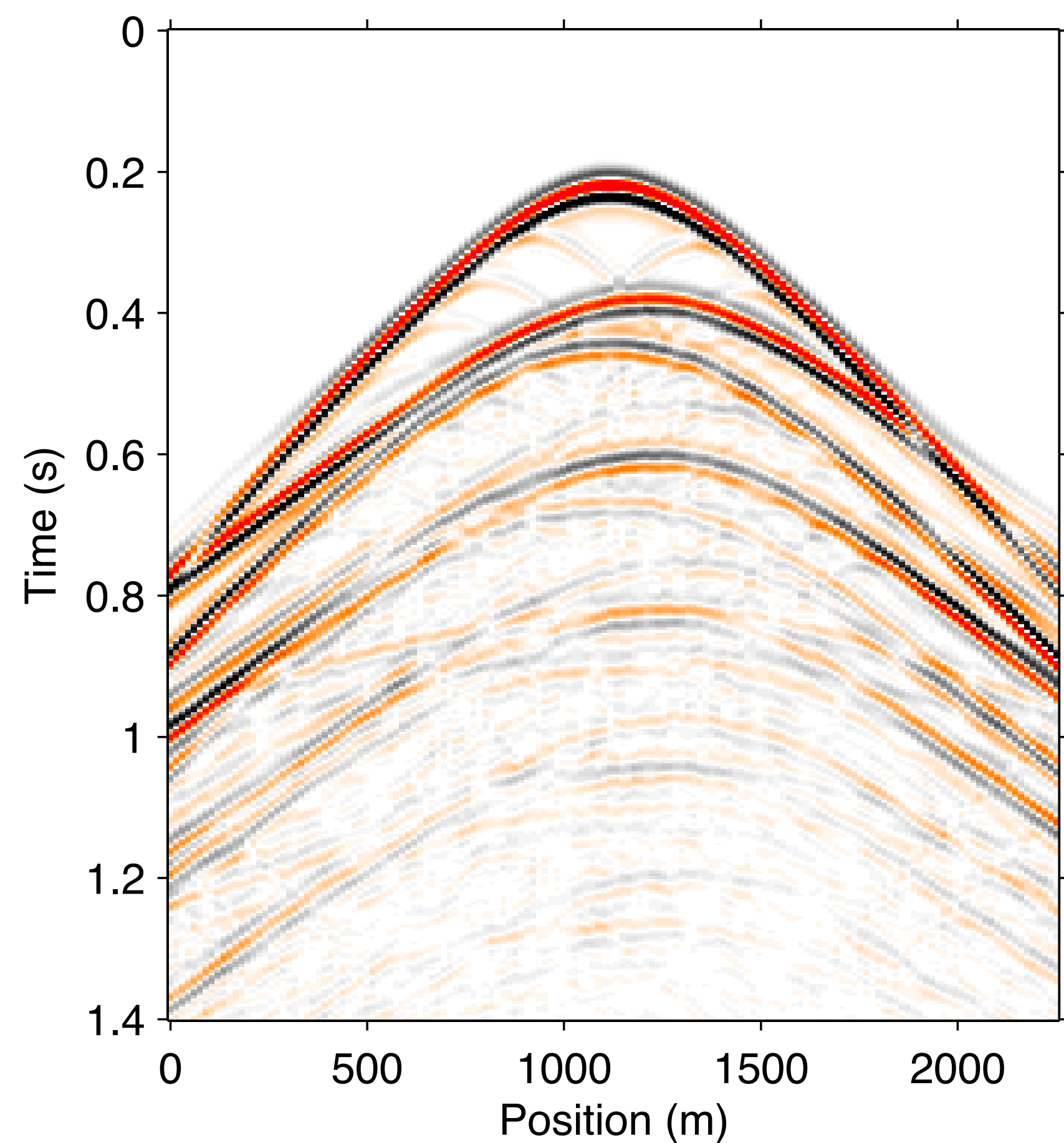
*“The impact of field-survey characteristics on surface-related multiple attenuation”*

Dragoset, Moore, Kostov 2006

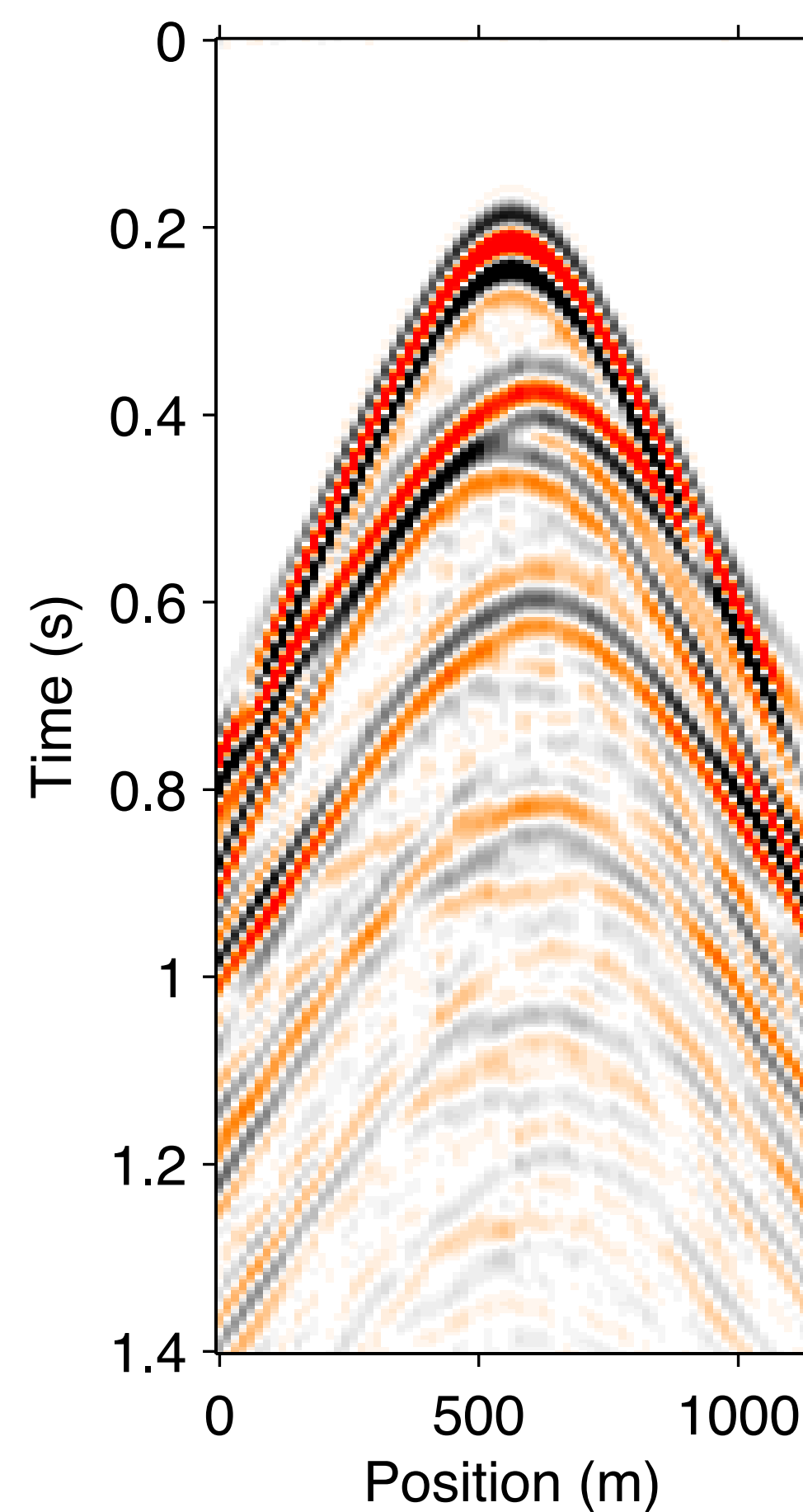


# Idea 1: Solve REPSI on decimated dataset (low-pass in temporal frequency)

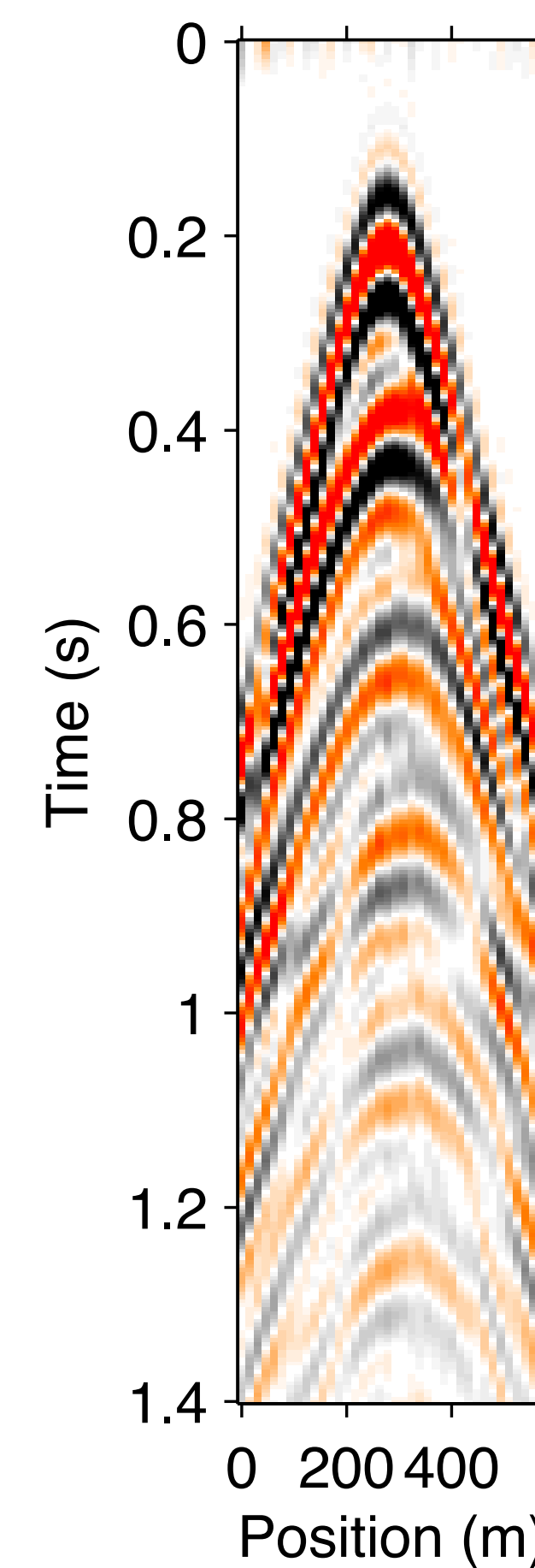
Original (dx = 15m)



2x decimated  
lowpass 30Hz



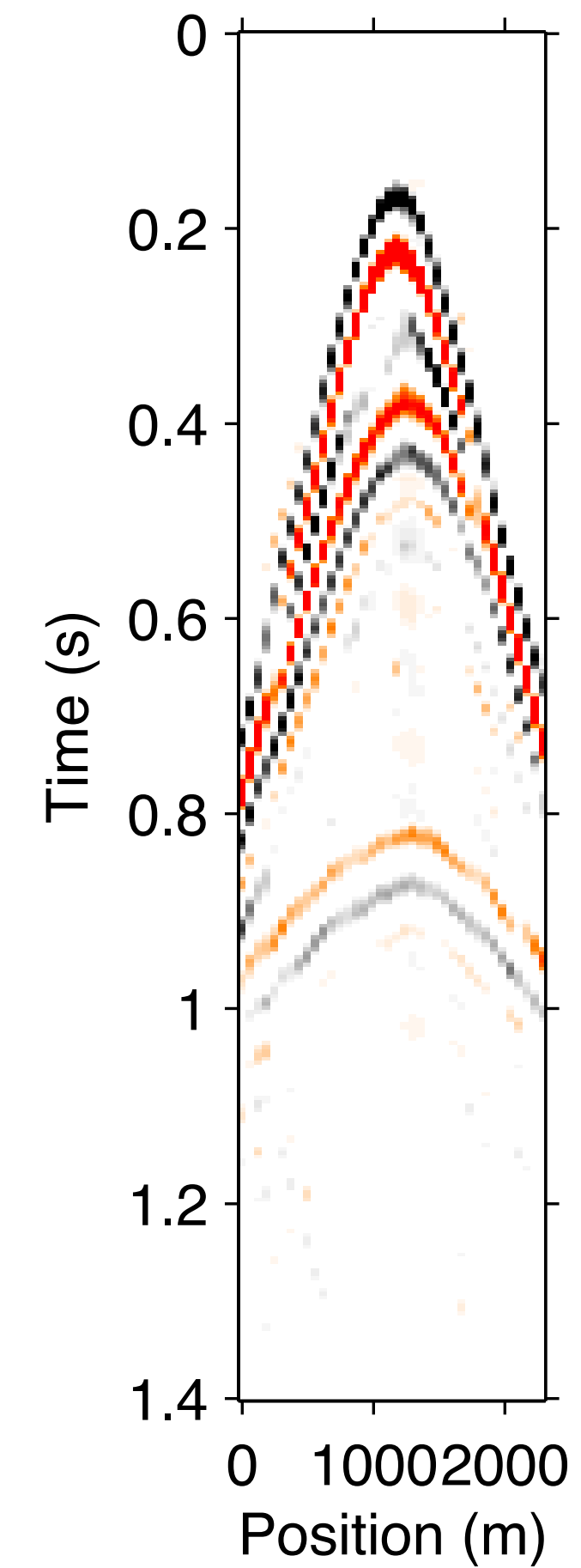
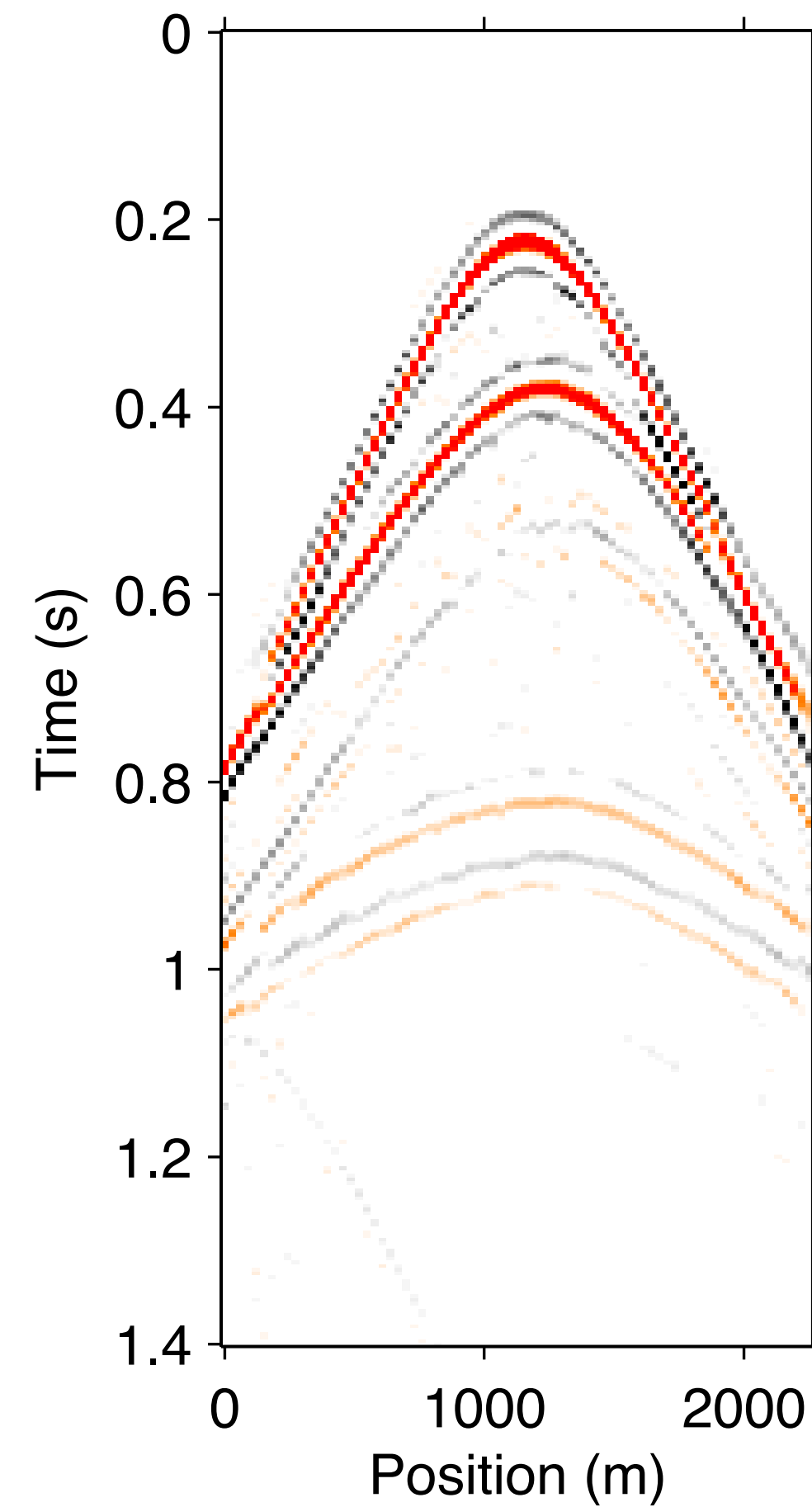
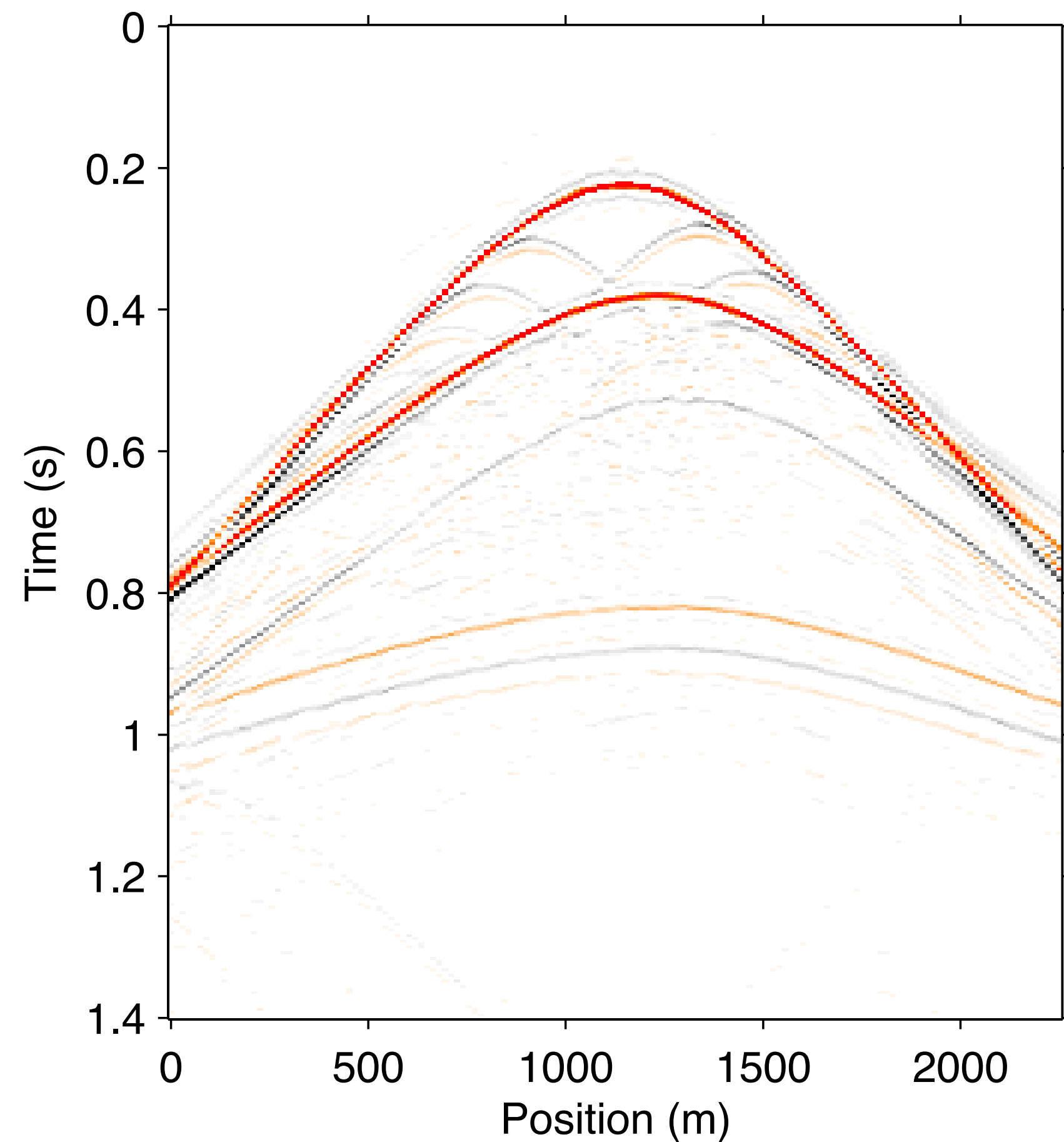
4x decimated  
lowpass 60Hz





# Idea 1: Solve REPSI on decimated dataset (low-pass in temporal frequency)

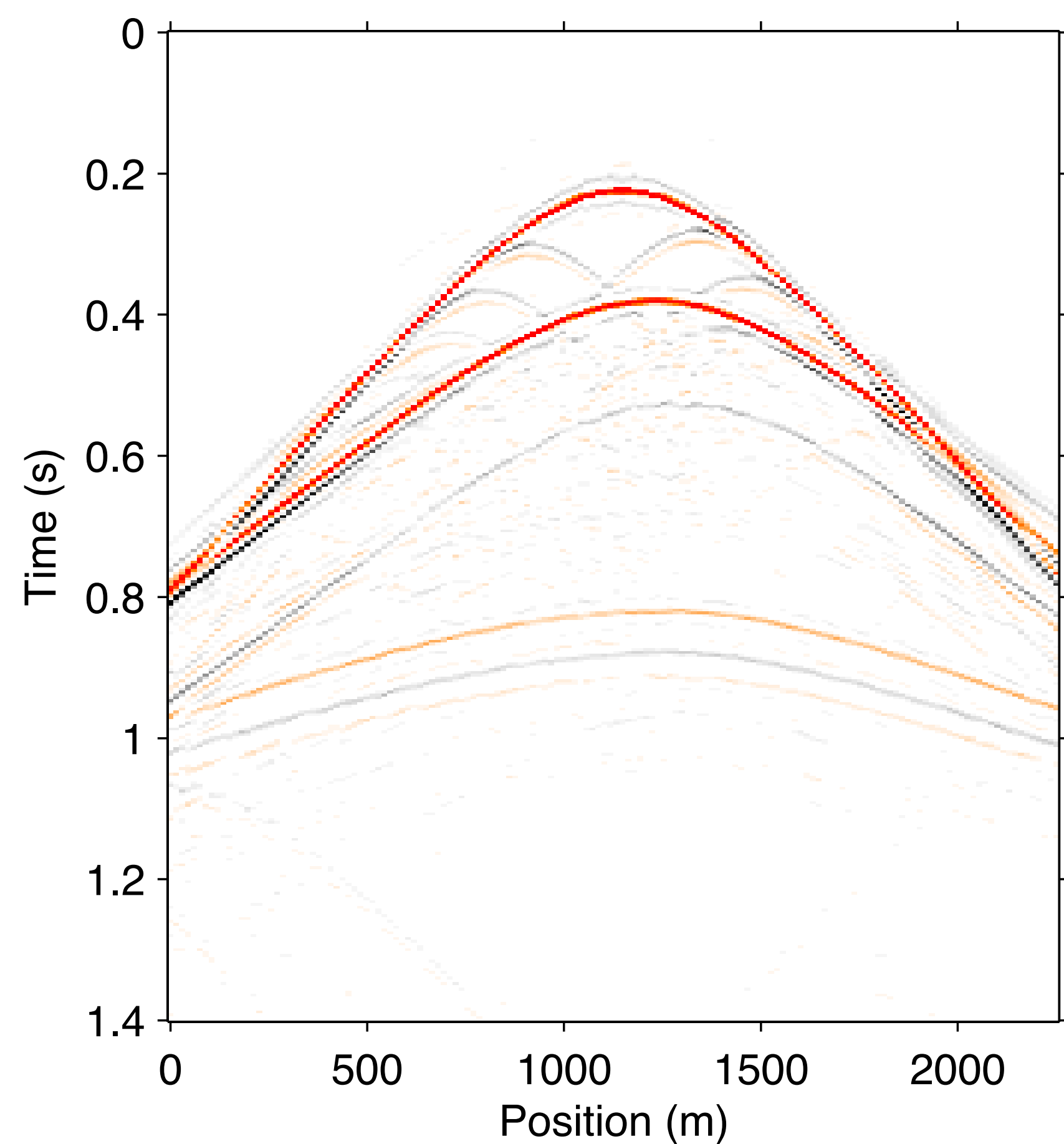
## Impulse response solutions



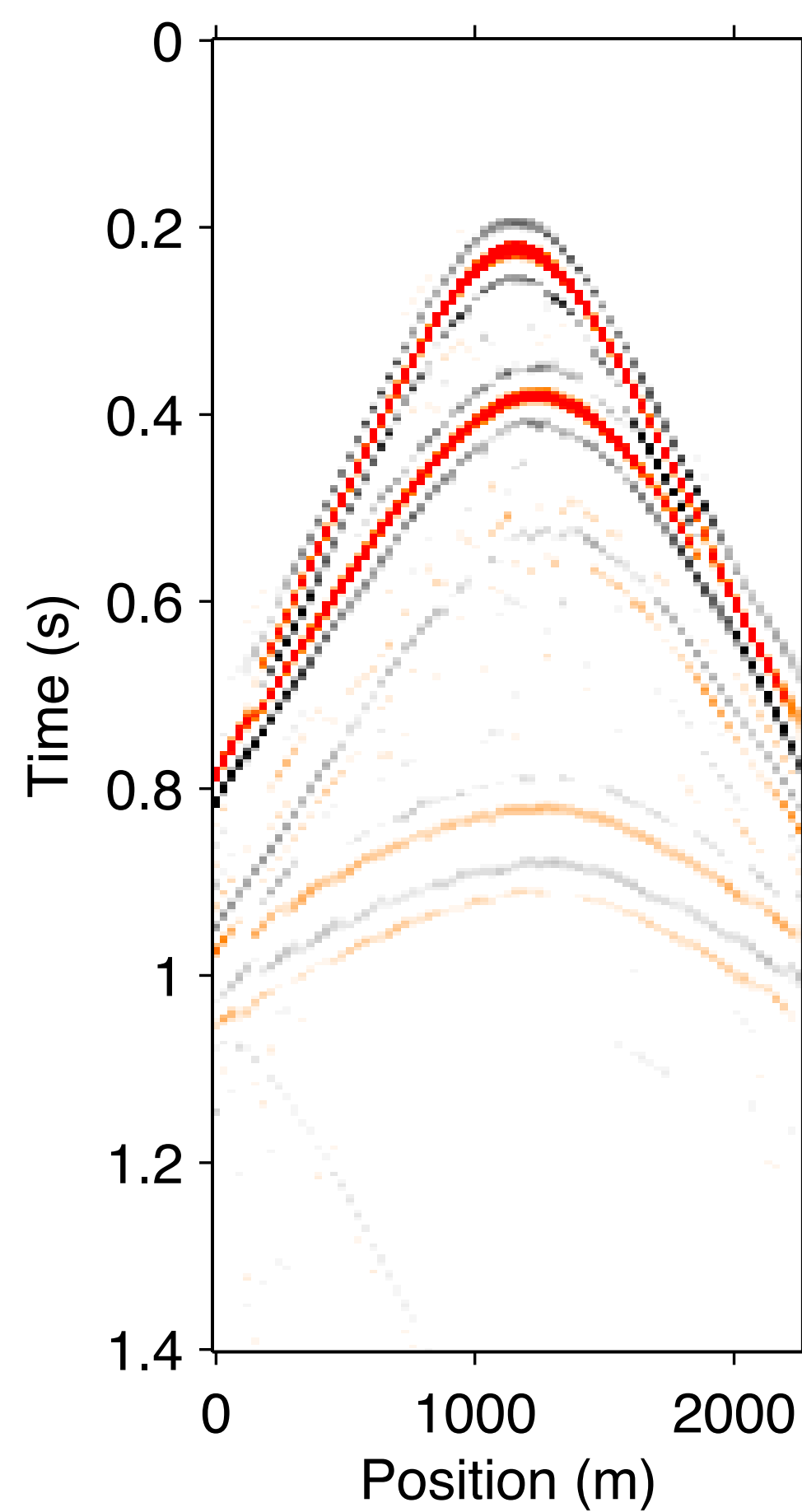


# Idea 1: Solve REPSI on decimated dataset (low-pass in temporal frequency)

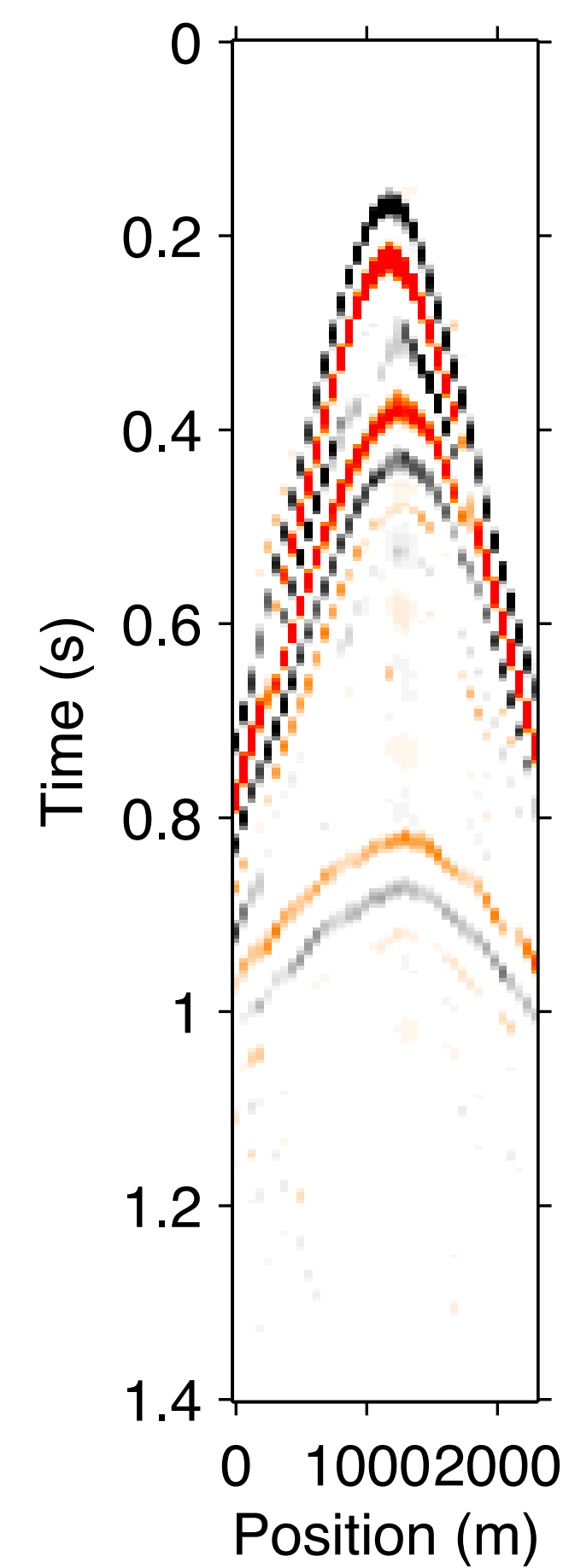
40 min



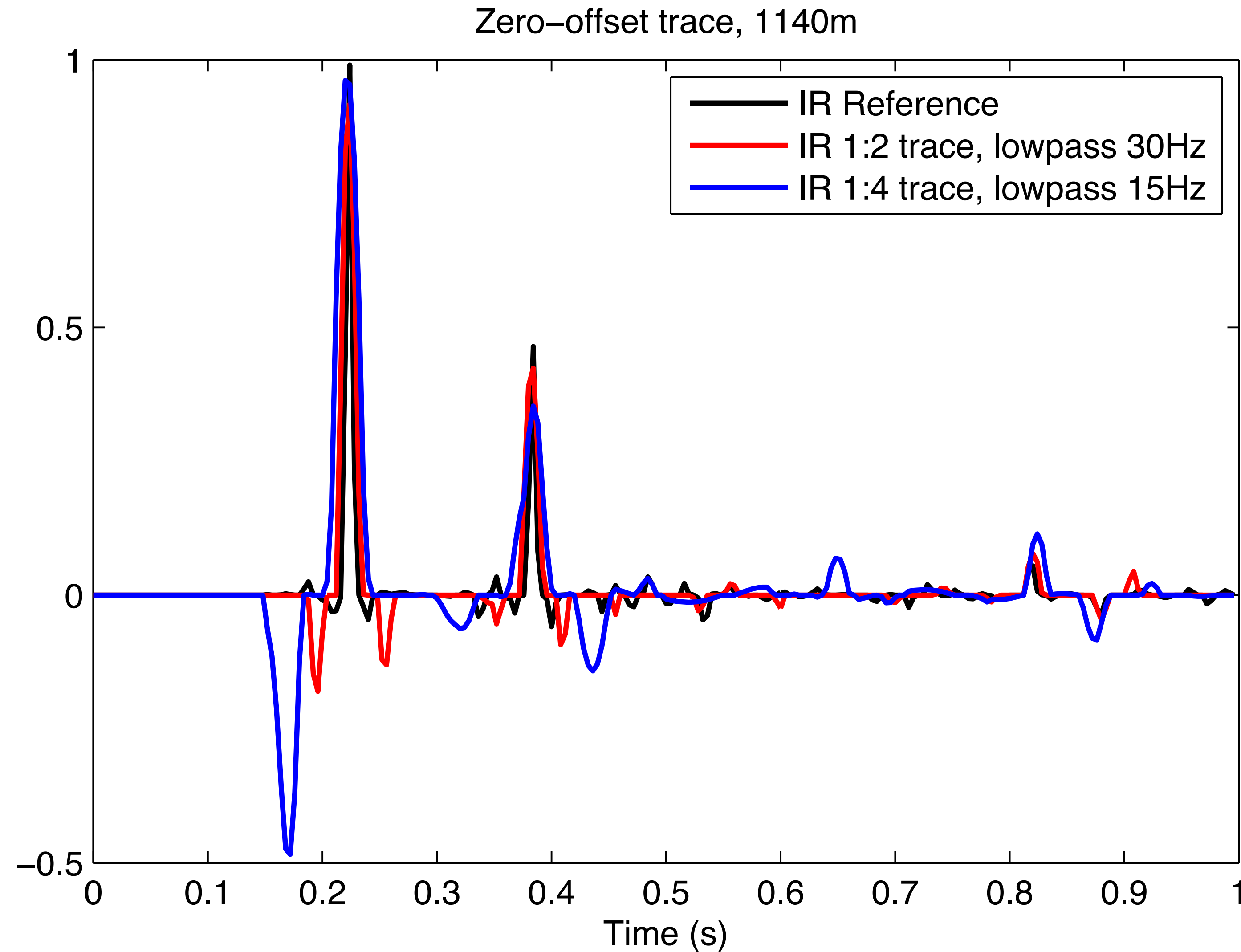
6 min



1.5 min



# Idea 1: Solve REPSI on decimated dataset (low-pass in temporal frequency)



## Idea 2: Warm-start with coarse data solutions

Since decimated datasets solve much faster, we use its (slightly inaccurate) results to replace early estimates to full problem

Initial  $\mathcal{T}_k$  (one-norm constraint) of full problem obtained by interpolating coarse solution, calculate one-norm, then scale back by some ratio

Previous Q is discarded

Interpolation method of G not important, just can't alias

Simple constant NMO (i.e., at water velocity) + linear interpolation works fine

By the way...

**If you hear  
any of these:**

MNO  
Animo  
Animal  
Asimov  
Nemo  
Anemone  
Eminem  
M&M's  
Nominal  
Dominos  
Ememo  
Wrararar  
Waka-waka

**it's just me  
trying to say:**

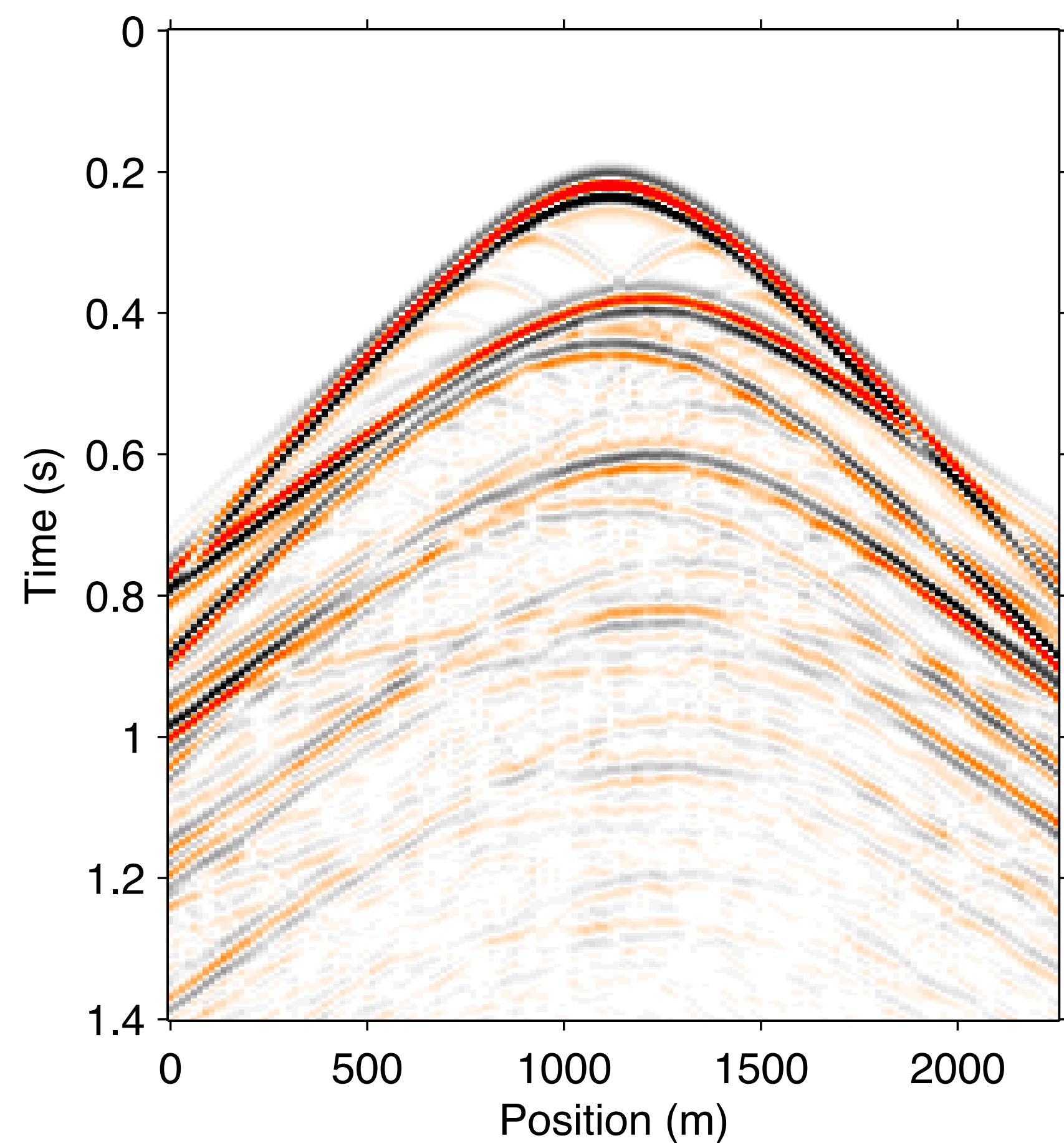
**NMO** (very badly)



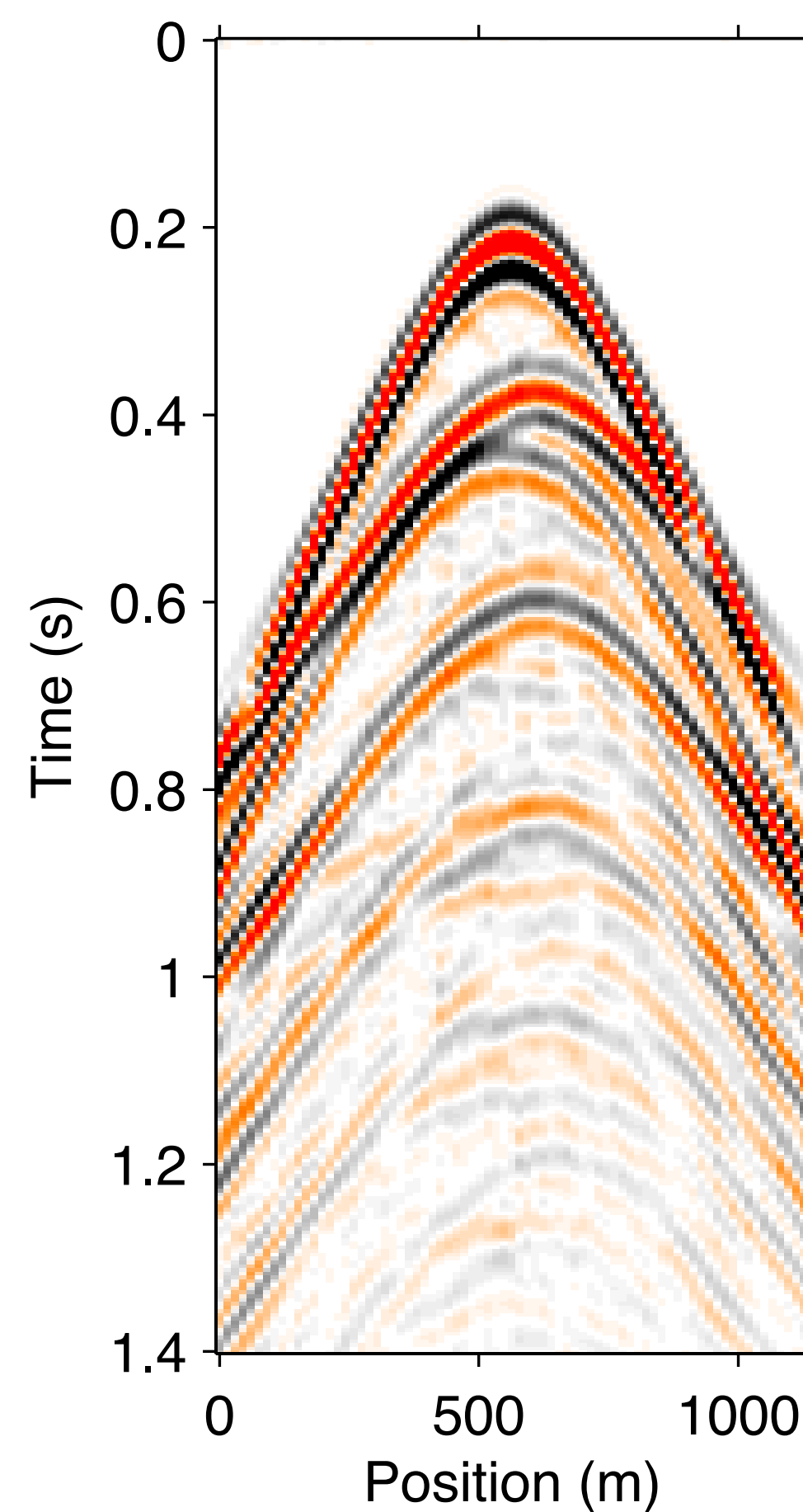
# Warm-starting/continuation from coarse solution

## Example

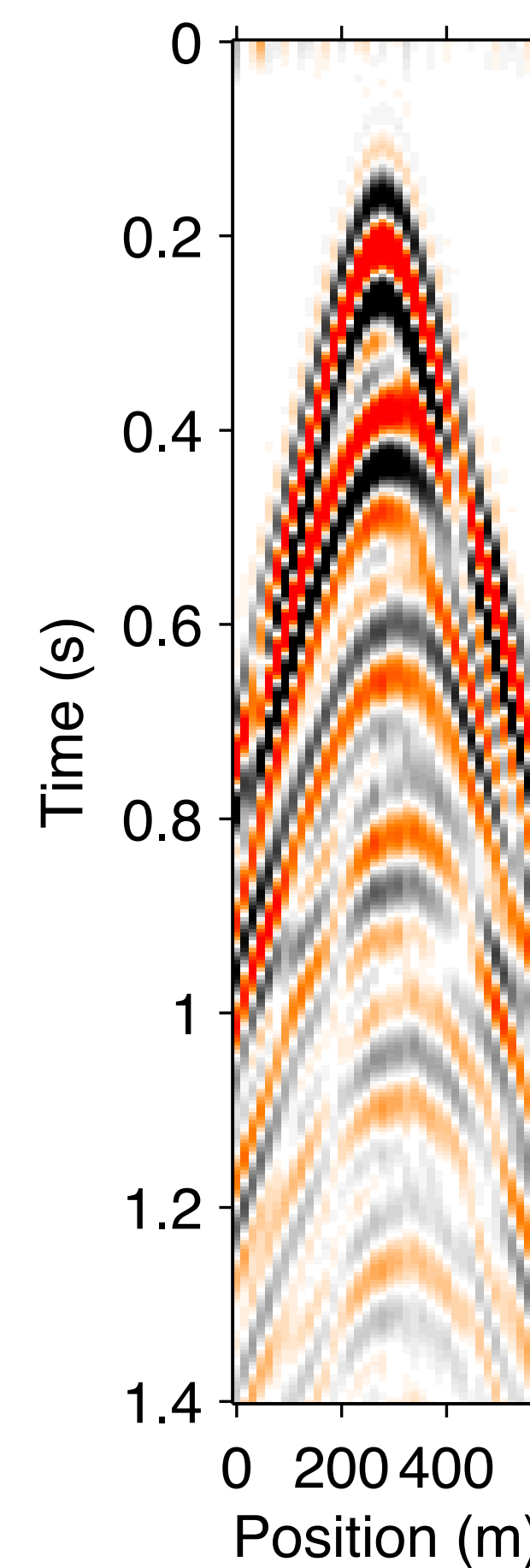
Original (dx = 15m)



2x decimated  
lowpass 30Hz



4x decimated  
lowpass 60Hz

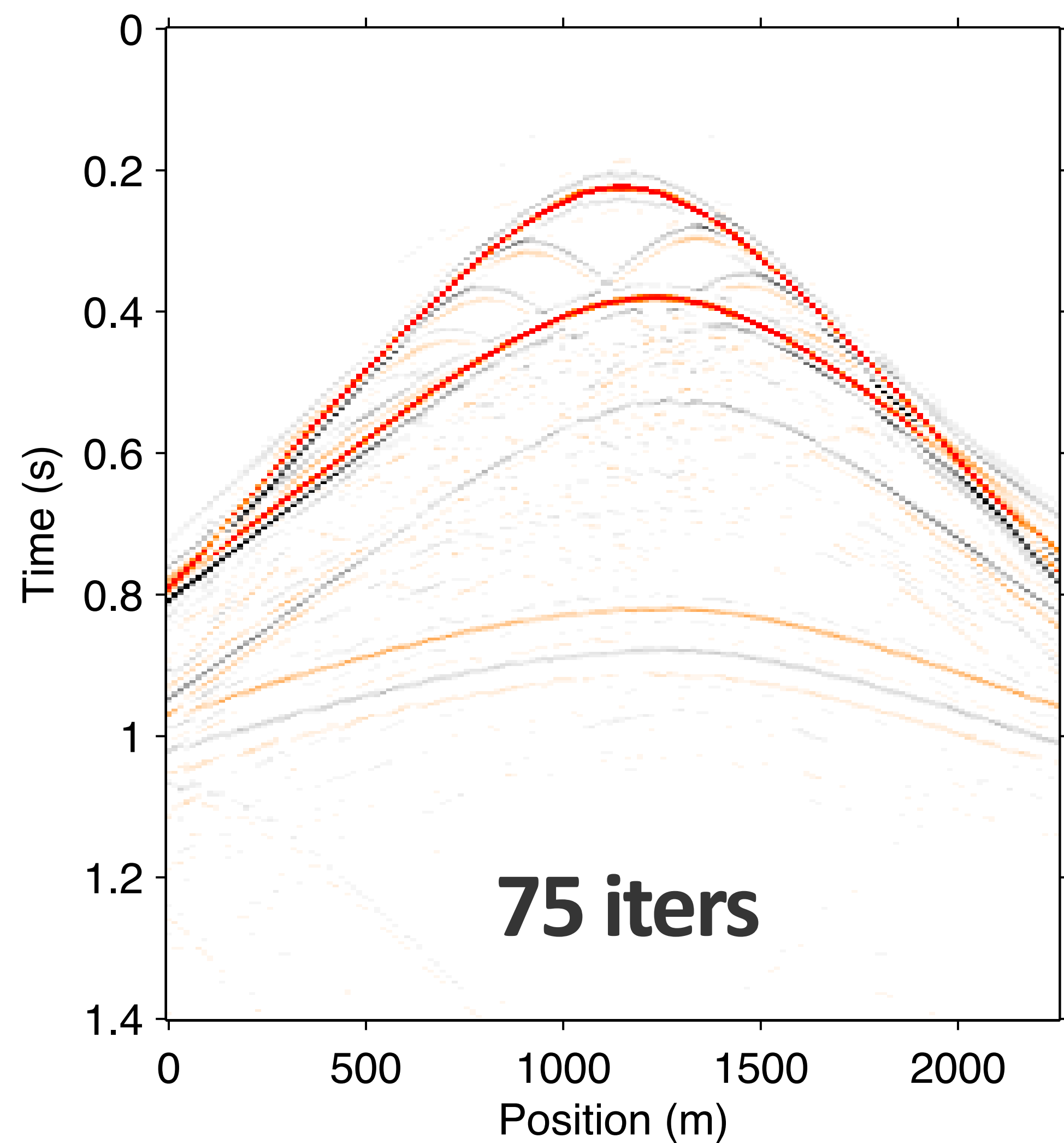




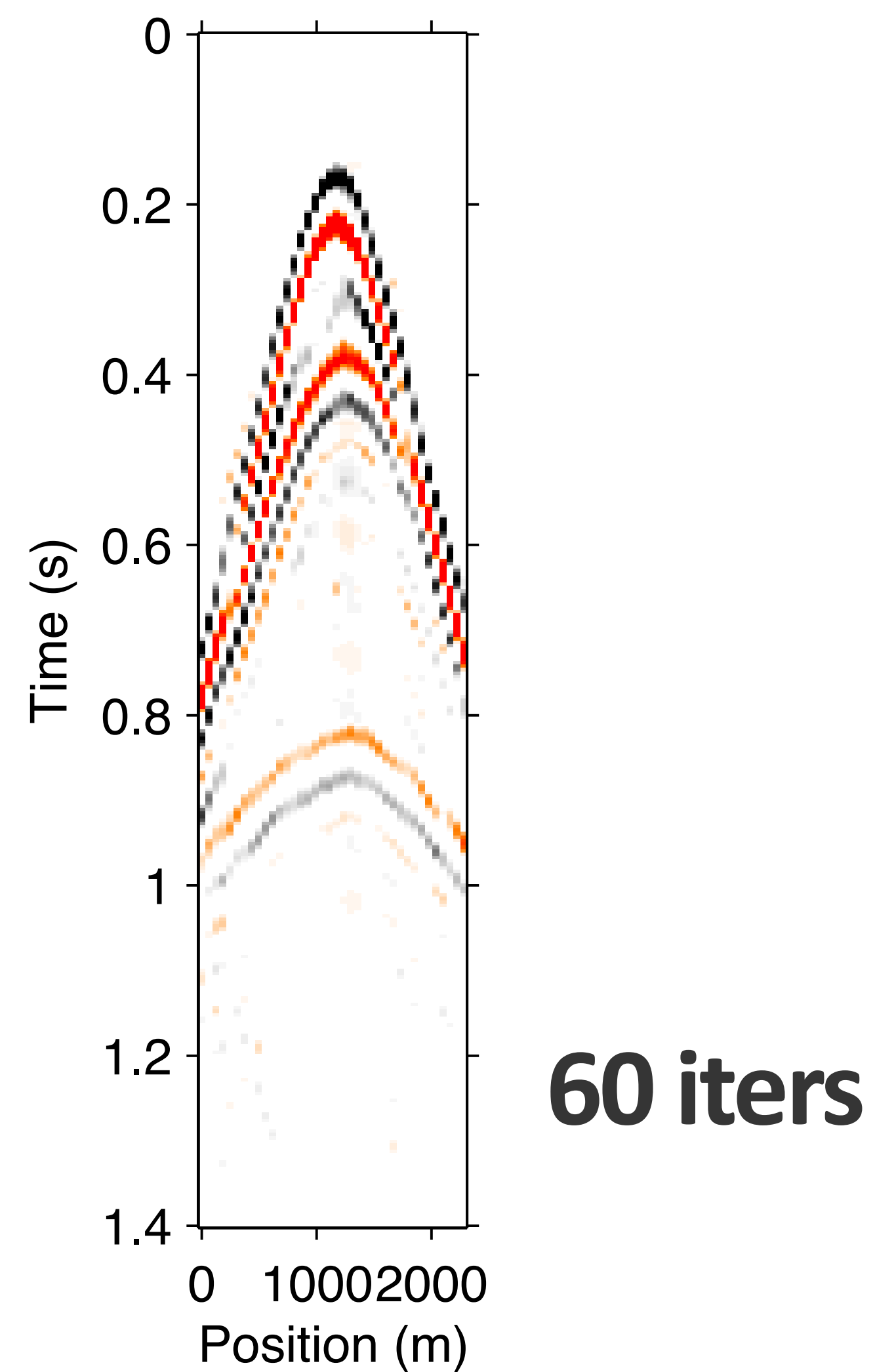
# Warm-starting/continuation from coarse solution

## Example

Solution of full data



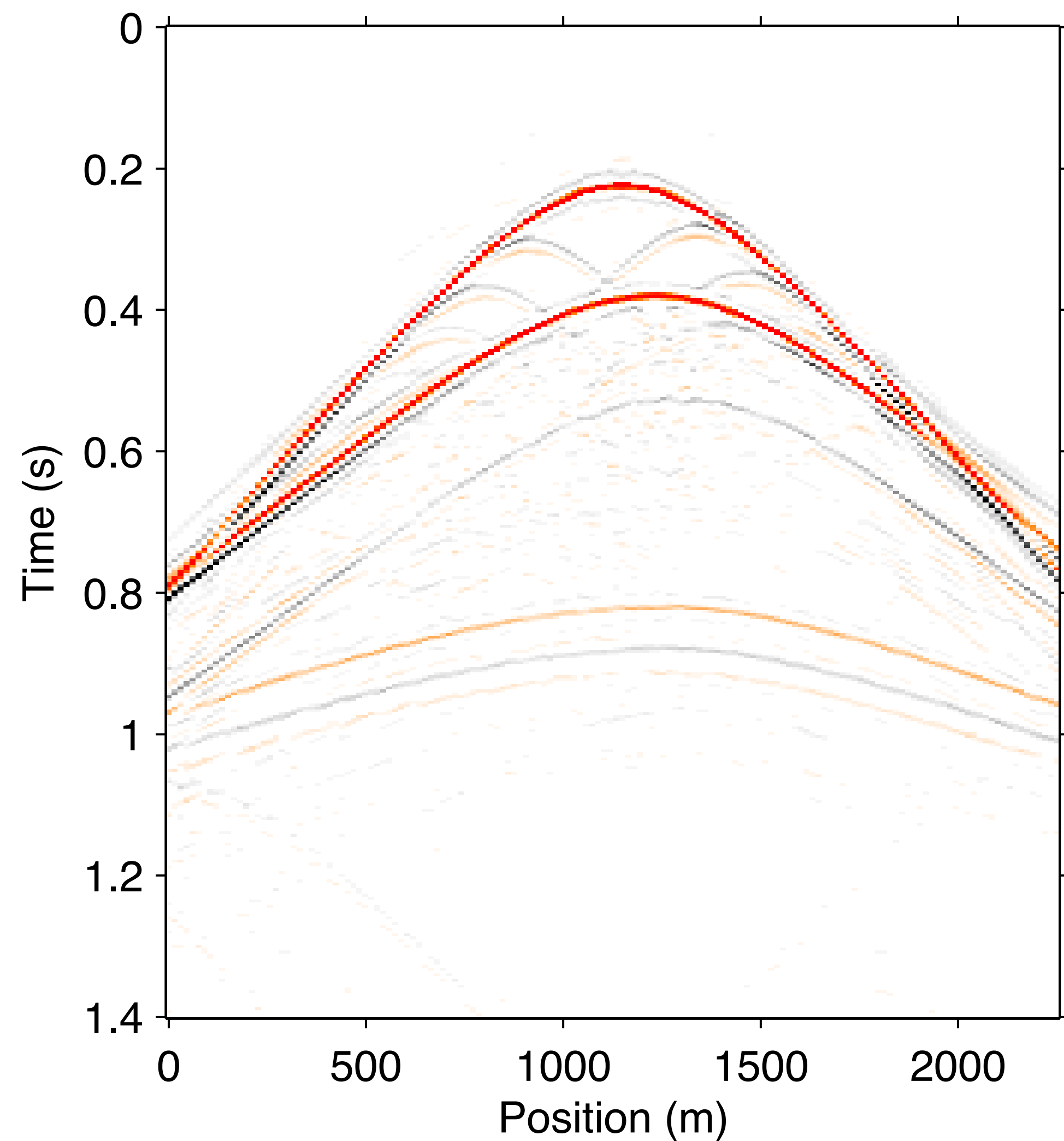
Solution of 4x decimated data



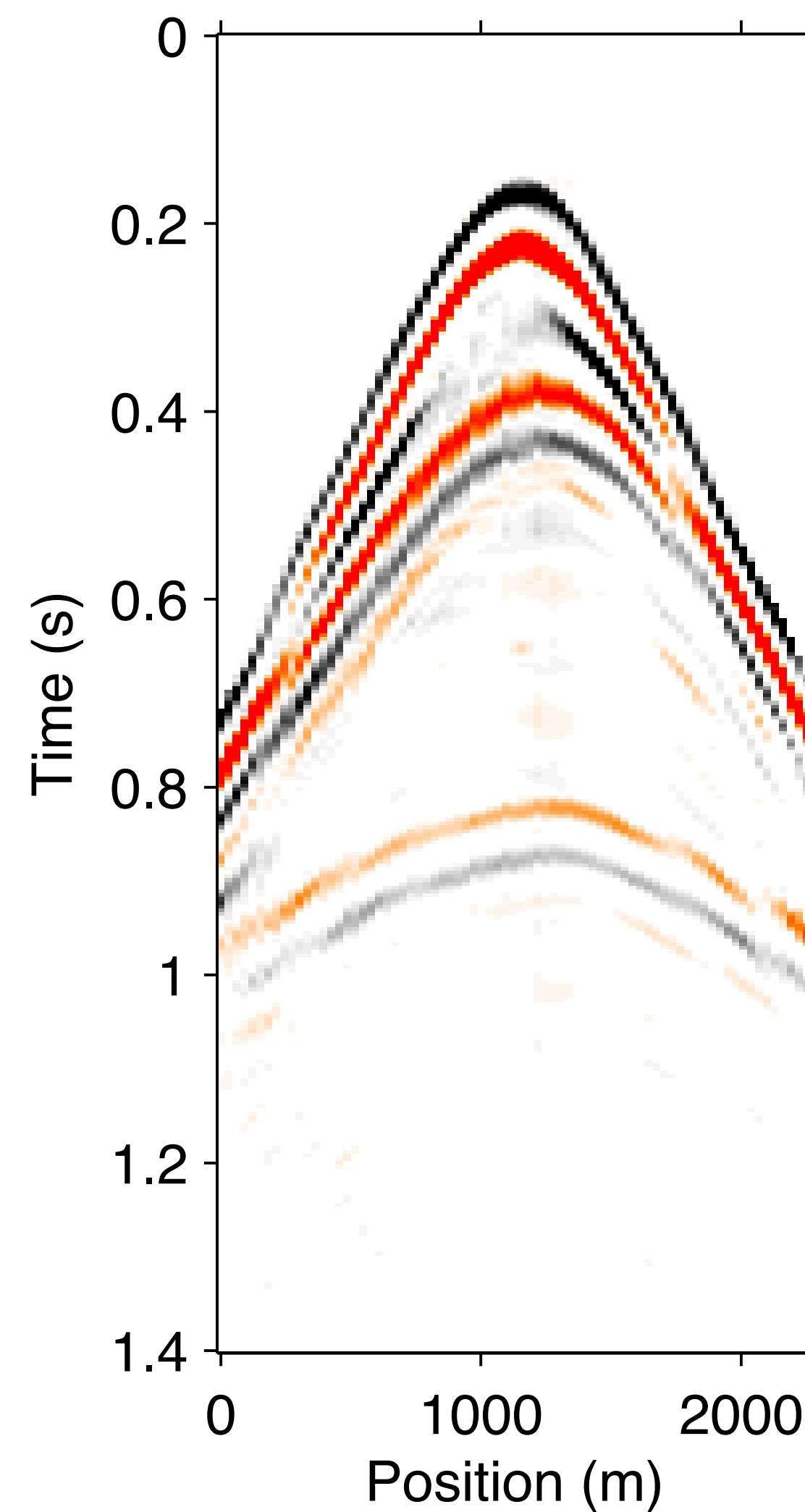
# Warm-starting/continuation from coarse solution

## Example

Solution of full data



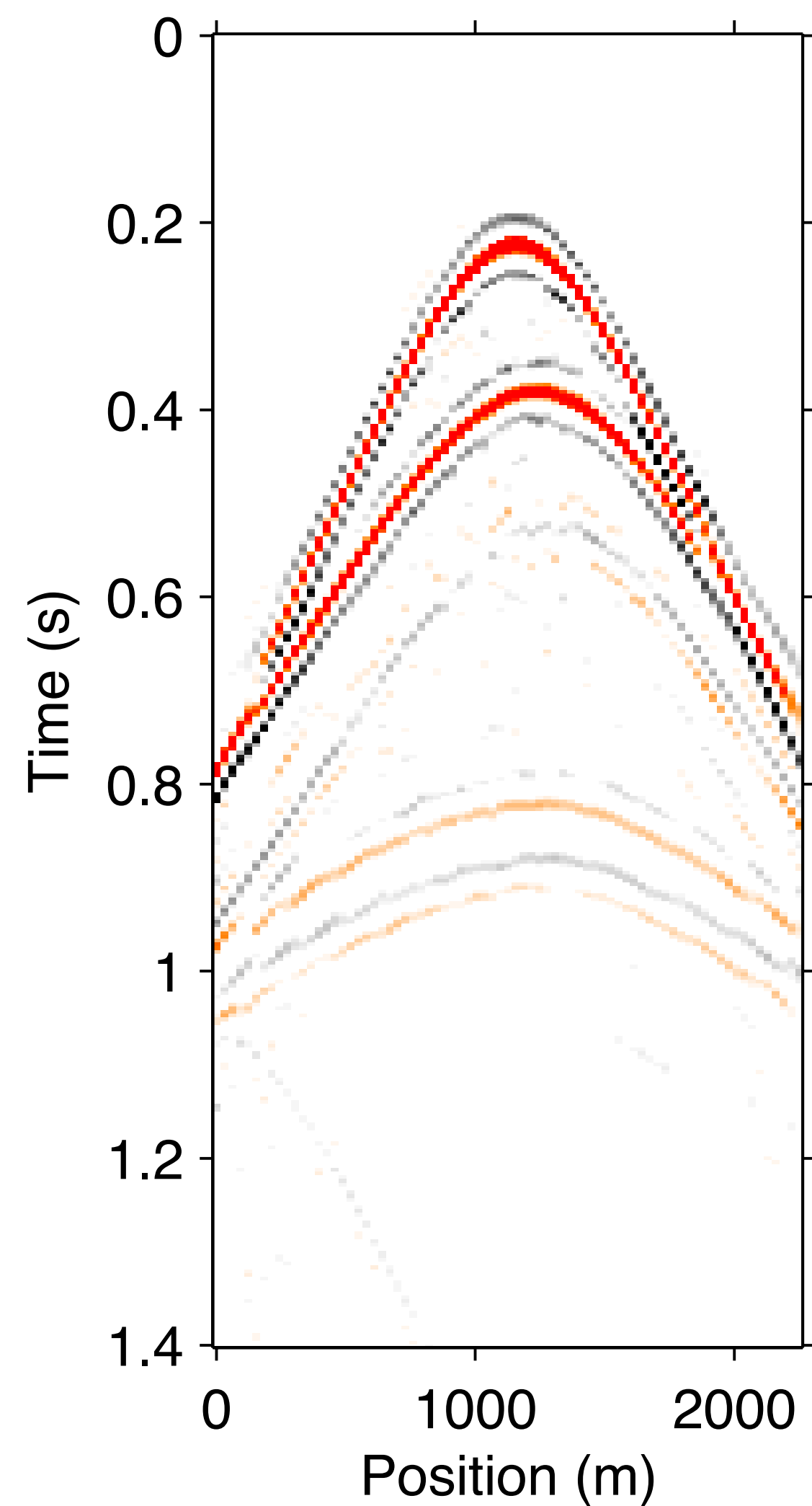
Solution of 4x decimated data  
constant NMO linear interp 2x



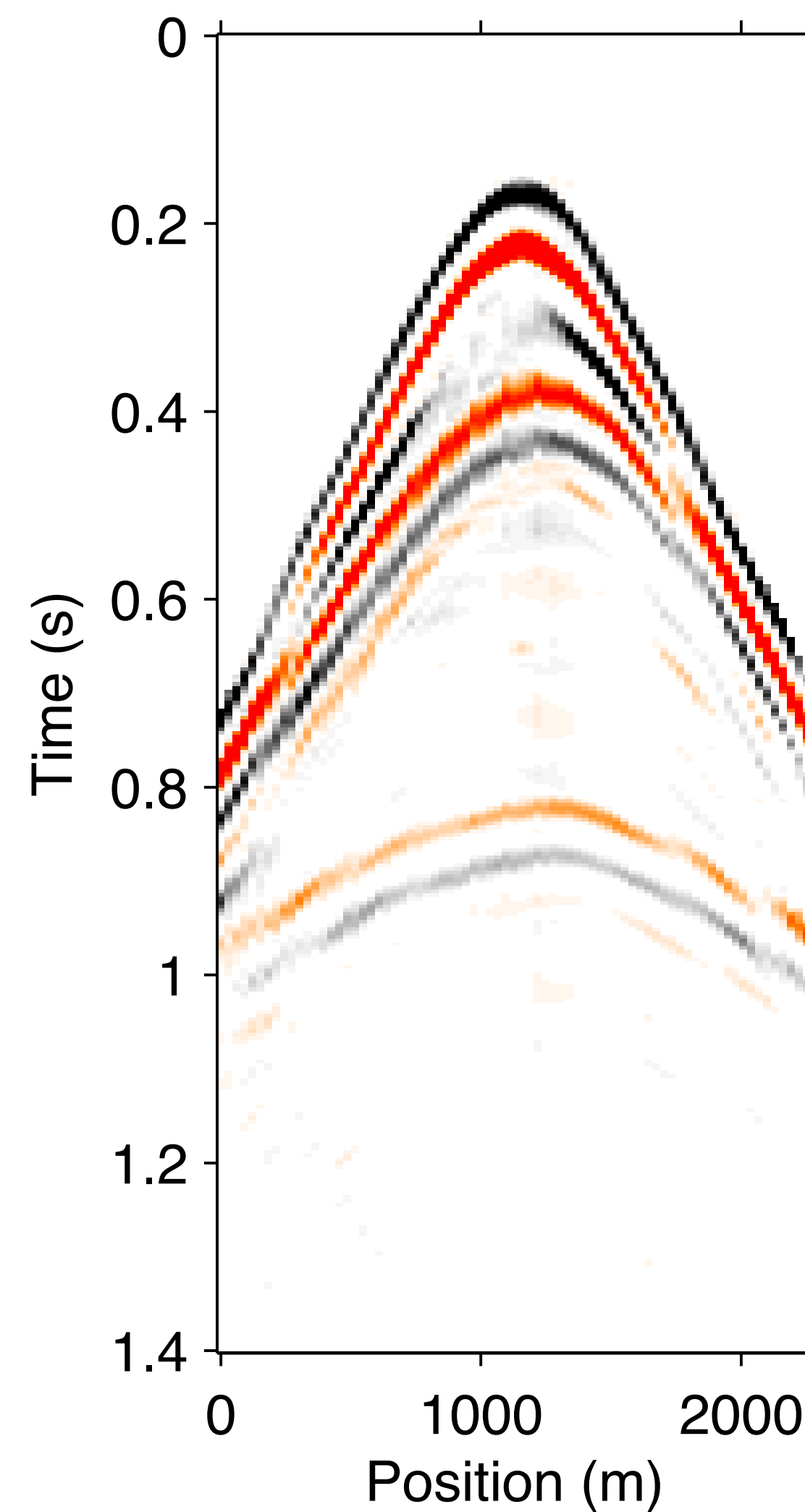
# Warm-starting/continuation from coarse solution

## Example

Solution of 2x decimated data



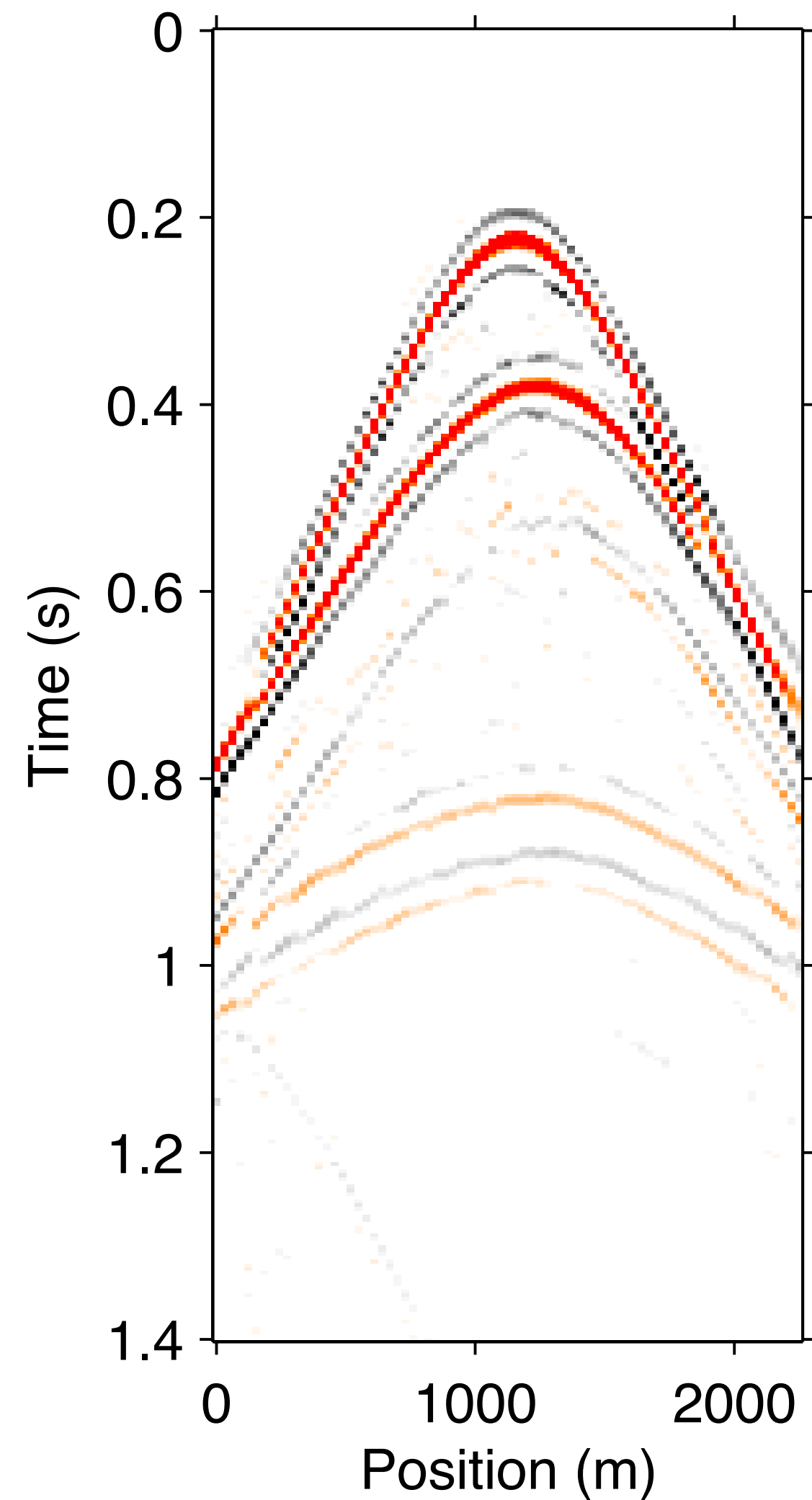
Solution of 4x decimated data  
constant NMO linear interp 2x



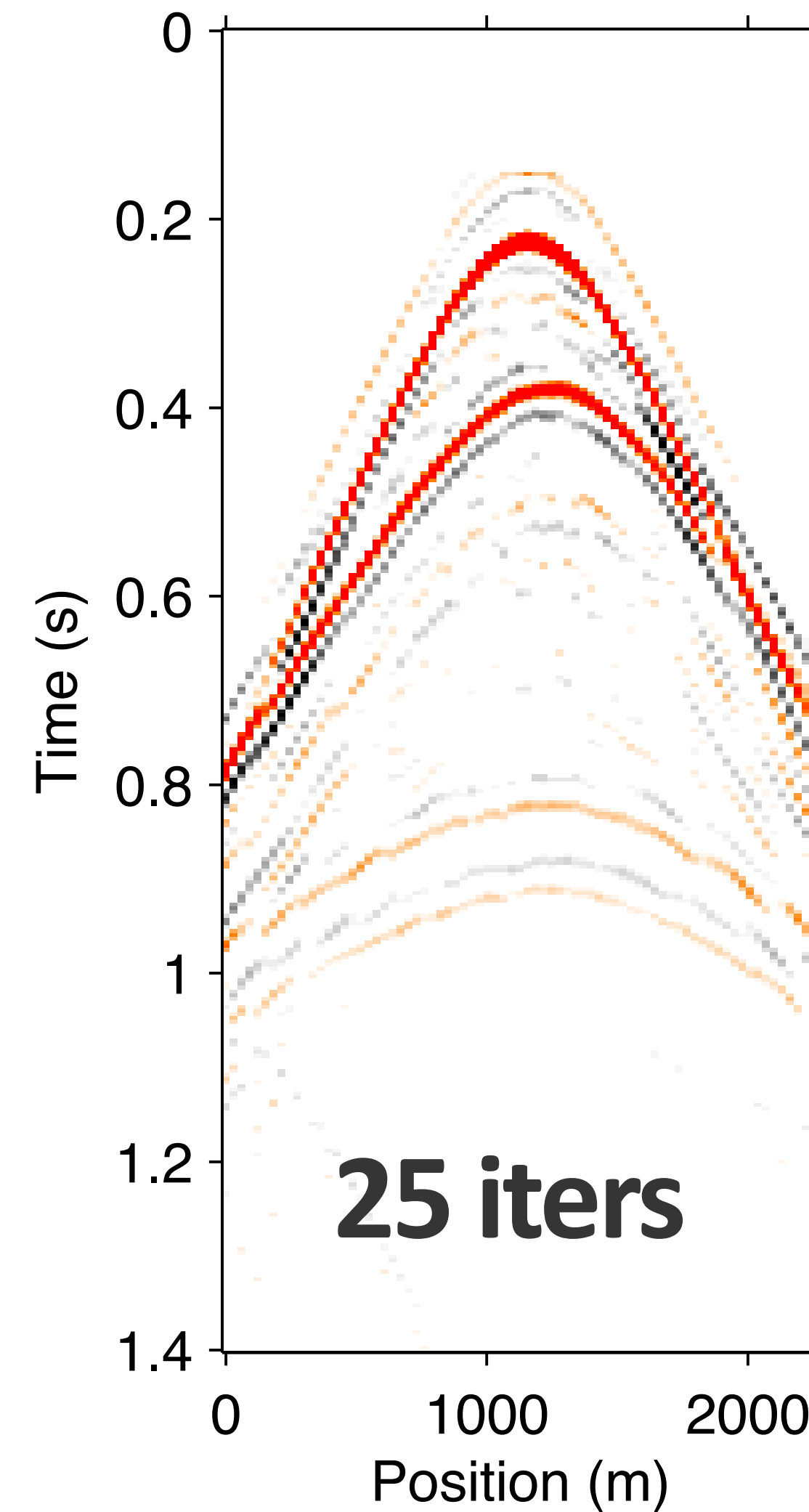
# Warm-starting/continuation from coarse solution

## Example

Solution of 2x decimated data



Solution on 2x dec data  
*continuation* from 4x dec solution

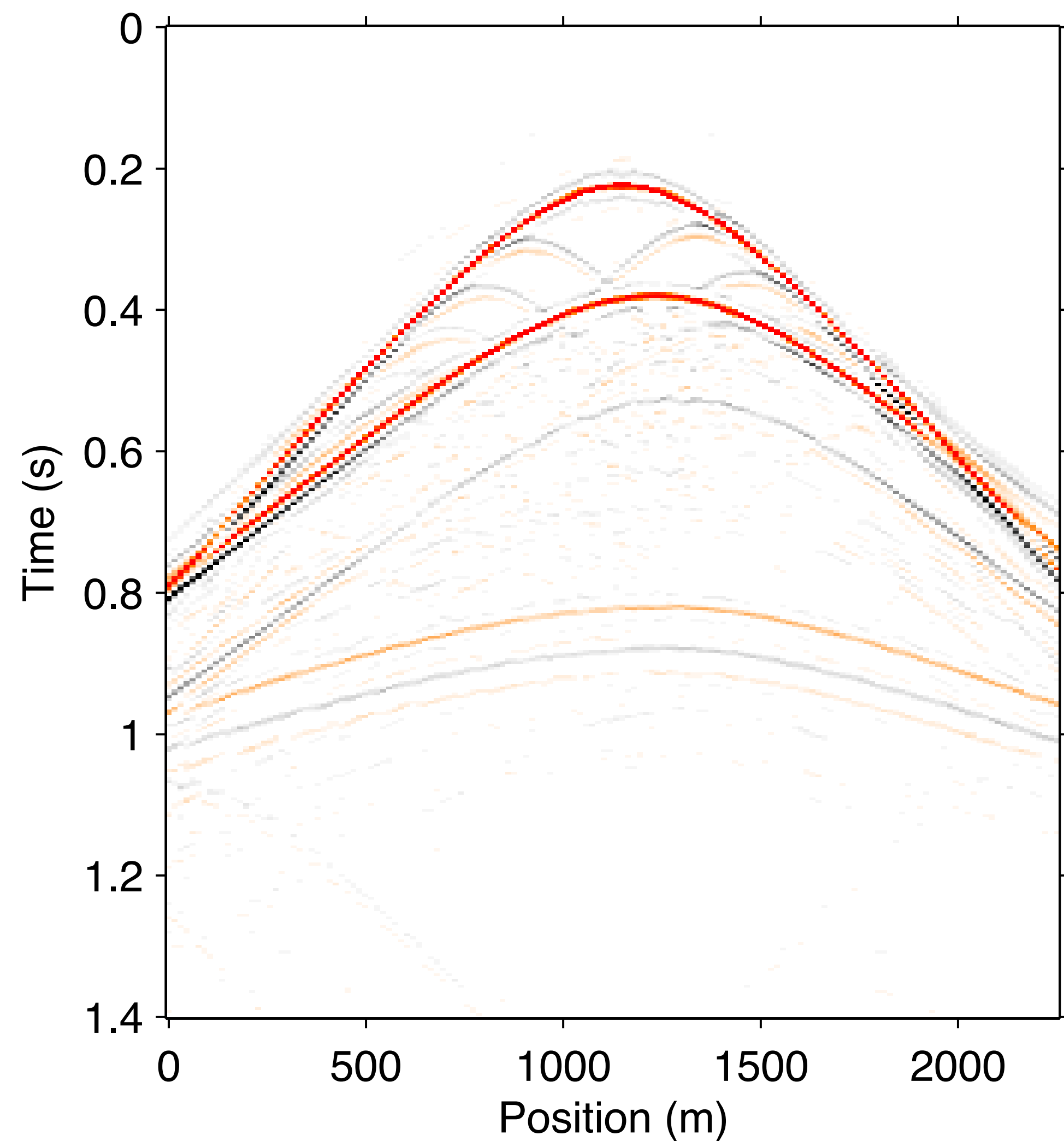




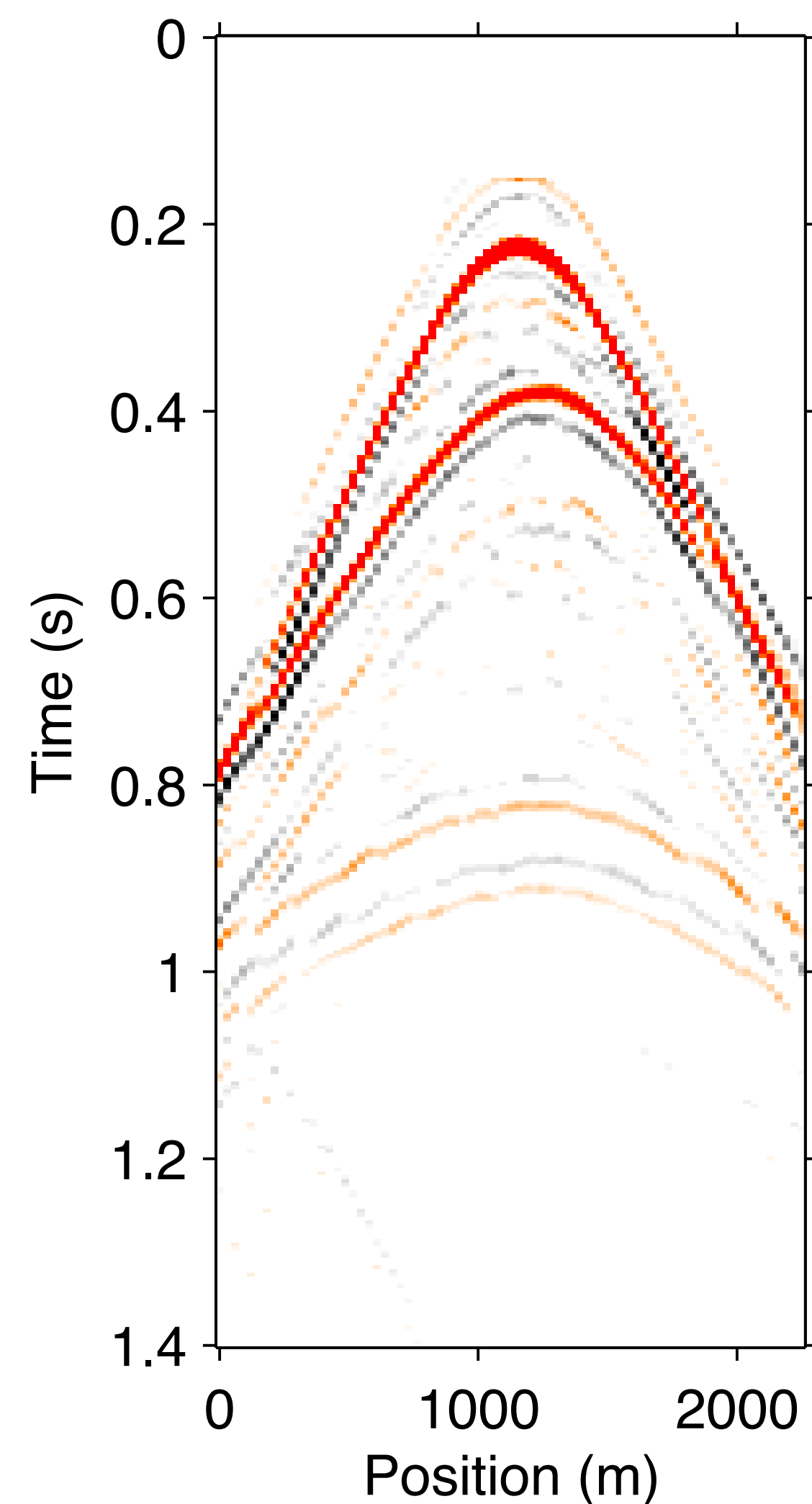
# Warm-starting/continuation from coarse solution

## Example

Solution of full data



Solution on 2x dec data  
*continuation* from 4x dec solution

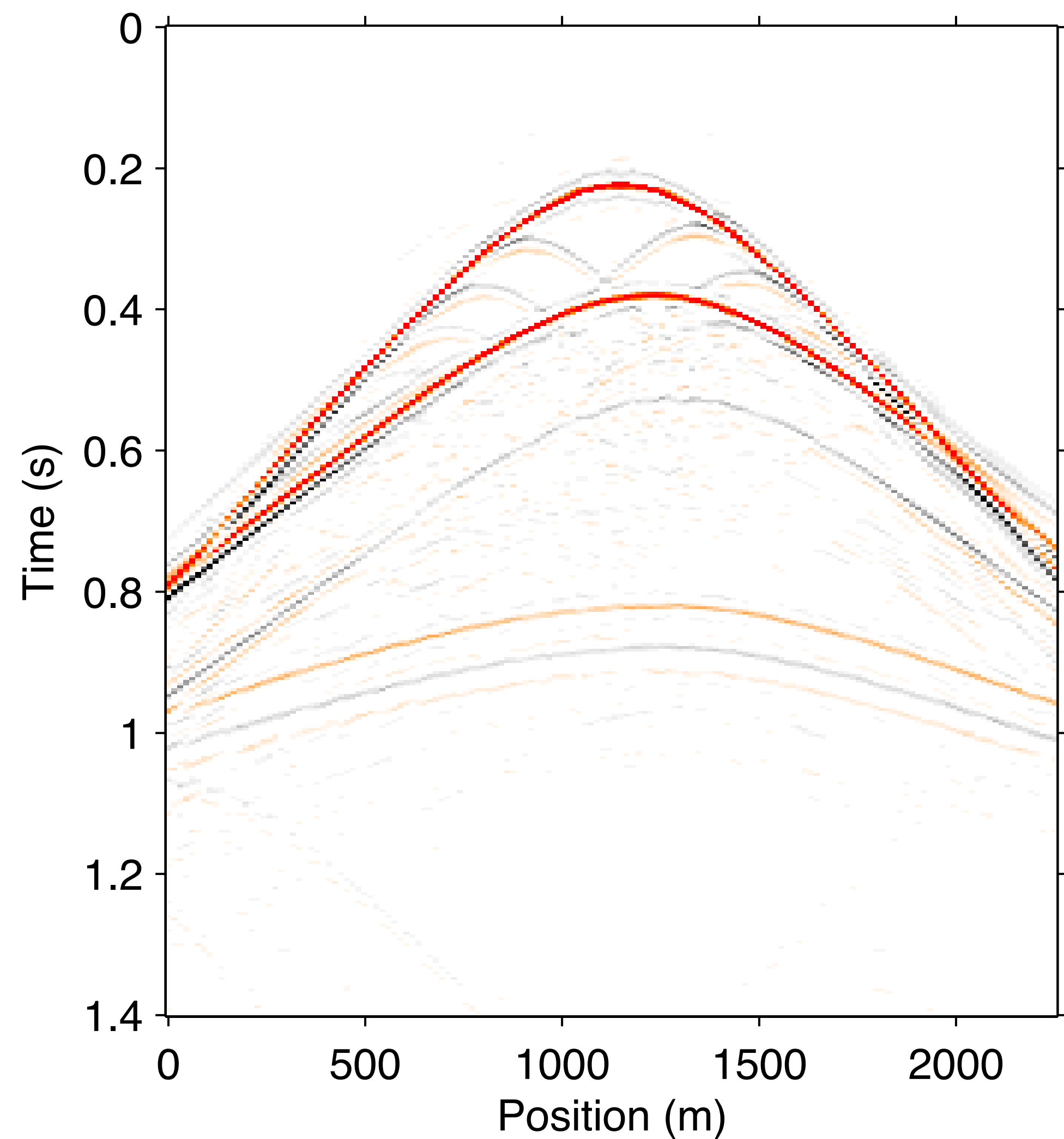




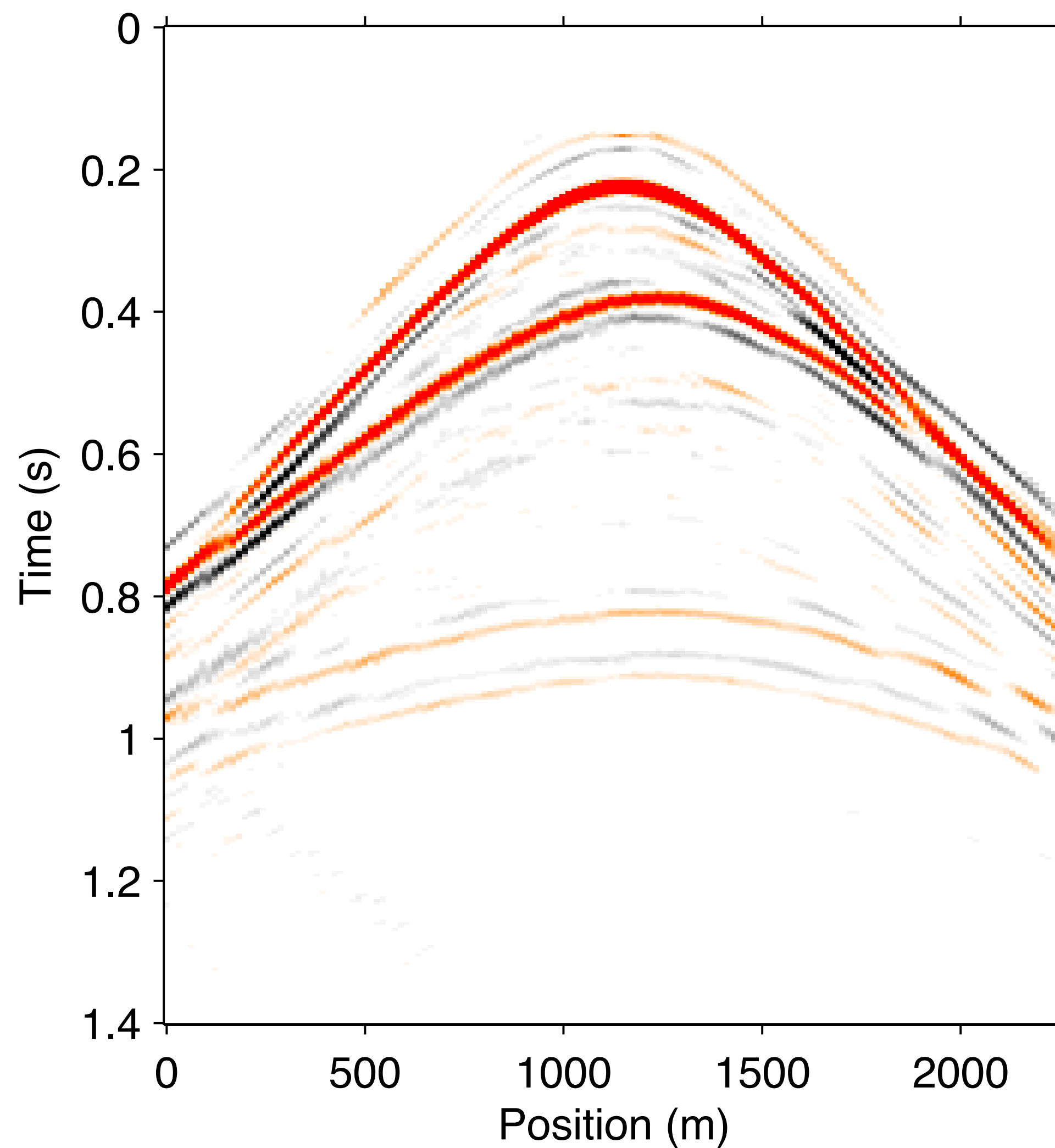
# Warm-starting/continuation from coarse solution

## Example

Solution of full data



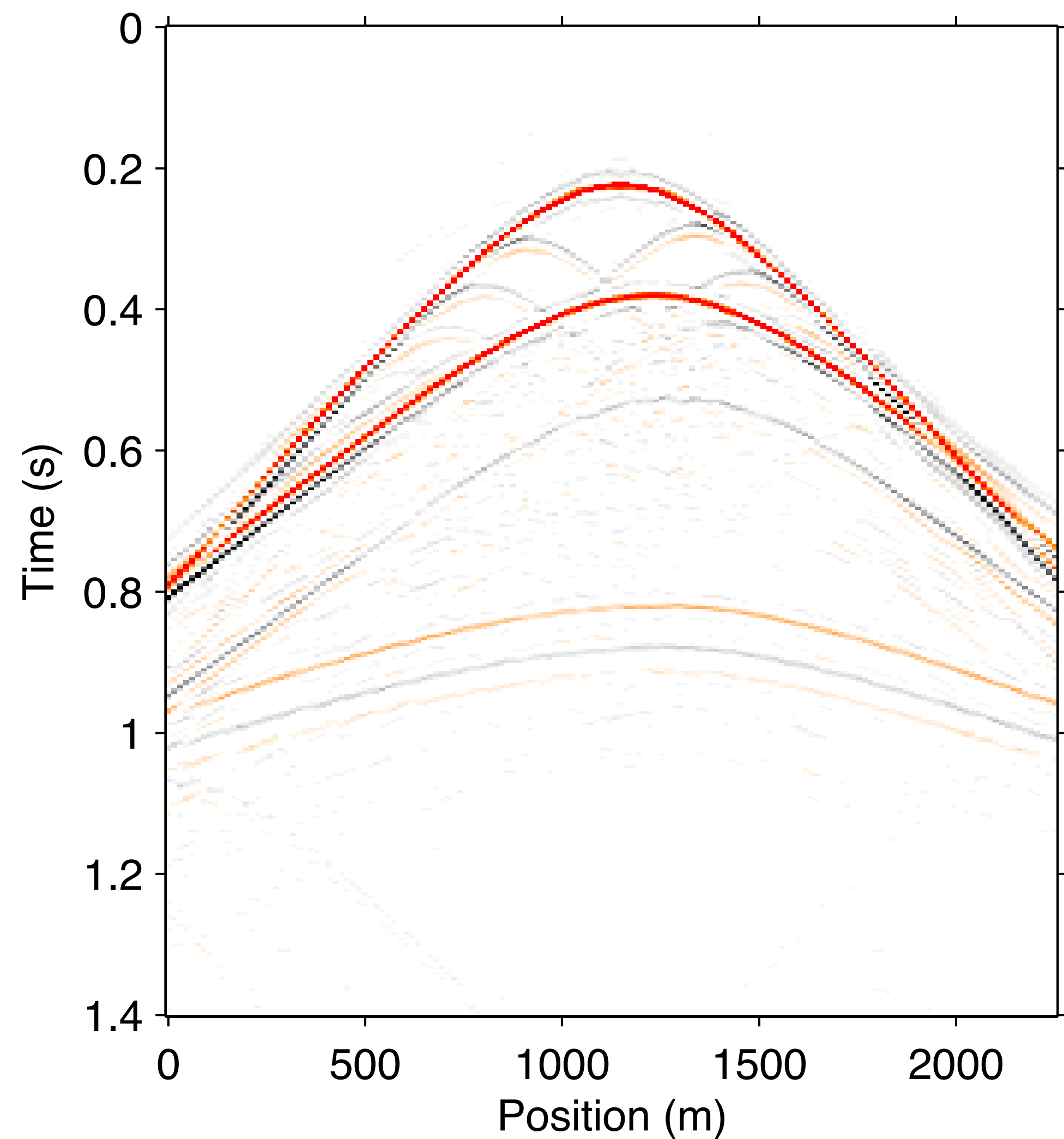
Solution on 2x dec data > interp 2x  
*continuation* from 4x dec solution



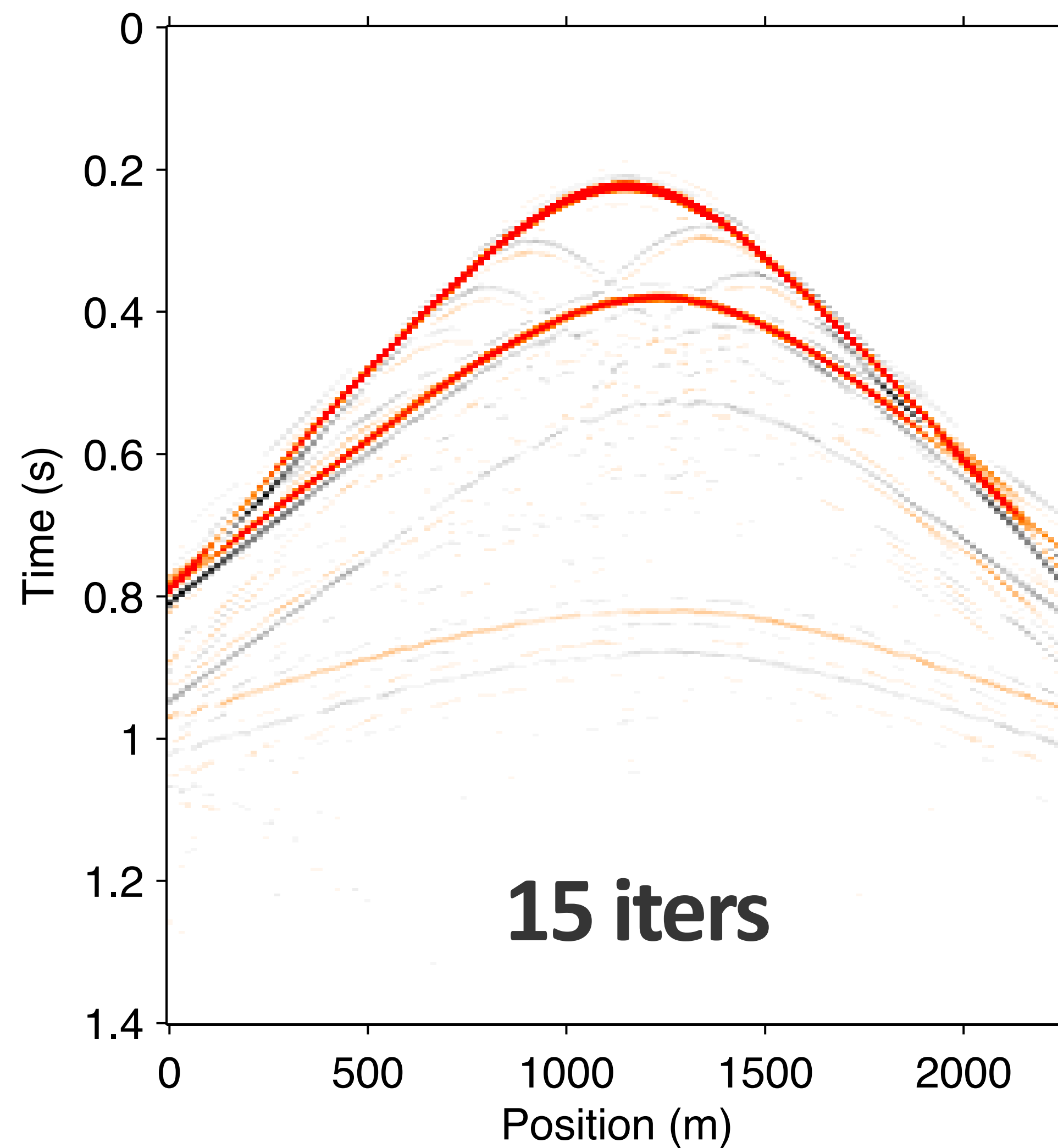
# Warm-starting/continuation from coarse solution

## Example

Solution of full data

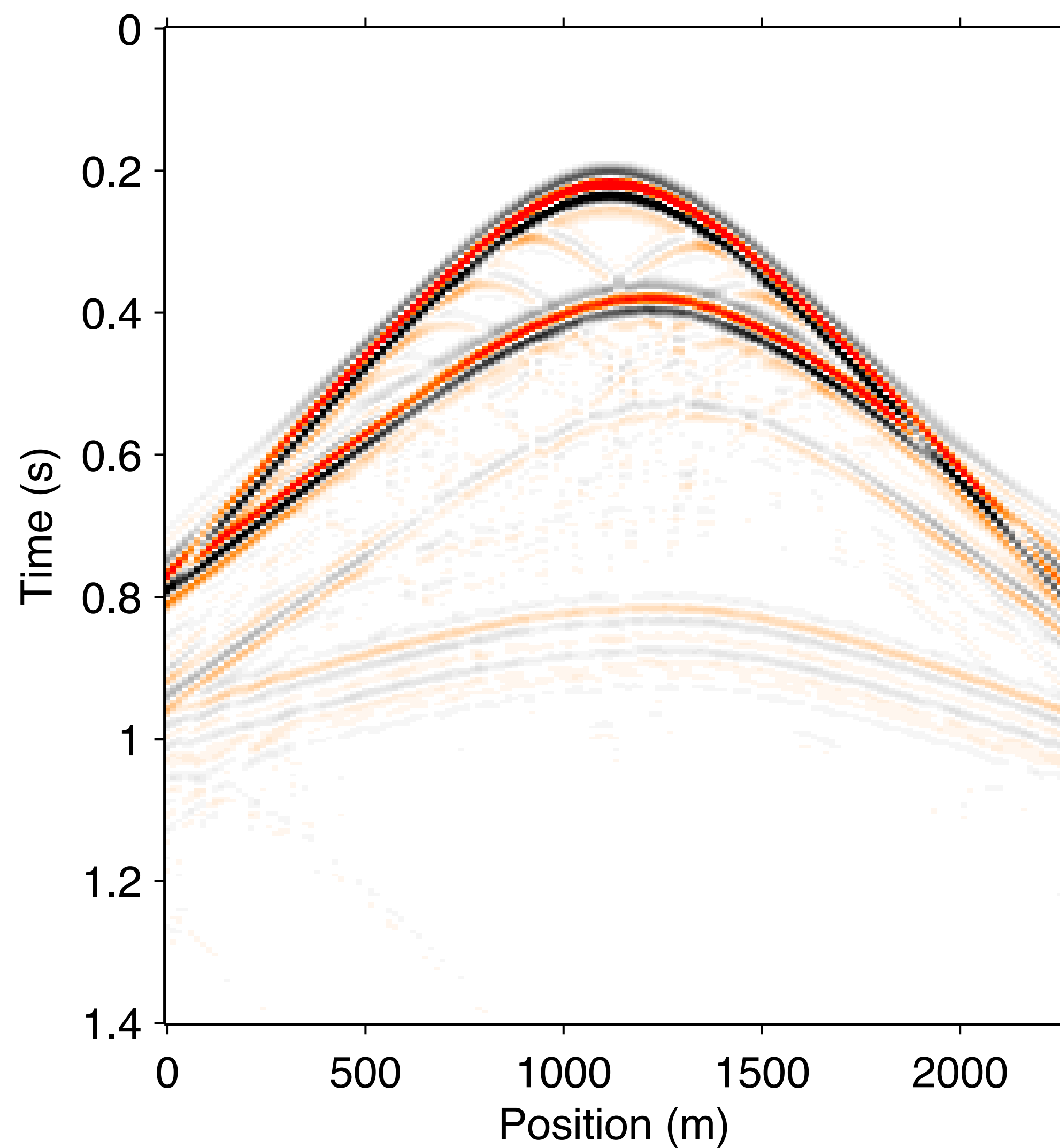


Solution on 2x dec data > interp 2x  
*continuation* from 4x thru 2x solution



# Warm-starting/continuation from coarse solution

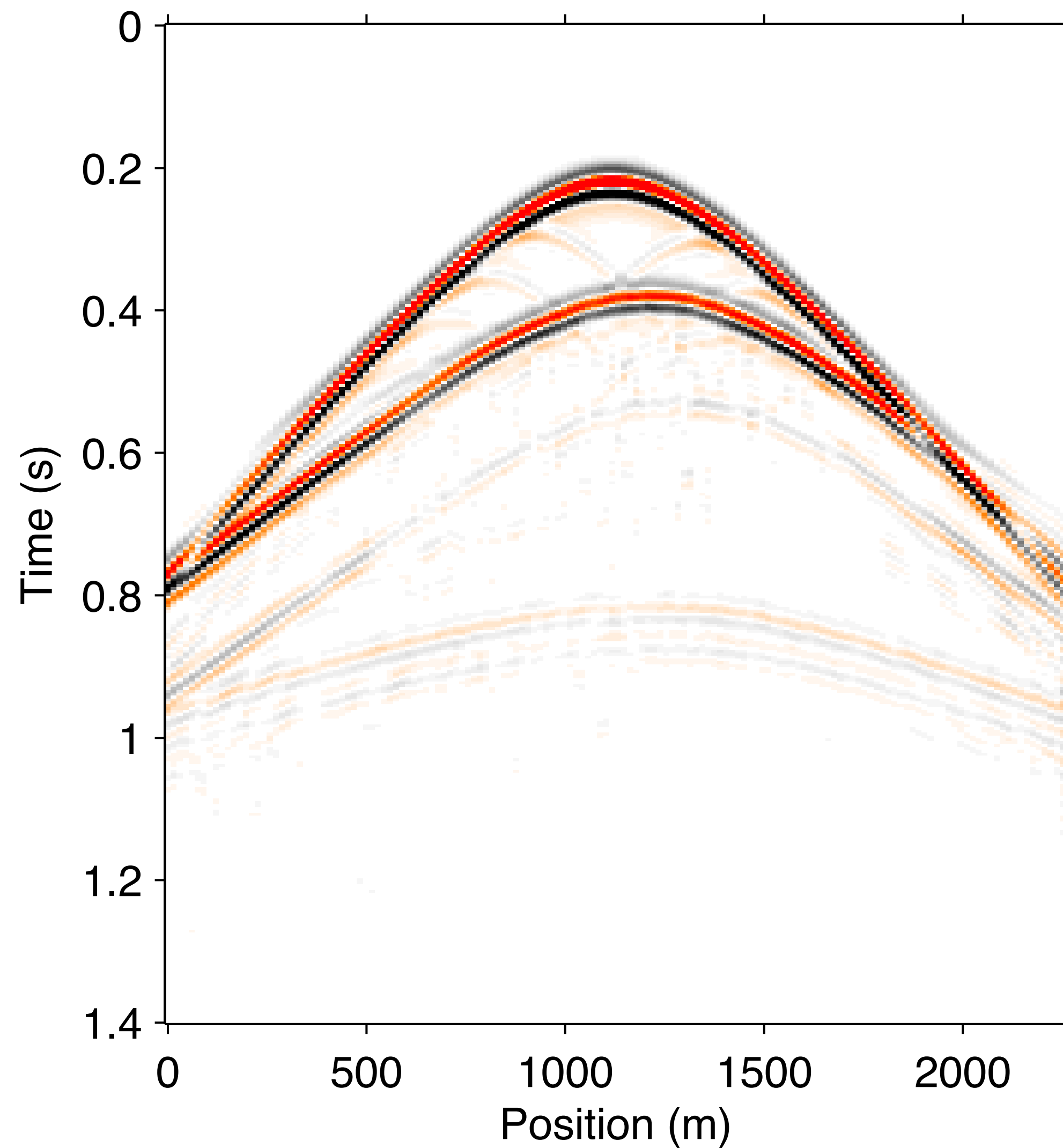
## Example



**Direct Primary**  
**Solved from full data**

# Warm-starting/continuation from coarse solution

## Example

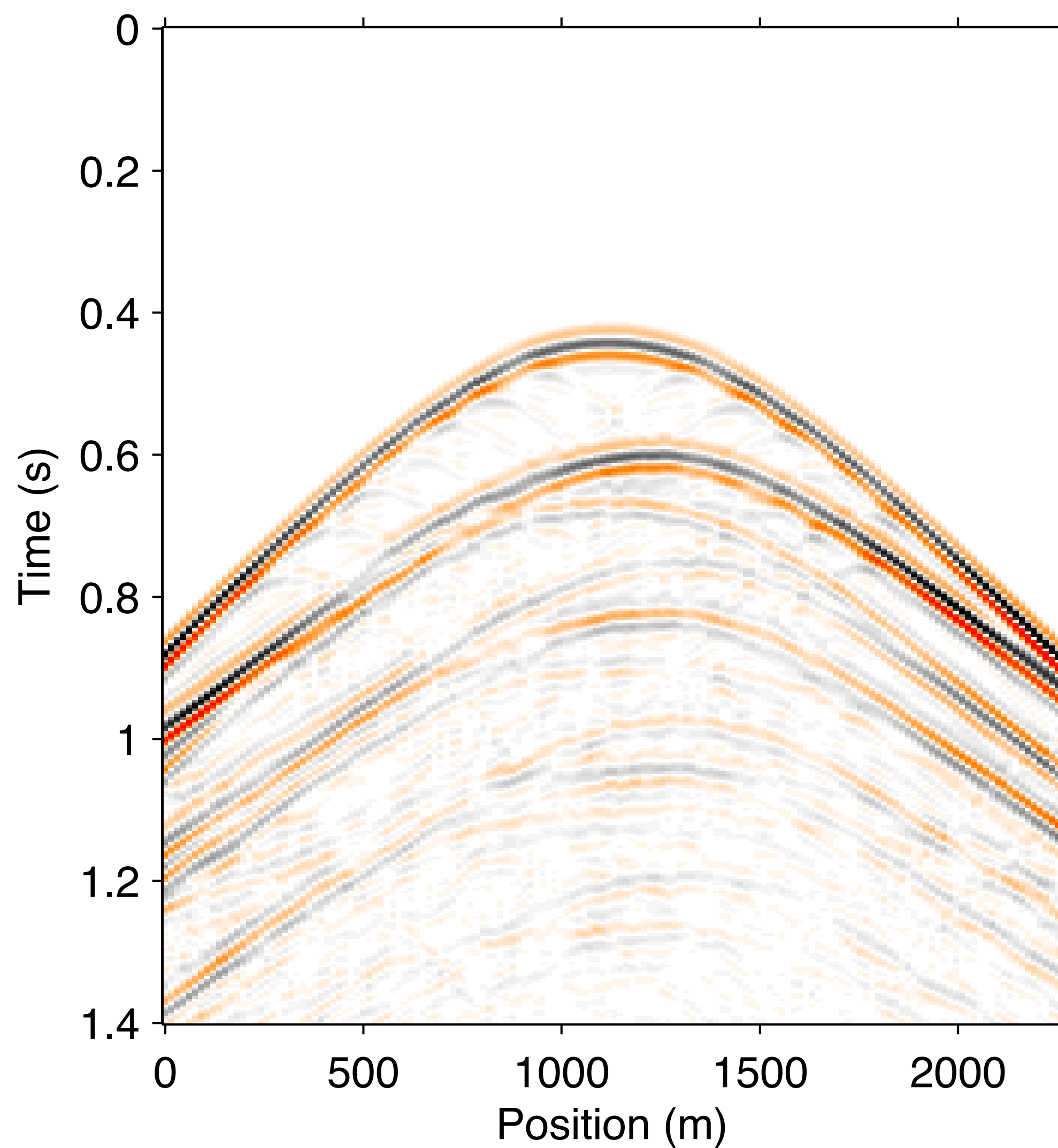


**Direct Primary**  
Solved with spatial sampling  
continuation  
 $dx = 60\text{m} > 30\text{m} > 15\text{m}$



# Warm-starting/continuation from coarse solution

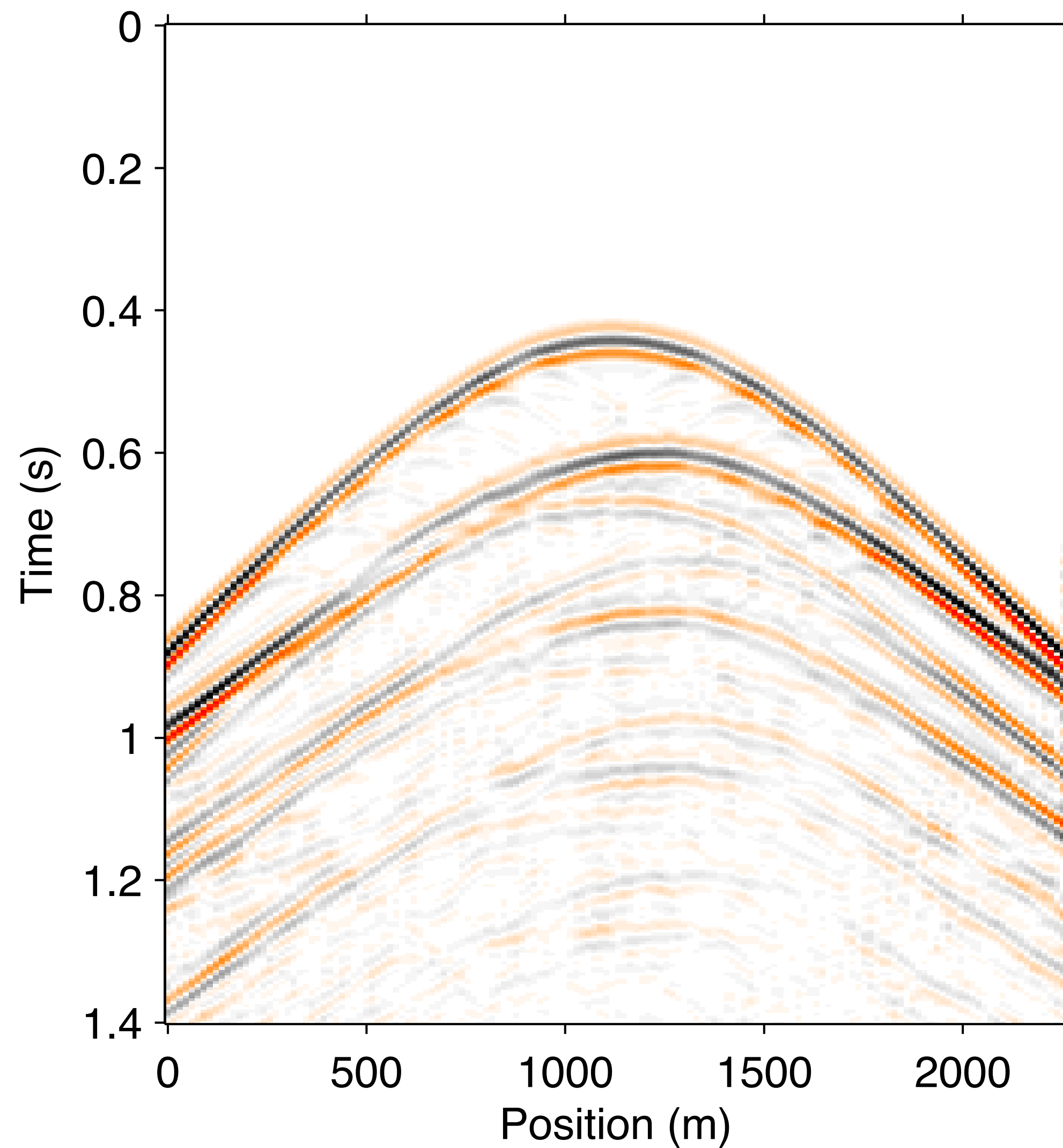
## Example



**REPSI Multiples**  
**Solved from full data**

# Warm-starting/continuation from coarse solution

## Example



**REPSI Multiples**  
Solved with spatial sampling  
continuation  
 $dx = 60m > 30m > 15m$

## Significant speedup from bootstrapping

Per-iteration FLOPs cost (one forward/adjoint):  $n = n_{\text{rcv}} = n_{\text{src}}$

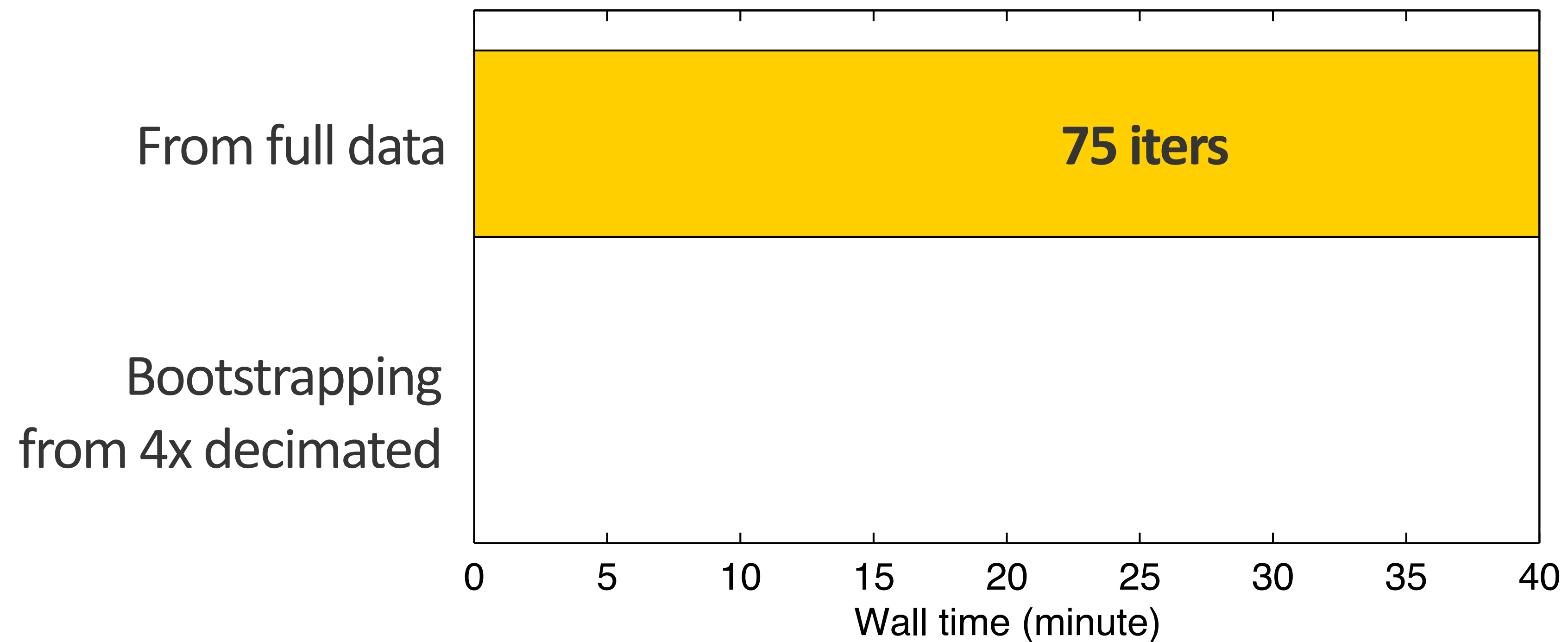
$$\text{Cost}(n) = \underbrace{\mathcal{O}(2n_t \log n_t n^2)}_{\text{2 times FFT}} + \underbrace{\mathcal{O}(n_t n^3)}_{\text{computing MCG \& sum in FX}}$$

$$\text{Cost}\left(\frac{1}{2}n\right) = \frac{1}{4}\mathcal{O}(2n_t \log n_t n^2) + \frac{1}{8}\mathcal{O}(n_t n^3)$$

$$\text{Cost}\left(\frac{1}{4}n\right) = \frac{1}{16}\mathcal{O}(2n_t \log n_t n^2) + \frac{1}{64}\mathcal{O}(n_t n^3)$$

# Significant speedup from bootstrapping

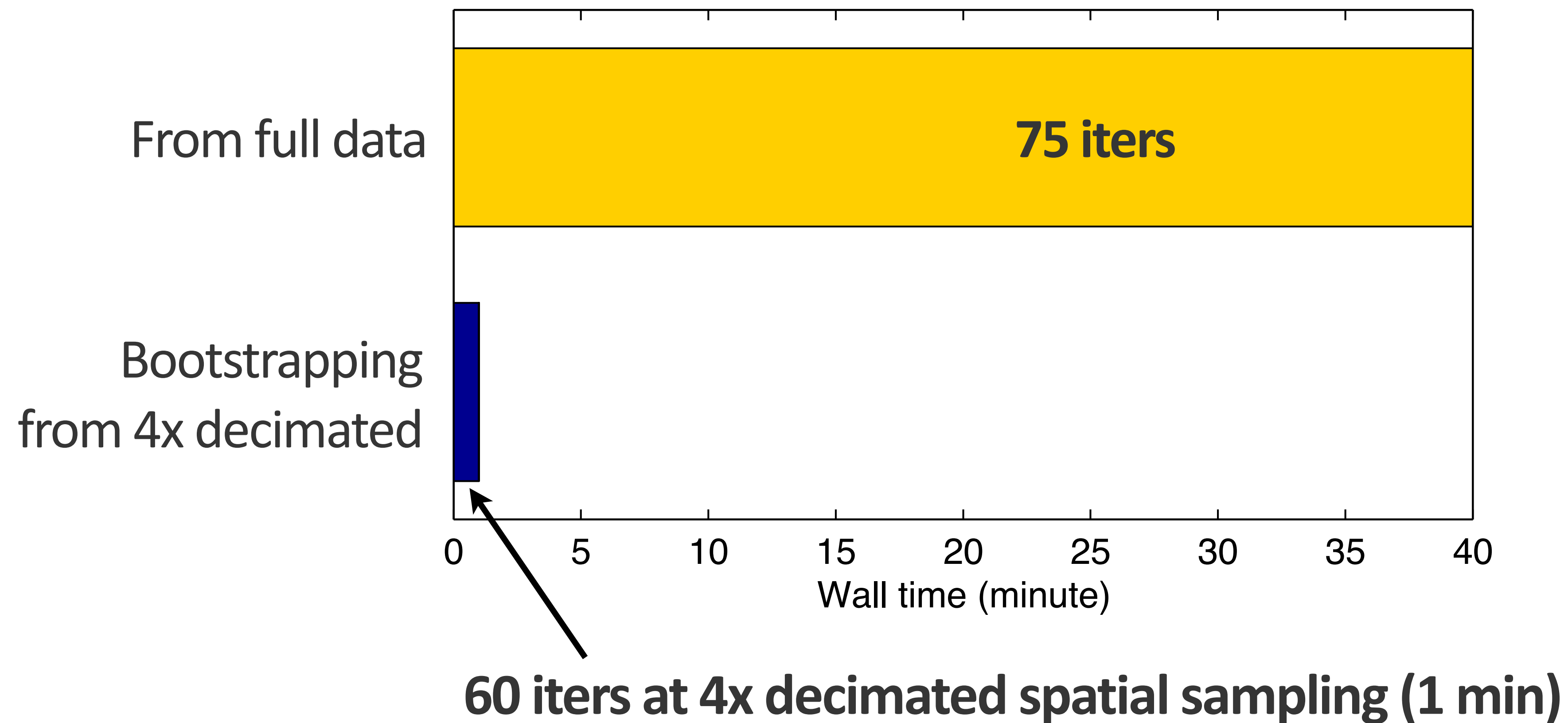
Wall times





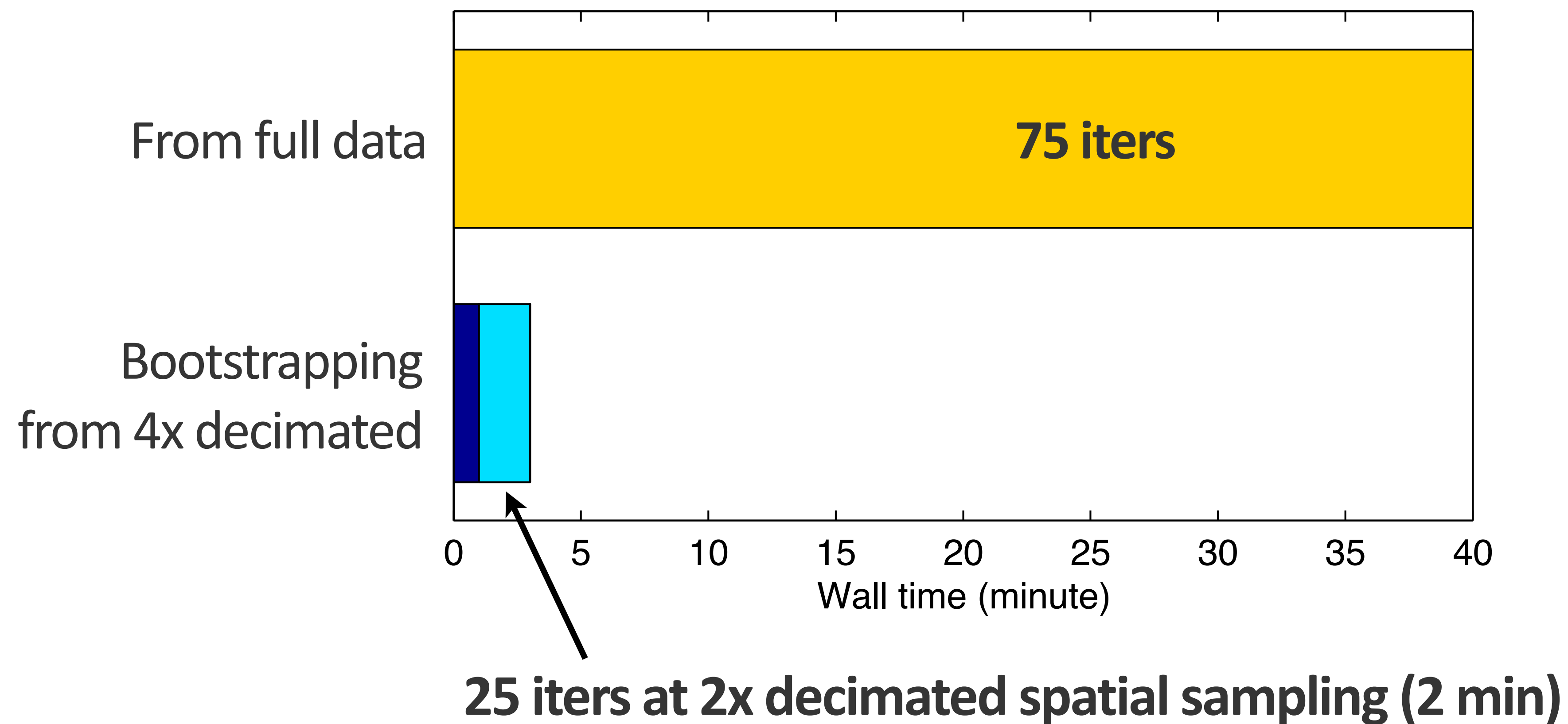
# Significant speedup from bootstrapping

Wall times



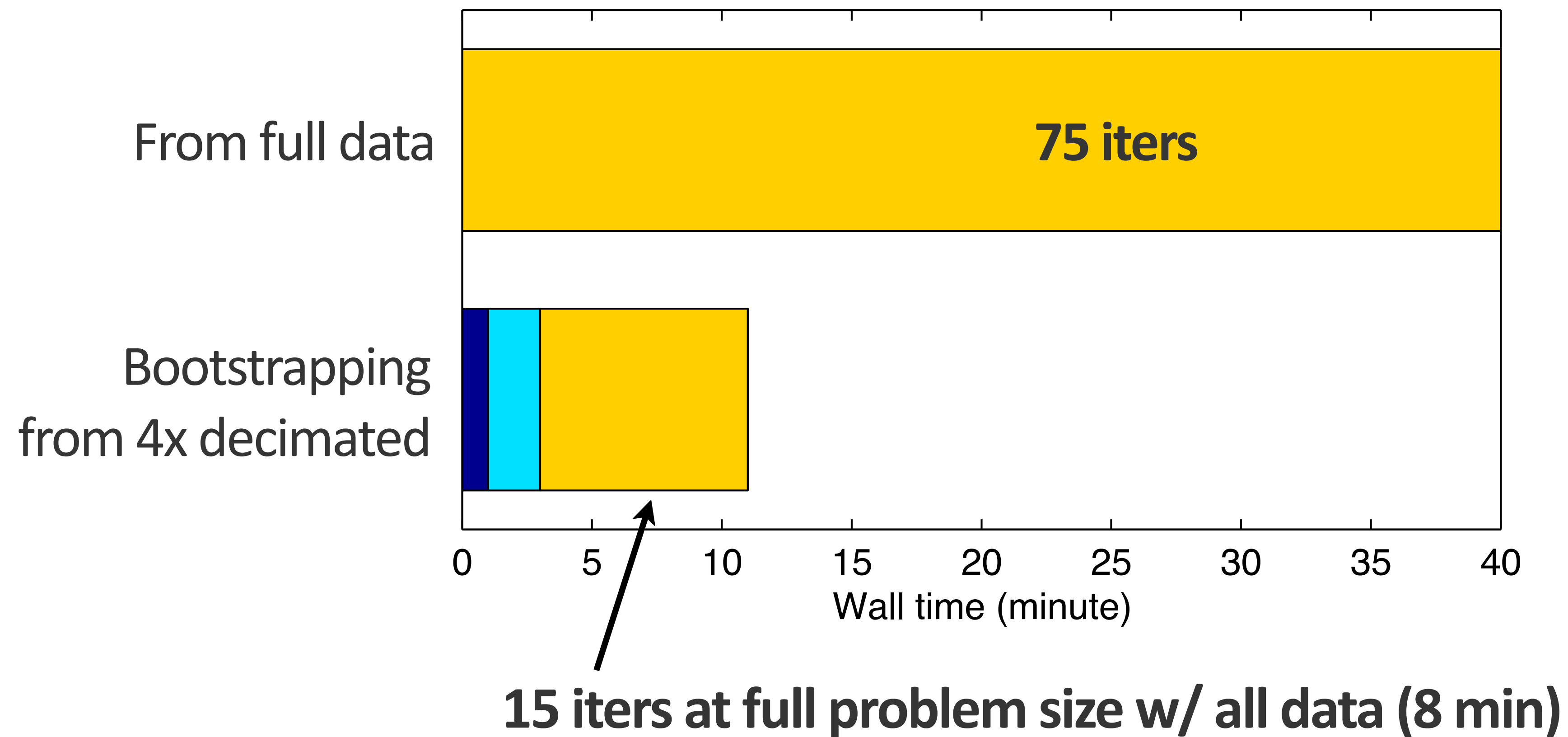
# Significant speedup from bootstrapping

Wall times



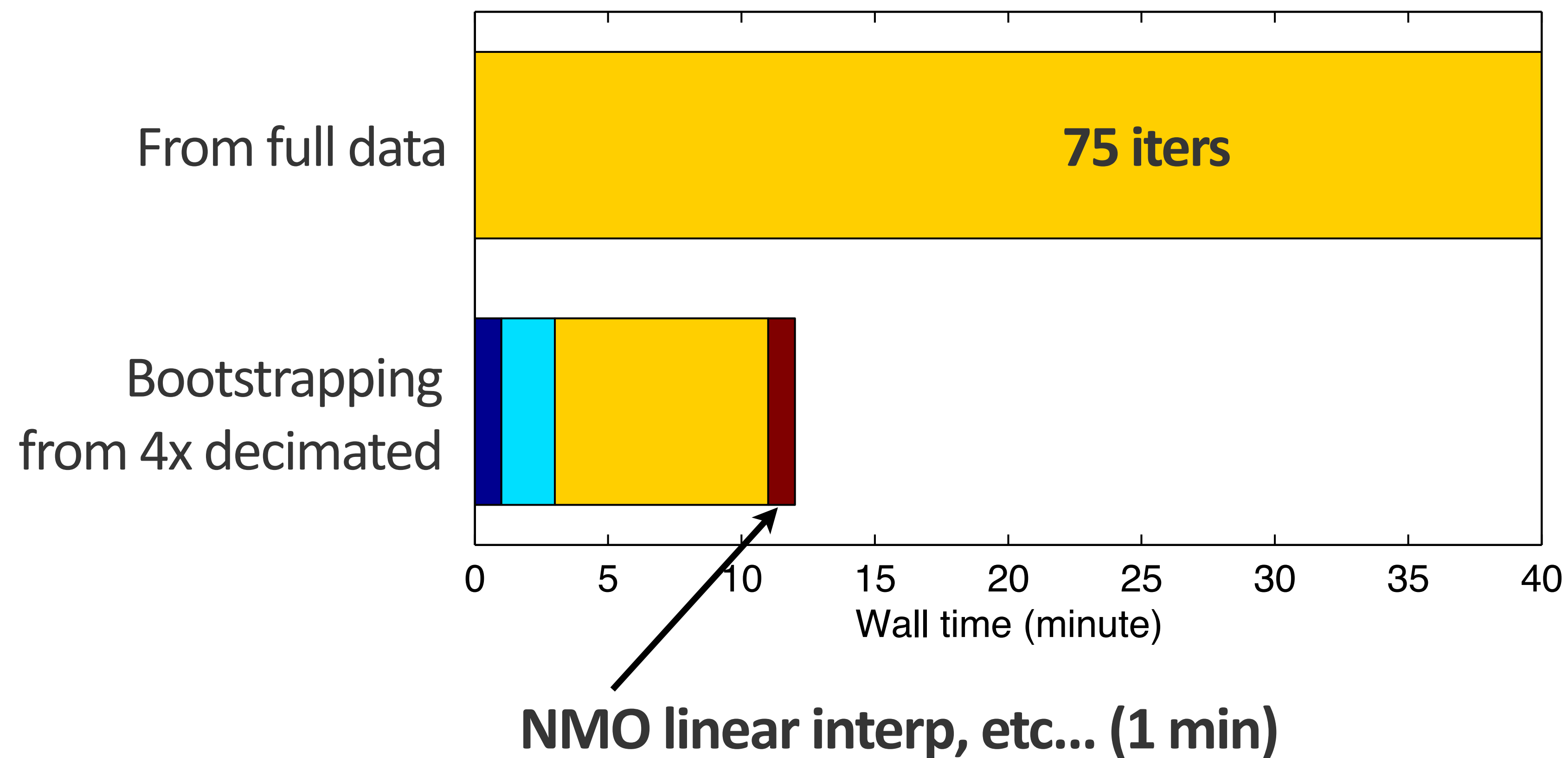
# Significant speedup from bootstrapping

Wall times



# Significant speedup from bootstrapping

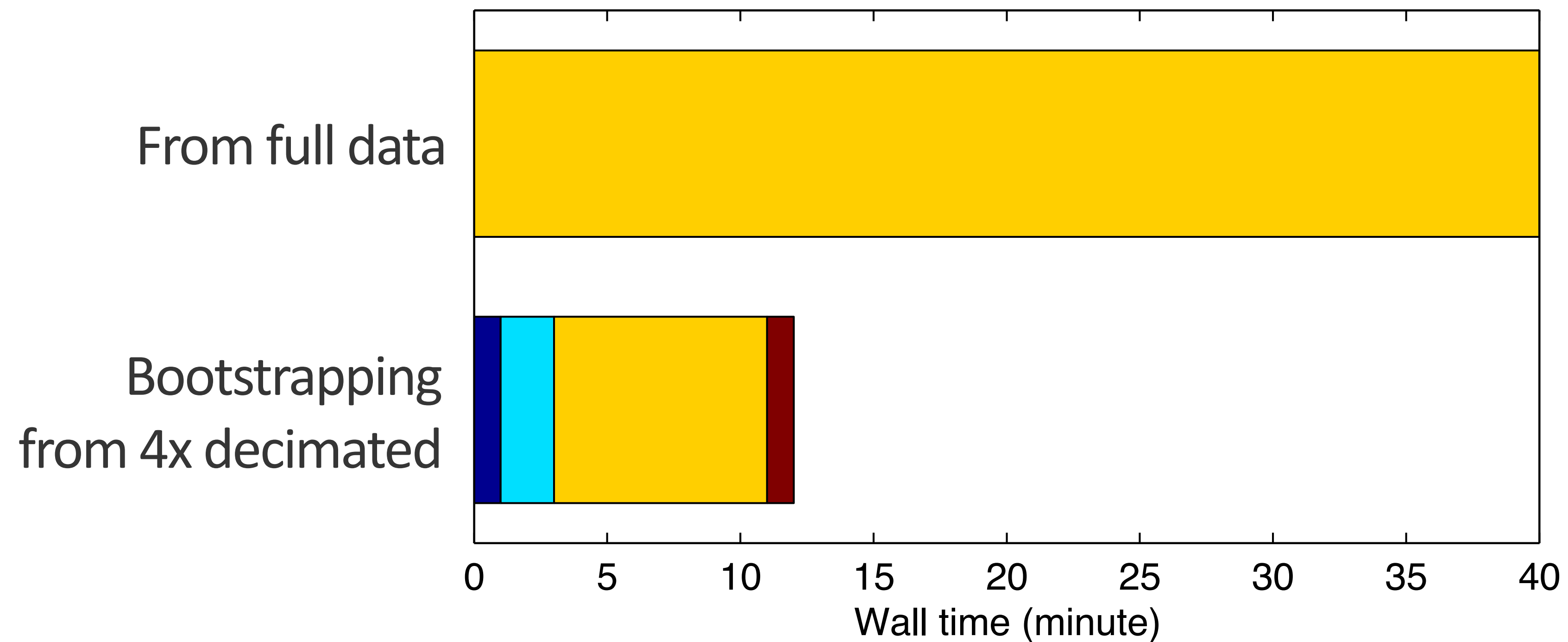
Wall times





# Significant speedup from bootstrapping

Wall times



# **Bootstrapping**

application to under-sampled data

# Robust EPSI

With updates to unknown data

**While**  $\|\mathbf{p}_k - \mathcal{M}(\mathbf{g}_k, \mathbf{q}_k)\|_2 > \sigma$

determine new  $\tau_k$  from the Pareto curve

$$\mathbf{g}_{k+1} = \arg \min_{\mathbf{g}} \|\mathbf{p}_k - \mathbf{M}_{q_k} \mathbf{g}\|_2 \text{ s.t. } \|\mathbf{g}\|_1 \leq \tau_k$$

(Solve with SPG part of SPGL1 until Pareto curve reached)

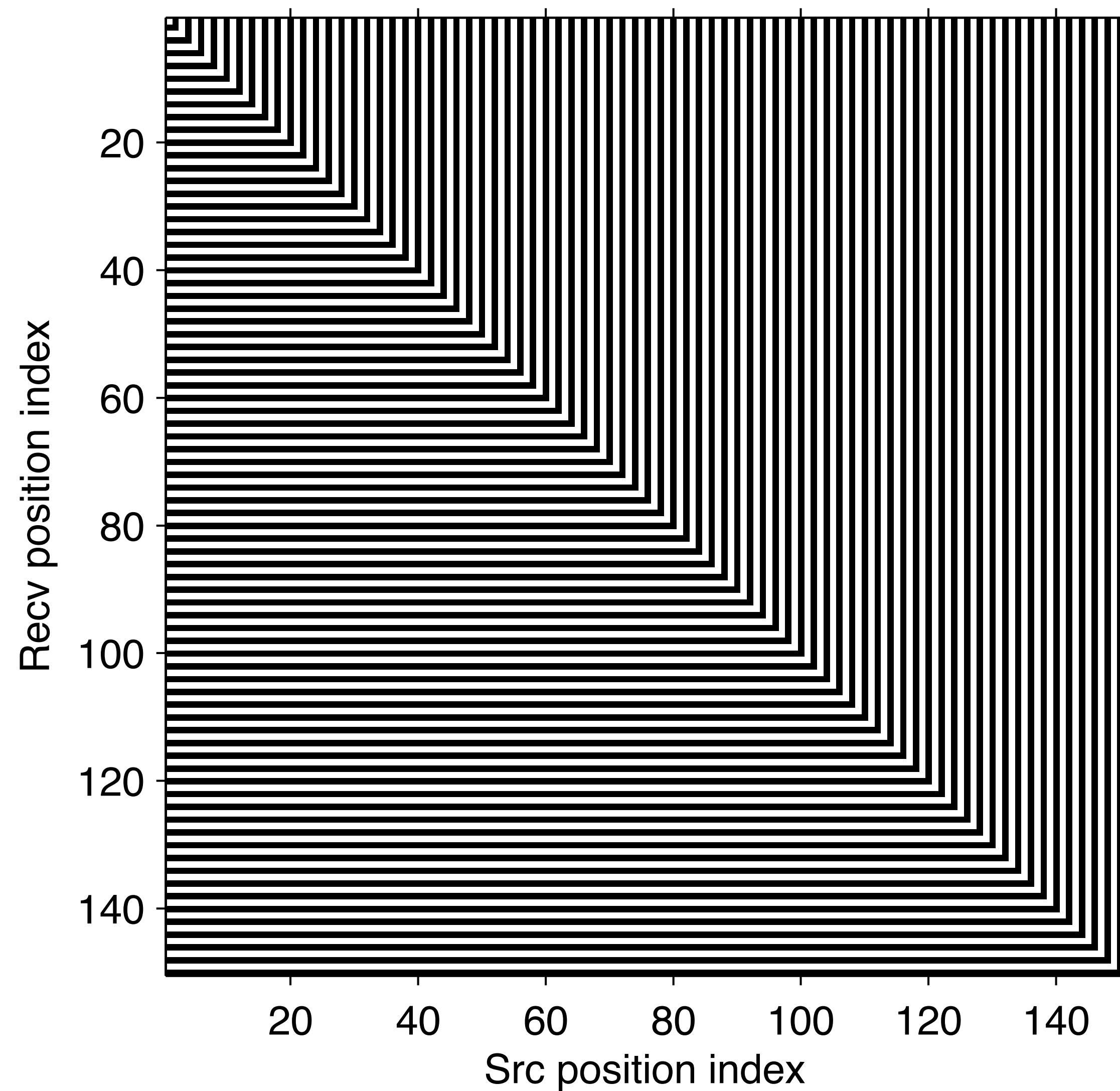
$$\mathbf{q}_{k+1} = \arg \min_{\mathbf{q}} \|\mathbf{p}_k - \mathbf{M}_{g_{k+1}} \mathbf{q}\|_2$$

(Solve with LSQR)

$$\mathbf{p}_{k+1} = \mathbf{p}_k + \alpha \Delta \mathbf{p}(\mathbf{g}_{k+1}, \mathbf{q}_{k+1}, \mathbf{p}_k)$$

(Gradient update on data)

# REPSI with 2:1 source undersampling



Trace mask

dRecv 15m

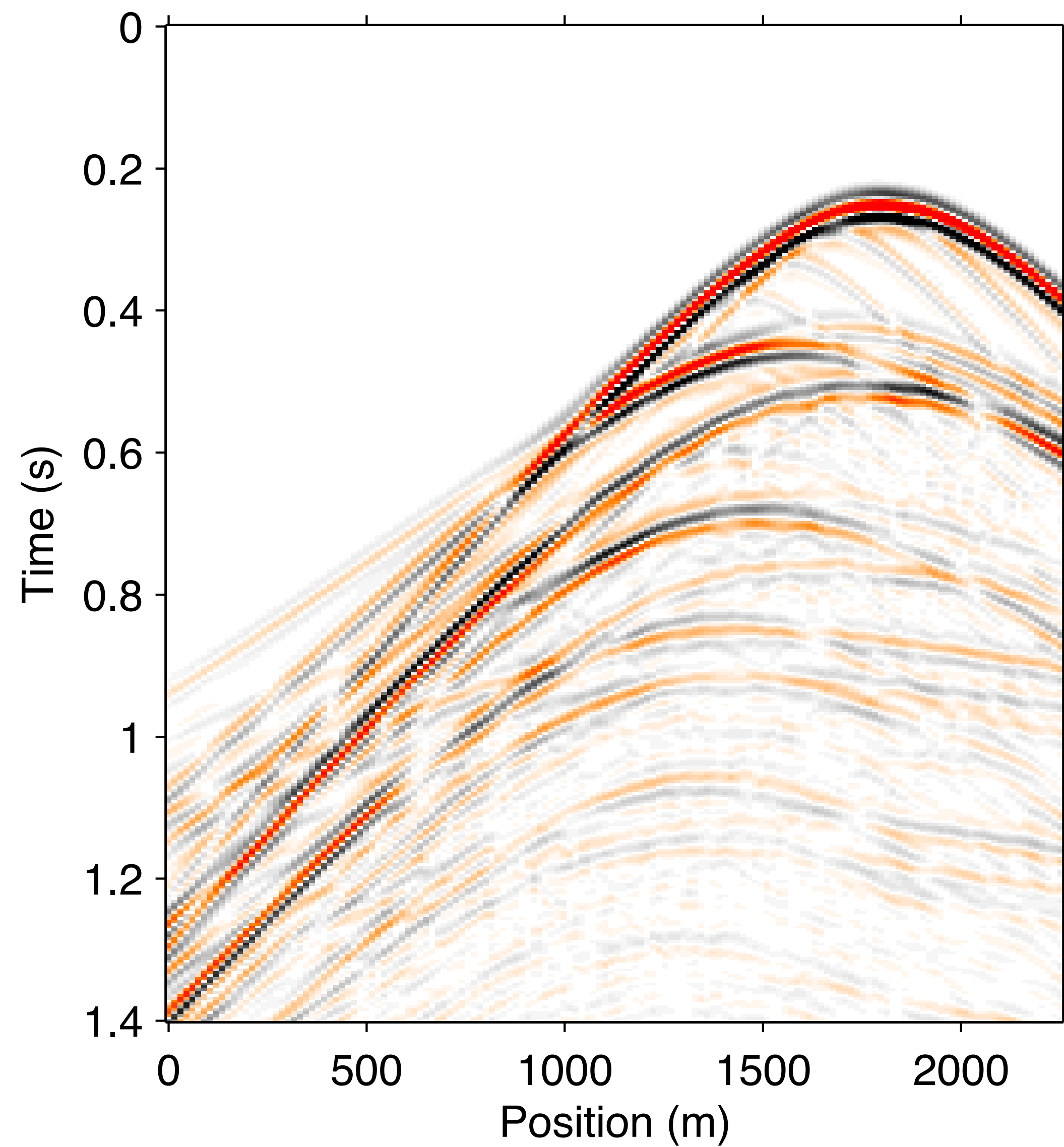
dSrc 30m

(assuming streamer acquisition)

■ : missing data

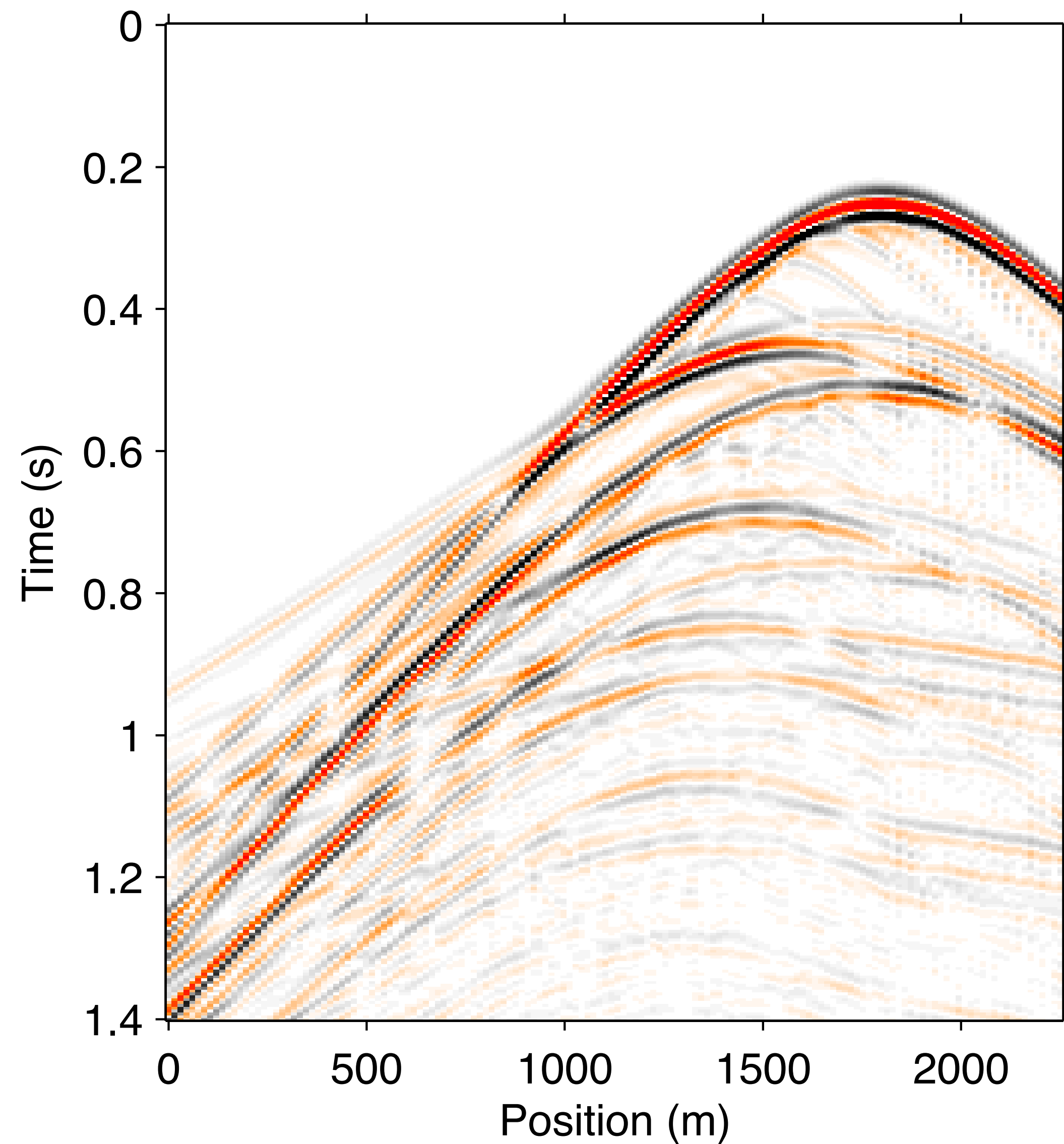


# REPSI with 2:1 source undersampling



**Fully-sampled data**  
shot gather  
src at 1800m

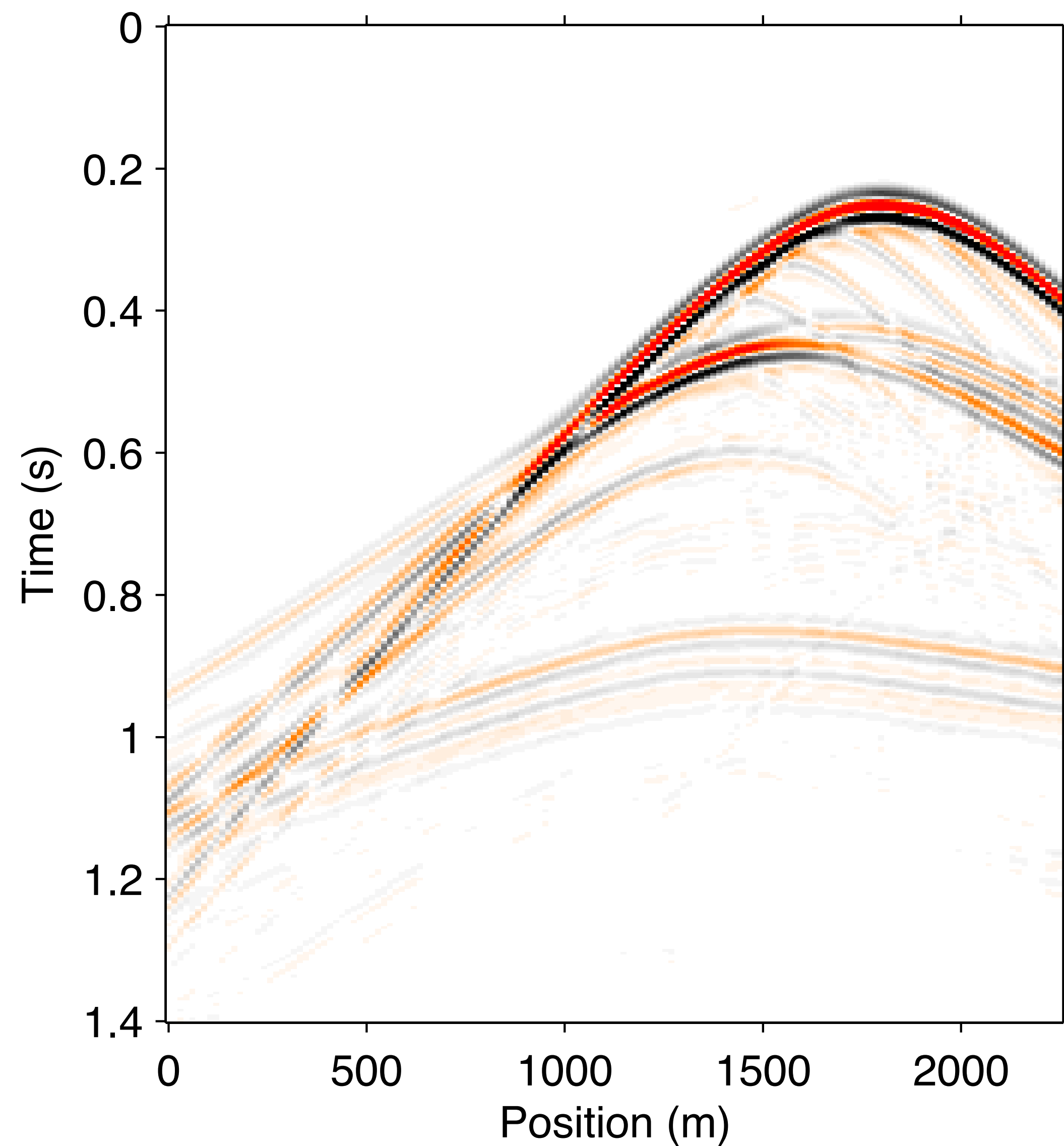
# REPSI with 2:1 source undersampling



## Interpolated data

via const NMO 1600m/s  
natural-neighbor trace copy  
negative offsets sampled  
src at 1800m

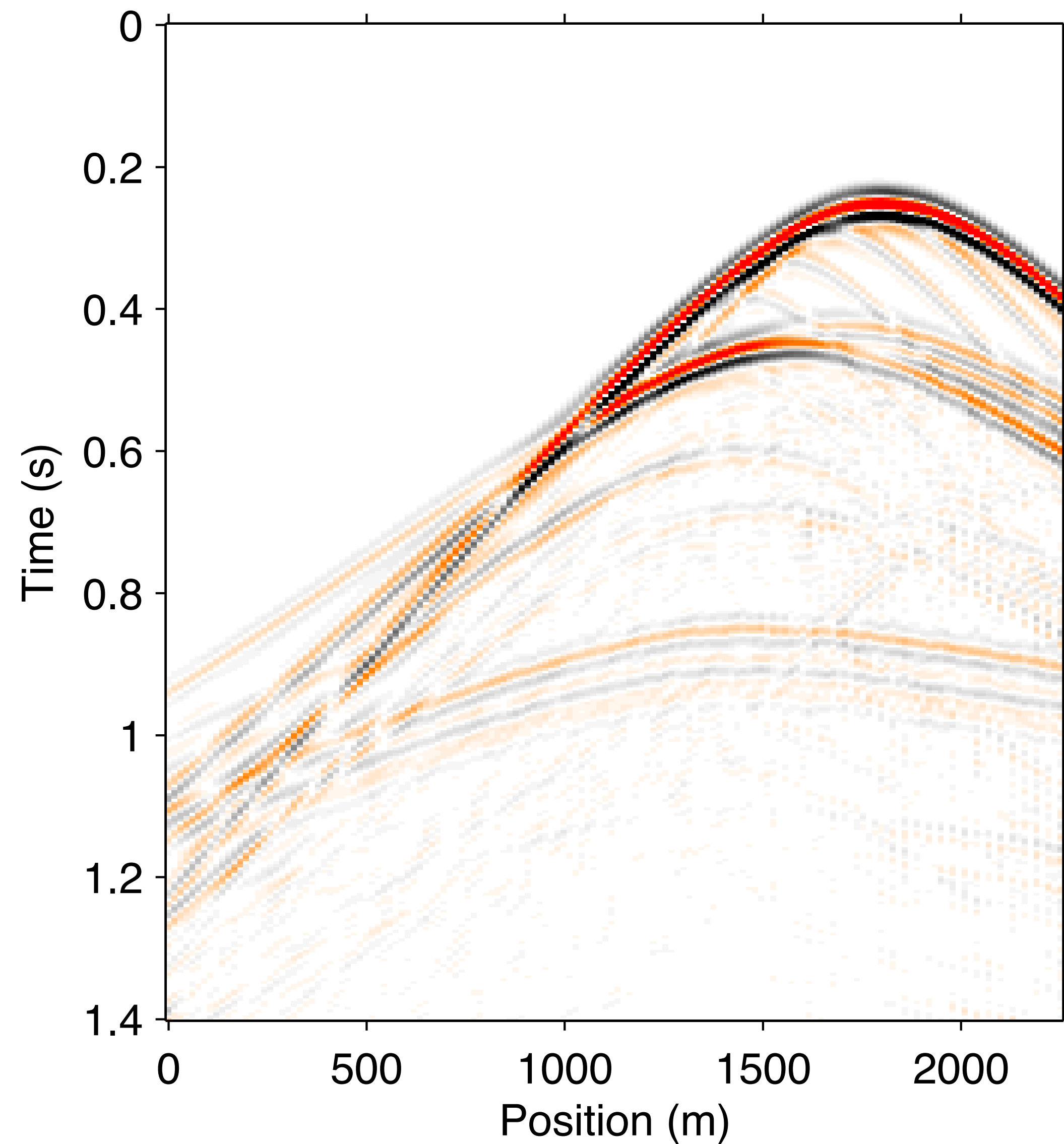
# REPSI with 2:1 source undersampling



## Reference solution

REPSI from fully-sampled data  
(conservative primary)

# REPSI with 2:1 source undersampling

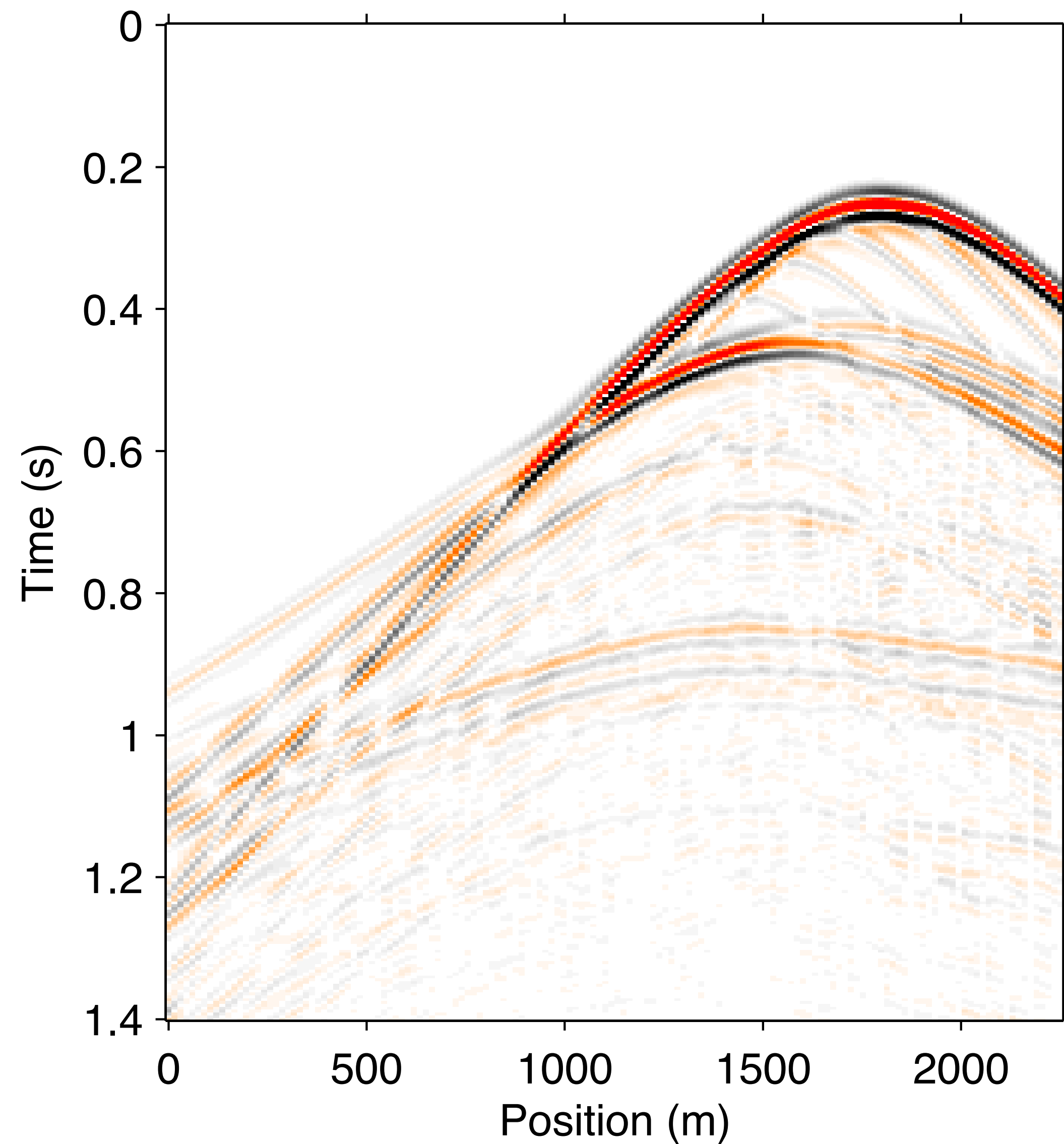


## REPSI primary

from 2:1 source undersampling  
with data updates  
 $d\text{Recv} = 15\text{m}$ ,  $d\text{Src} = 30\text{m}$   
(conservative primary)



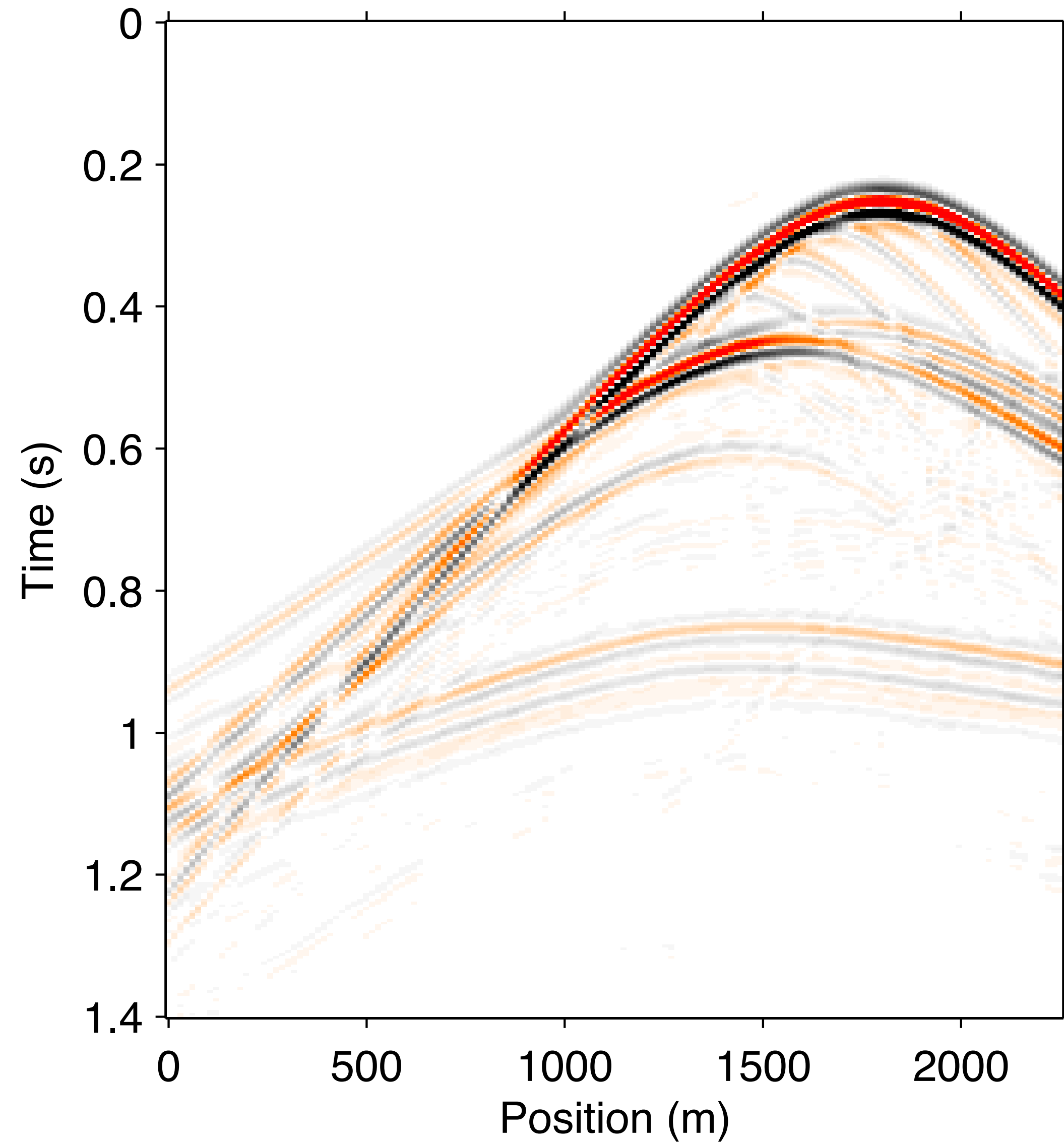
# REPSI with 4:1 source undersampling



## REPSI primary

from 4:1 source undersampling  
with data updates  
 $d\text{Recv} = 15\text{m}$ ,  $d\text{Src} = 60\text{m}$   
(conservative primary)

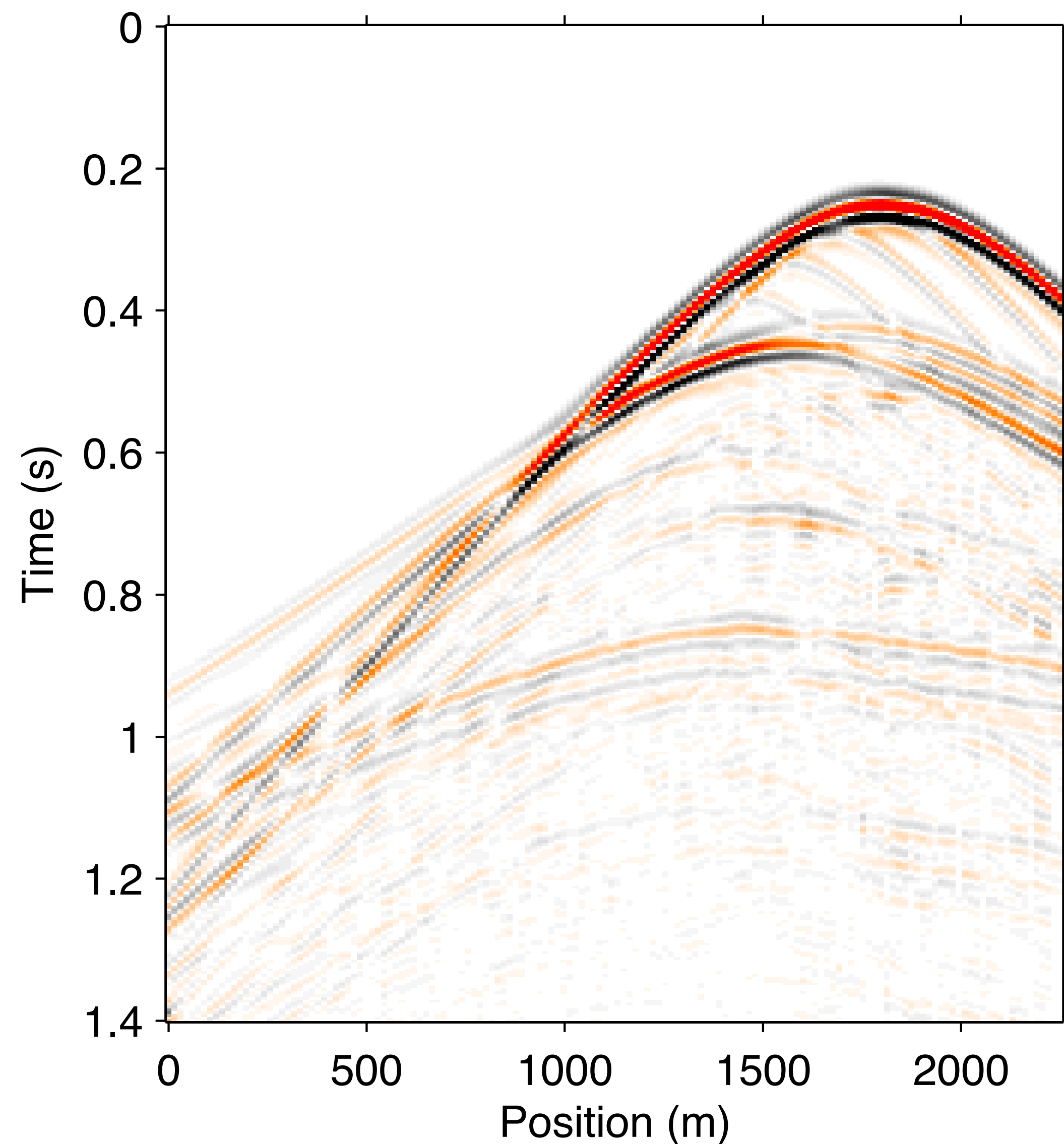
# REPSI with 2:1 source undersampling



## Reference solution

REPSI from fully-sampled data  
(conservative primary)

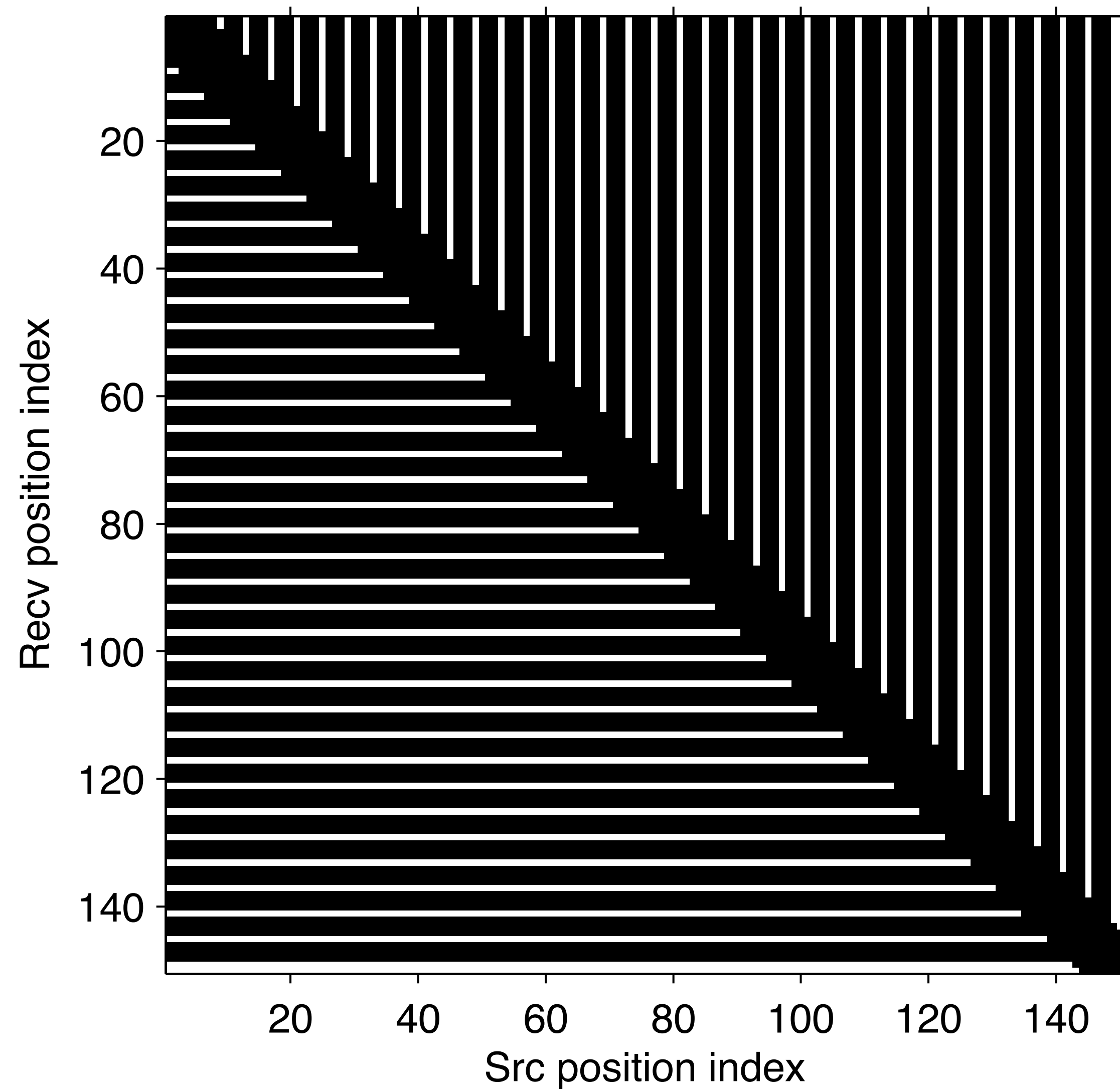
# REPSI with 4:1 source, nearest offset at 105m



## REPSI primary

from 4:1 source undersampling  
nearest offset at 105m  
with data updates  
 $d\text{Recv} = 15\text{m}$ ,  $d\text{Src} = 60\text{m}$   
(conservative primary)

# REPSI with 4:1 source, nearest offset at 105m



Trace mask

dRecv 15m

dSrc 60m

nearest offset 105m

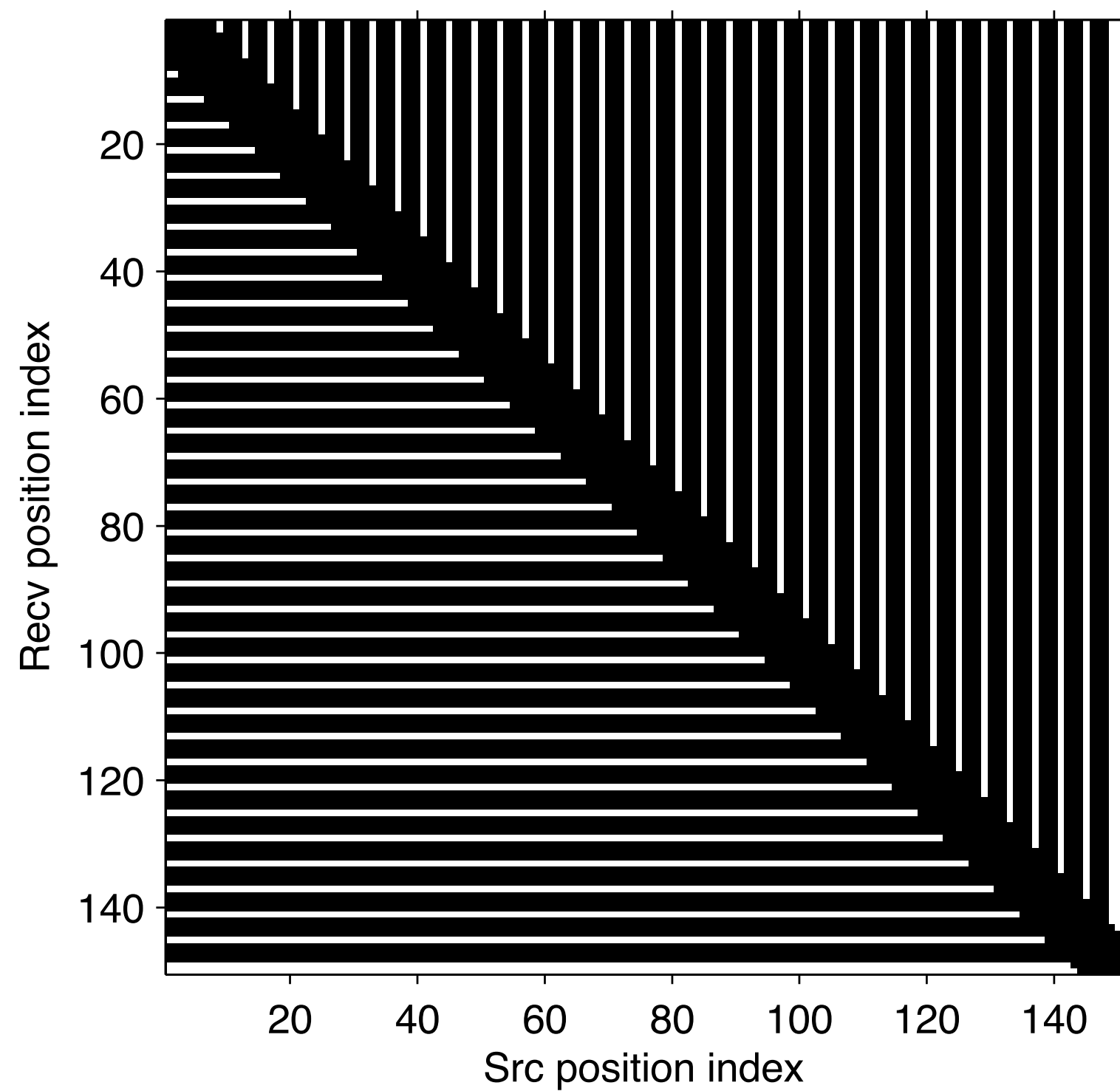
(assuming streamer acquisition)

■ : missing data

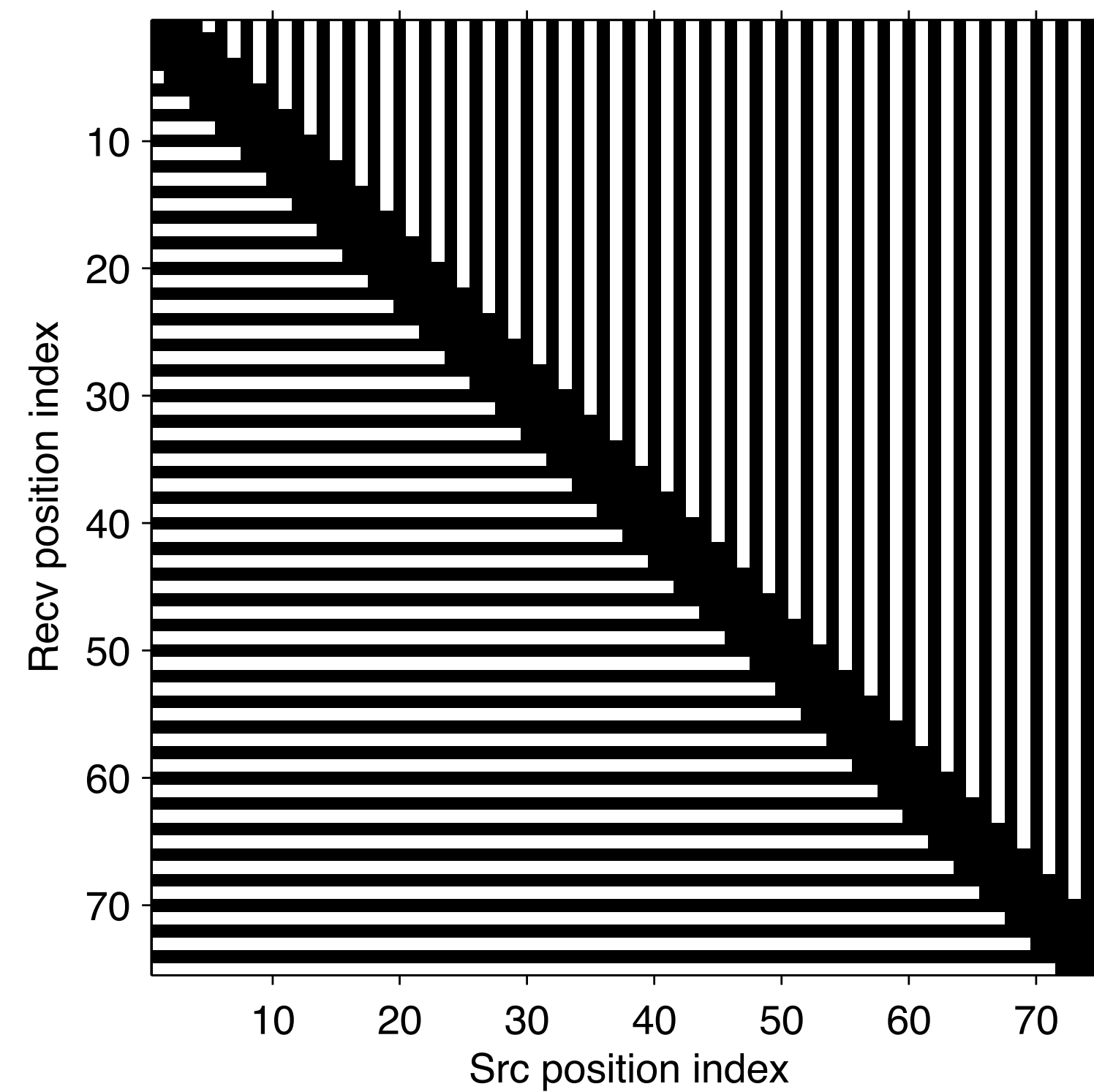


# Bootstrapping, for unknown data

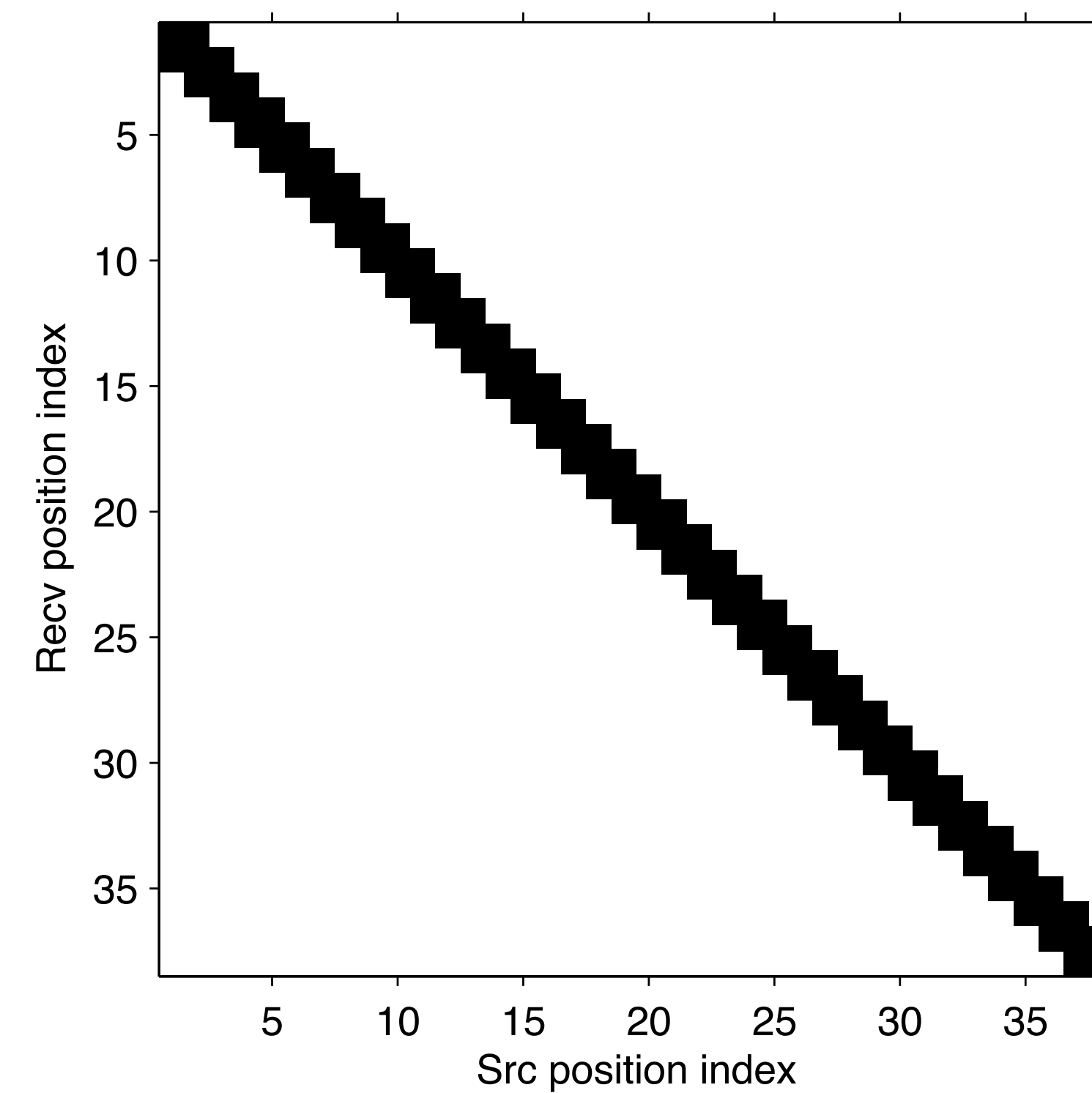
Mask at  $dx=15m$



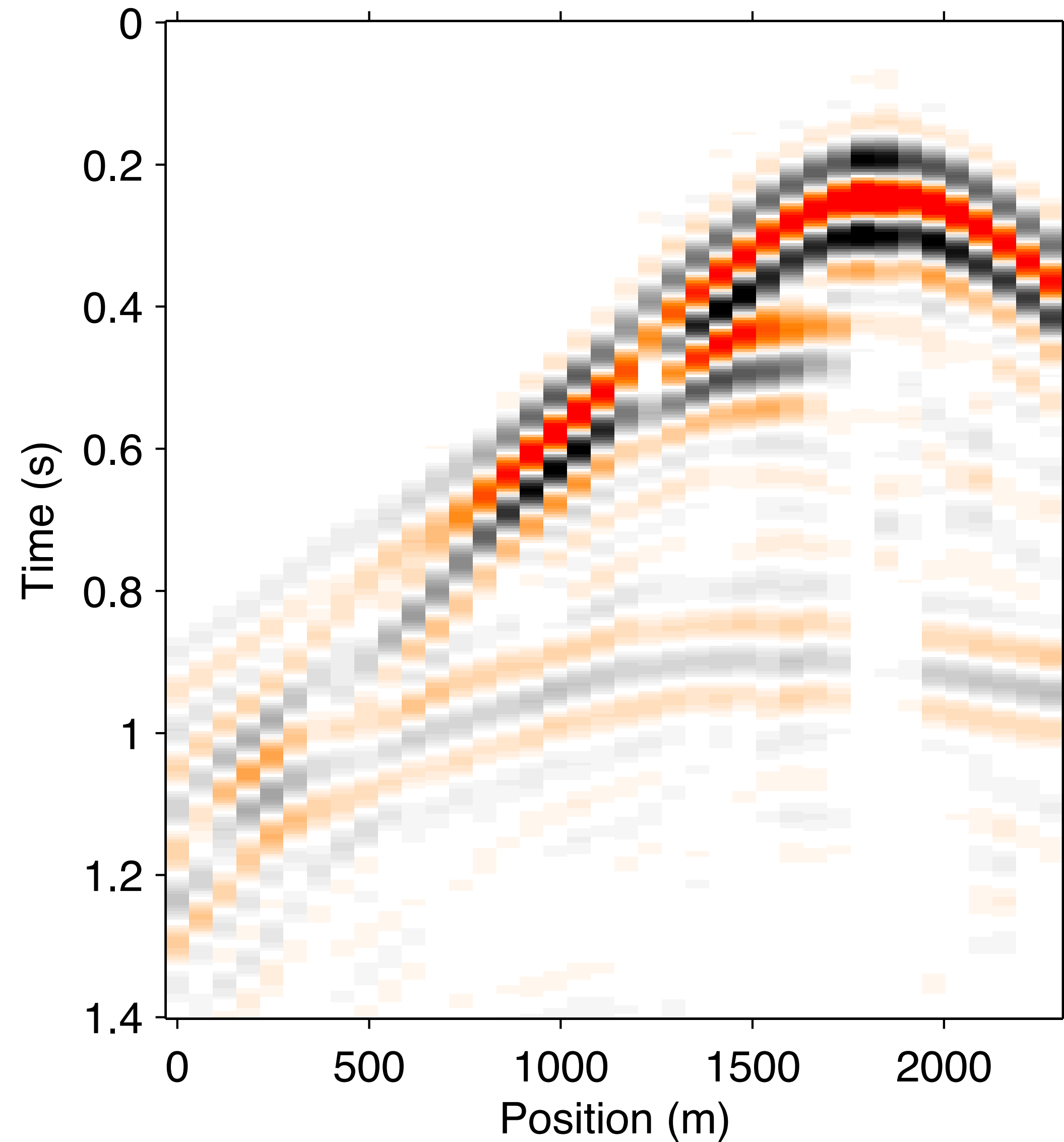
Mask at  $dx=30m$



Mask at  $dx=60m$



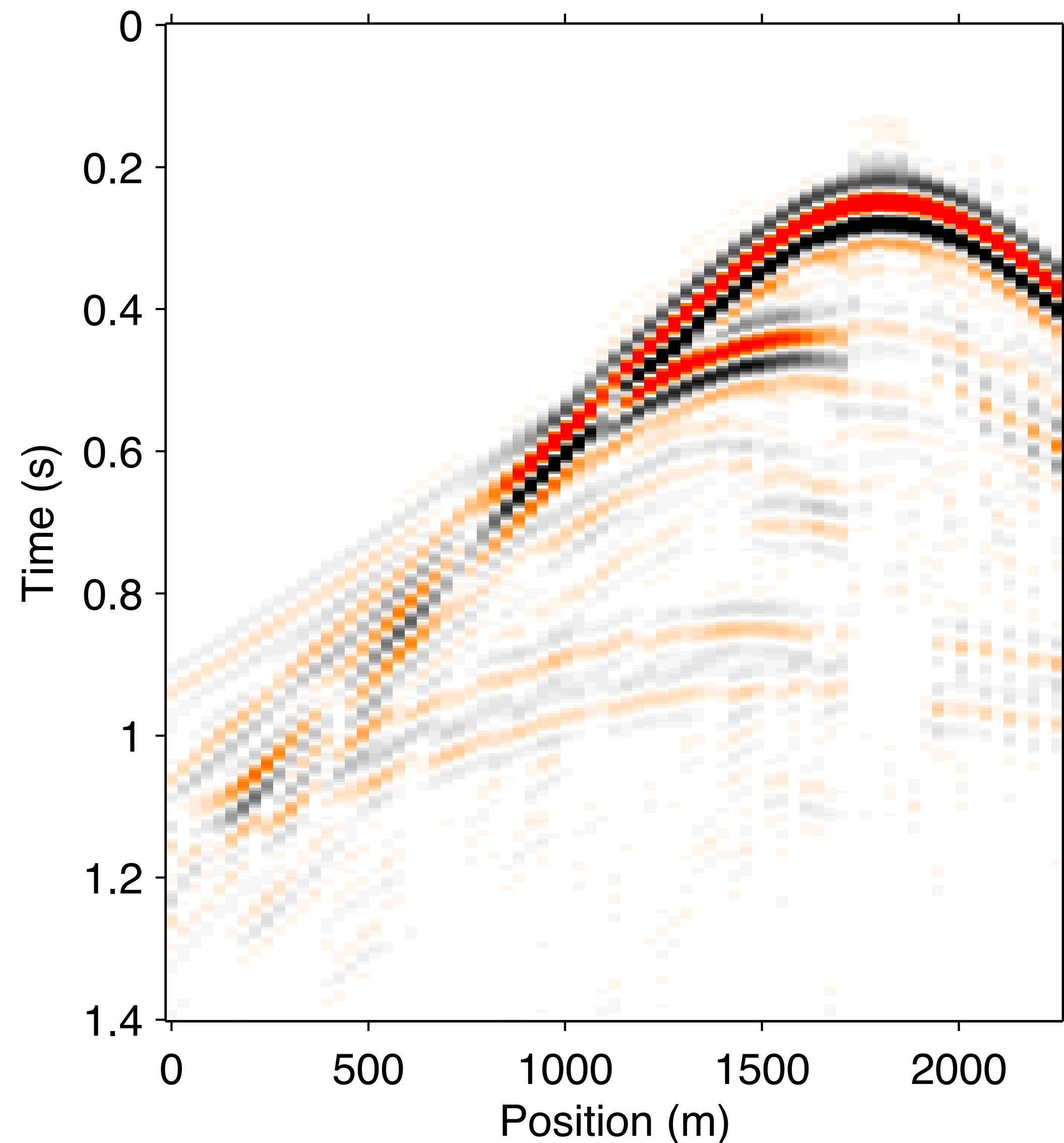
# Bootstrap REPSI with 4:1 source, nearest offset at 105m



## REPSI primary

from 4:1 source undersampling  
nearest offset at 105m  
with data updates  
 $d_{Recv} = 60m$ ,  $d_{Src} = 60m$   
(direct primary)

# Bootstrap REPSI with 4:1 source, nearest offset at 105m



## REPSI primary

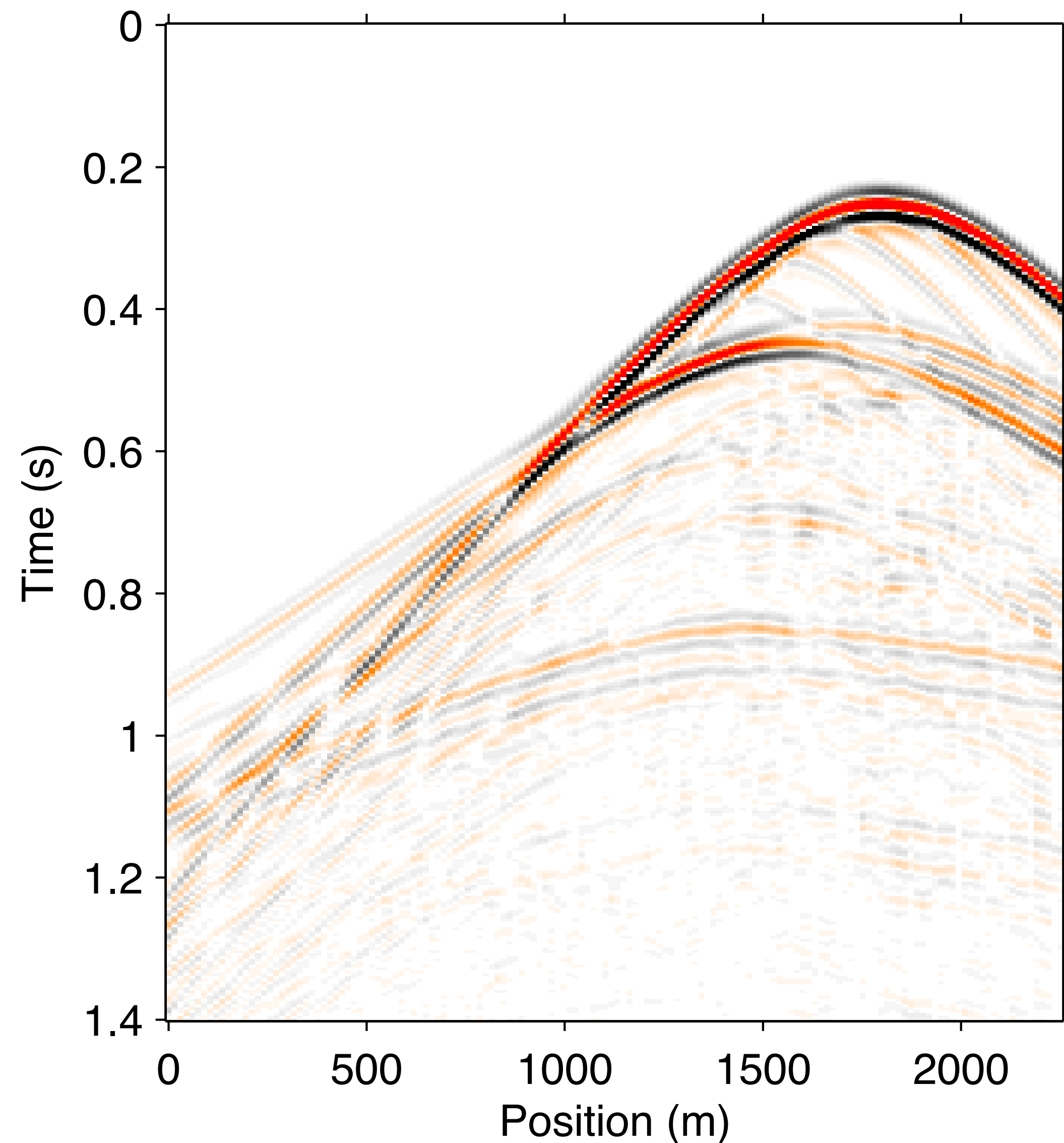
from 4:1 source undersampling  
nearest offset at 105m  
with data updates

**starting from dx=60m solution**

dRecv = 30m, dSrc = 60m

(direct primary)

# Bootstrap REPSI with 4:1 source, nearest offset at 105m



## REPSI primary

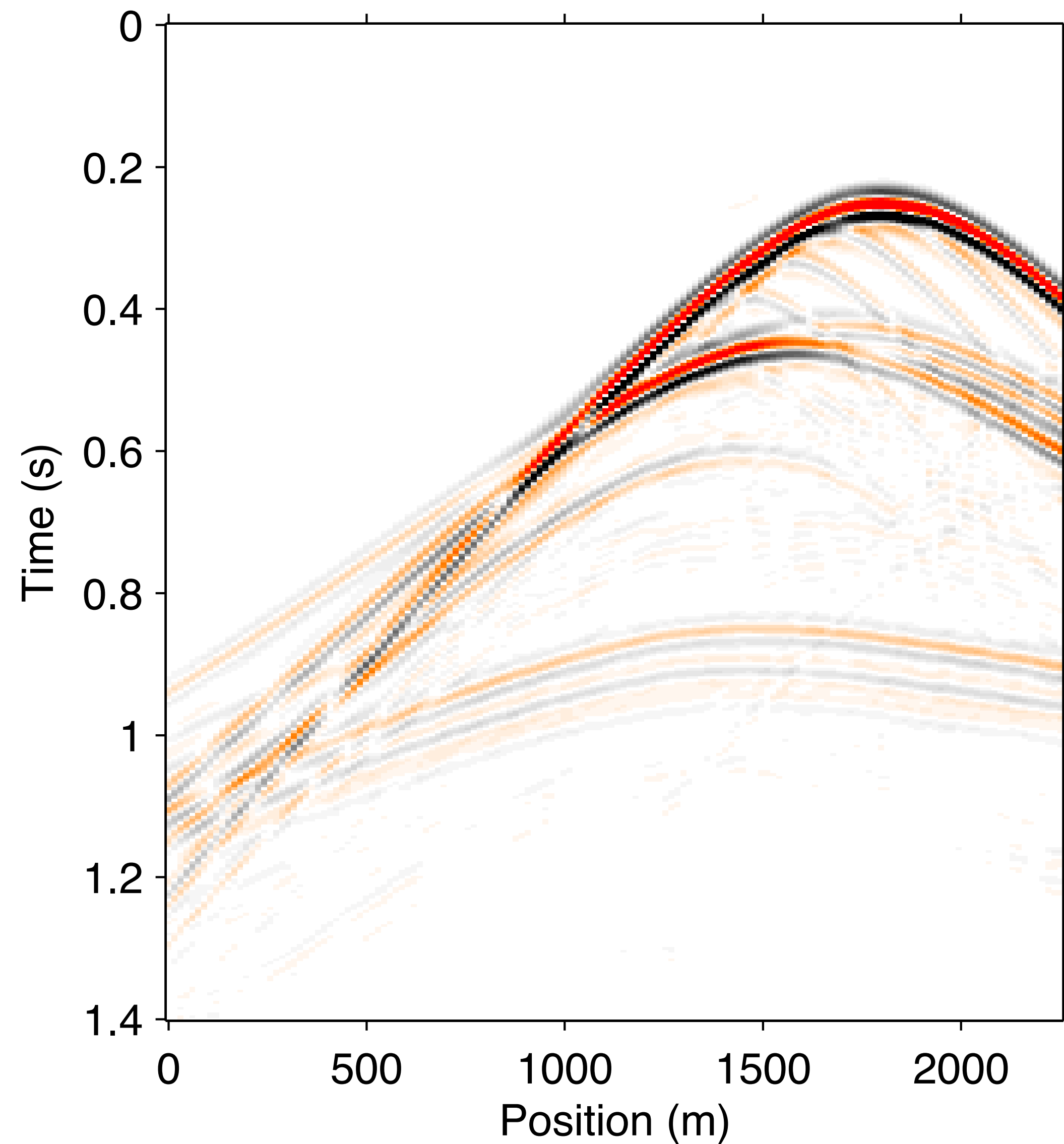
from 4:1 source undersampling  
nearest offset at 105m  
with data updates

**starting from dx=30m solution  
(which started from dx=60m)**

dRecv = 15, dSrc = 60m  
(conservative primary)



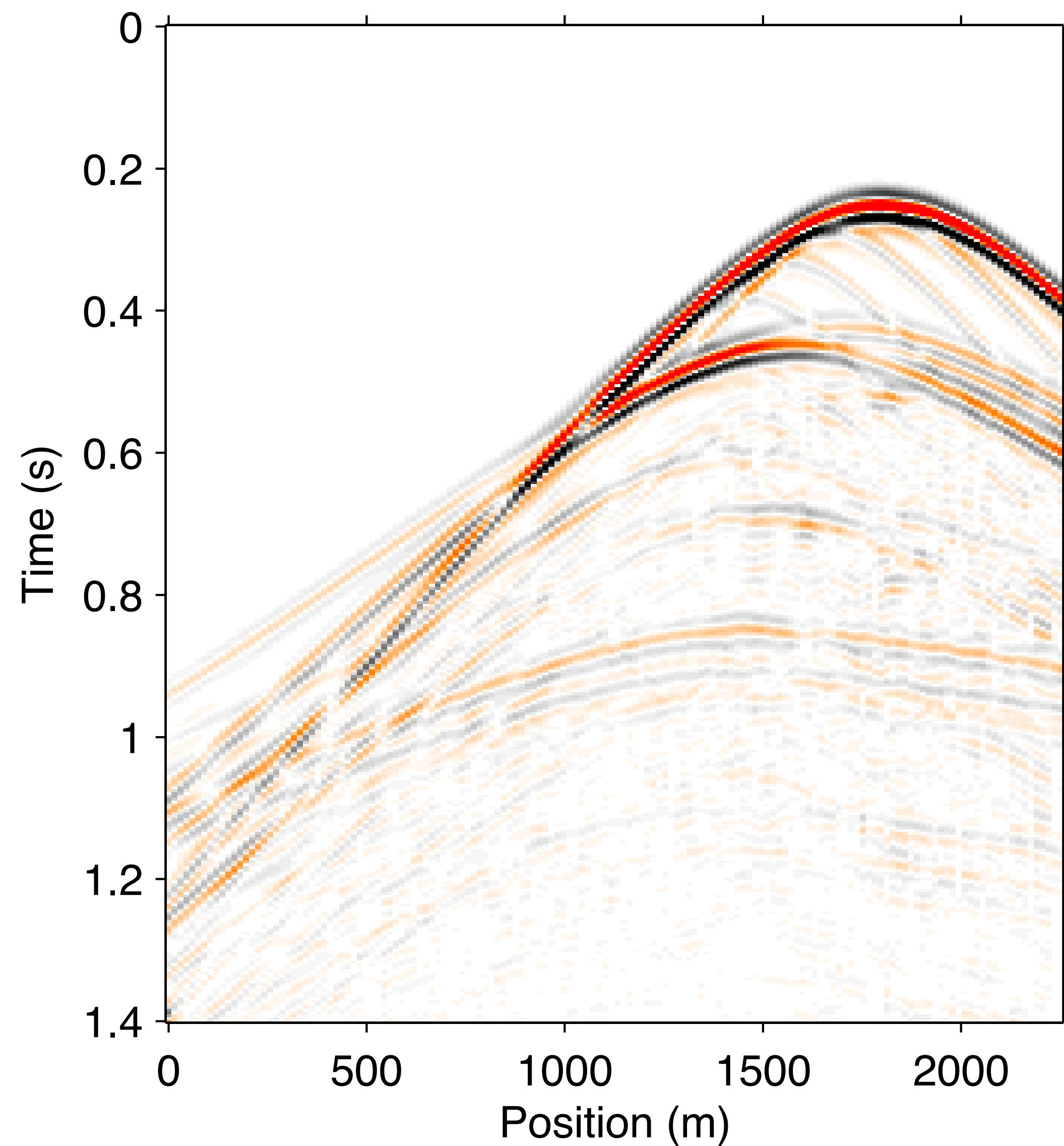
# REPSI with 2:1 source undersampling



## Reference solution

REPSI from fully-sampled data  
(conservative primary)

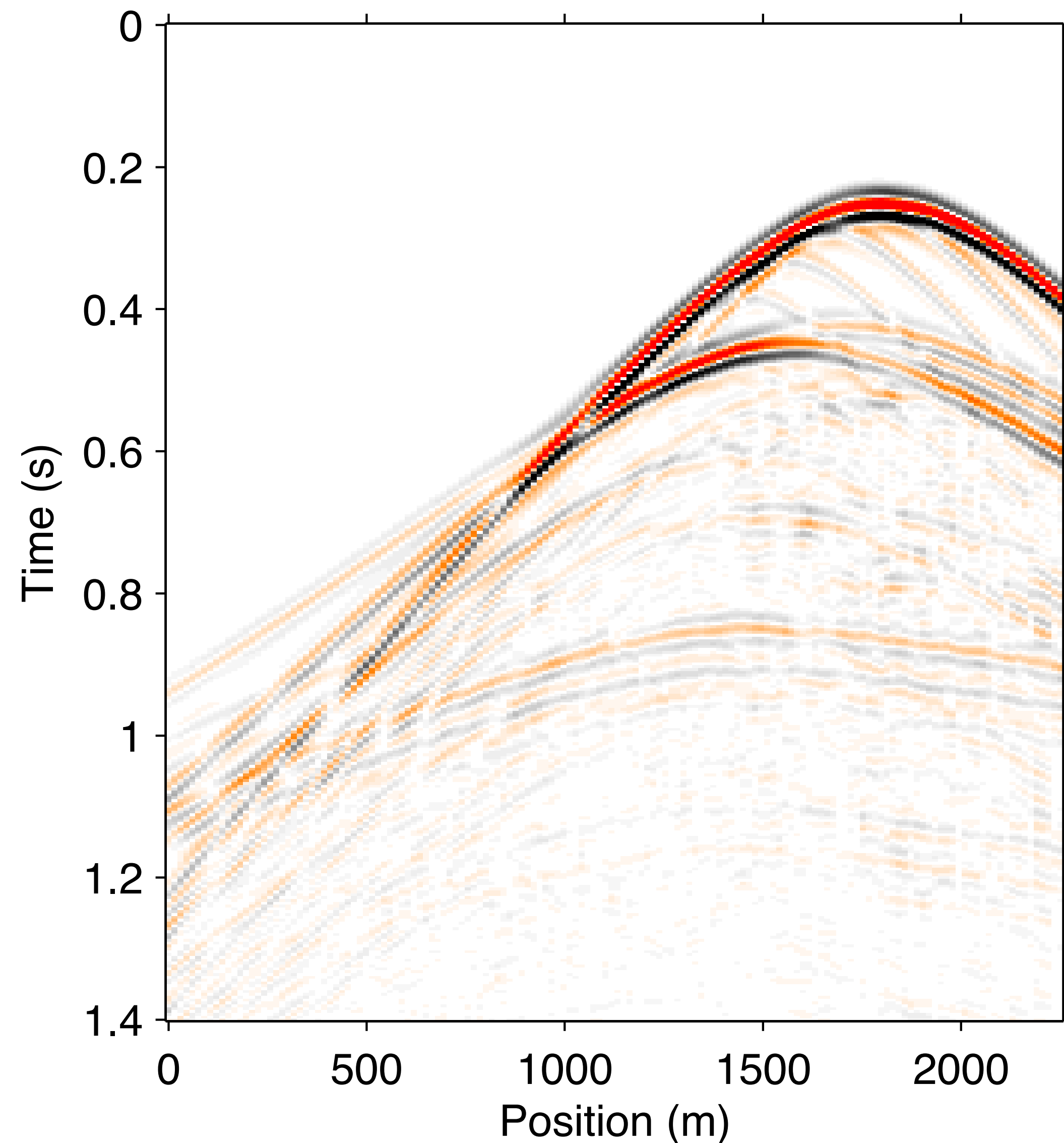
# REPSI with 4:1 source, nearest offset at 105m



## REPSI primary

from 4:1 source undersampling  
nearest offset at 105m  
with data updates  
 $dRecv = 15m$ ,  $dSrc = 60m$   
(conservative primary)

# Bootstrap REPSI with 4:1 source, nearest offset at 105m



## REPSI primary

from 4:1 source undersampling  
nearest offset at 105m  
with data updates

**starting from dx=30m solution  
(which started from dx=60m)**

dRecv = 15, dSrc = 60m  
(conservative primary)

## Sampling-continuation scheme summary

Start REPSI with decimated data, lowpass to avoid spatial aliasing; once “significant” progress is made, continue with less decimated problem

Significant saving in computation cost, 100x to 200x SRMP becomes more like 30x to 40x

Can keep ratio of unknown data at a controlled level

How low can we go? Two limits:

- Coarsest sampling interval in your datatype (crossline, OBN spacing, etc)
- Some lower-bound on feasible low-pass frequency, either from theory, data quality, or geophysical reasons (under investigation)



## Remaining areas of investigation for REPSI

At coarsest levels, use more advanced/costly sparsifying methods? Go grid-free in the time domain? (*i.e., super-resolution methods*)

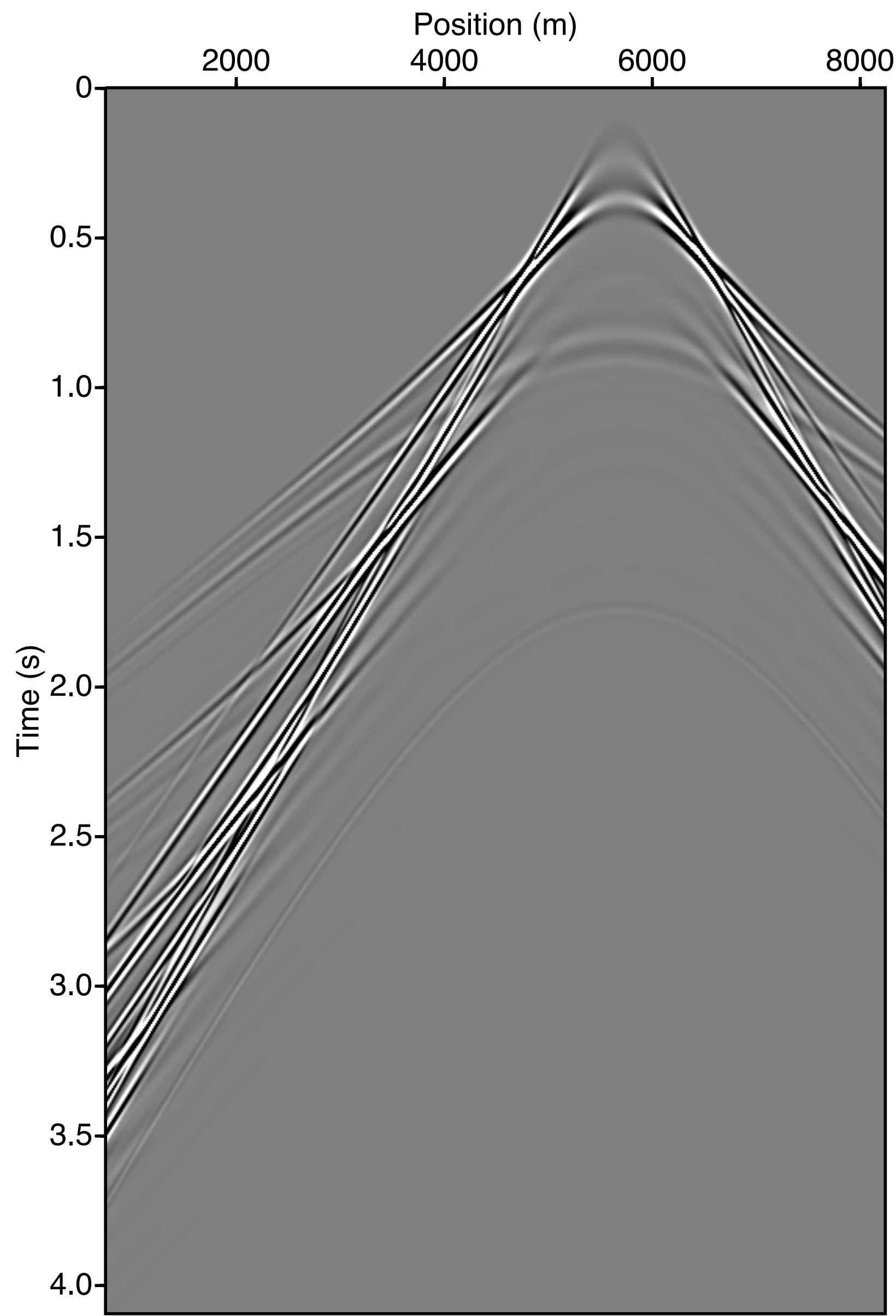
More sophisticated data-update method, less prone to local minima (use correlation between P and G, etc)

Incorporate up/down decomposition operator to work on P & Vz data

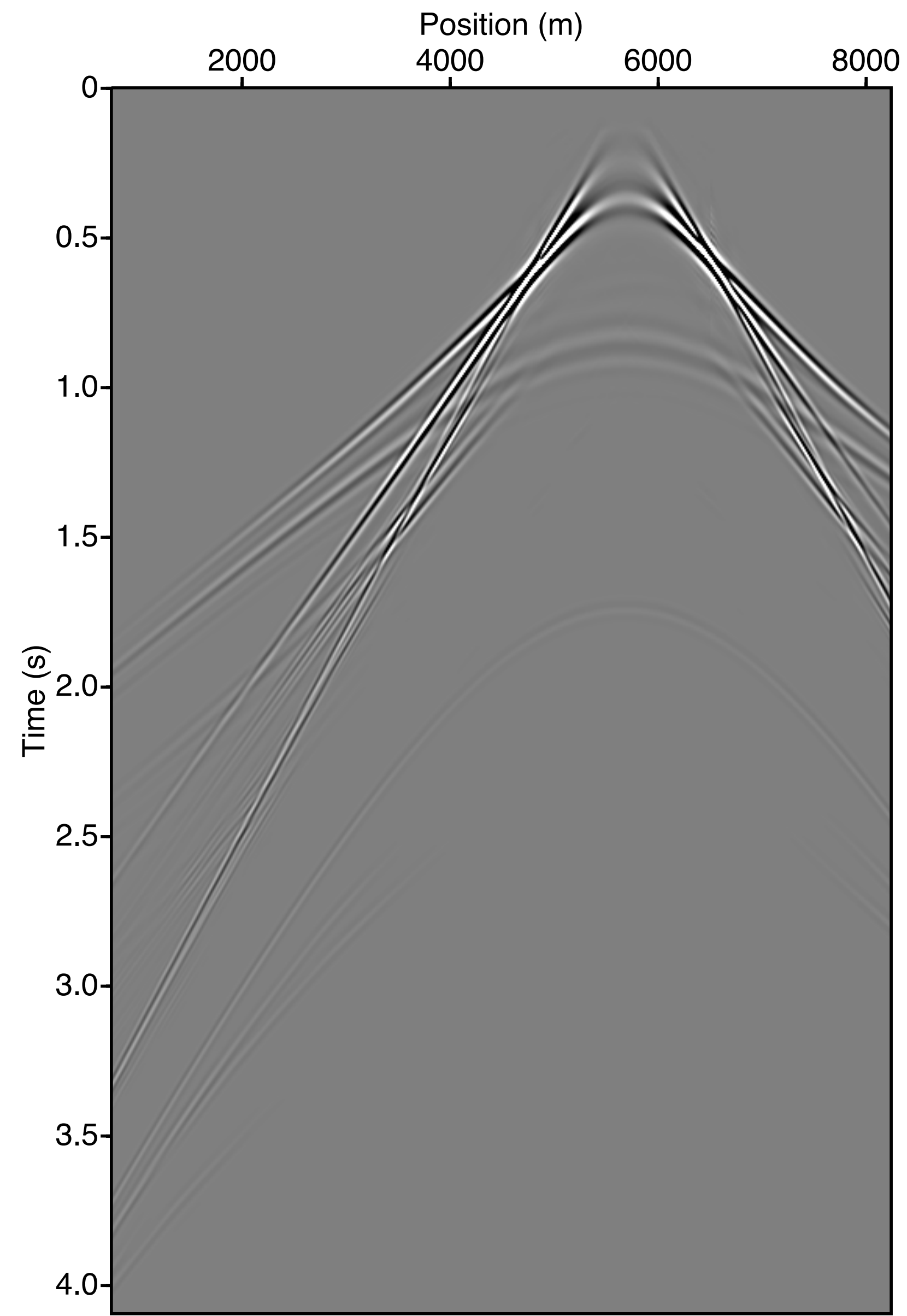
Potentially extract low-frequency information from G for diving-wave full-waveform inversion...?

# Bonus presentation

Preliminary study in low-frequency recovery of diving waves using Robust EPSI



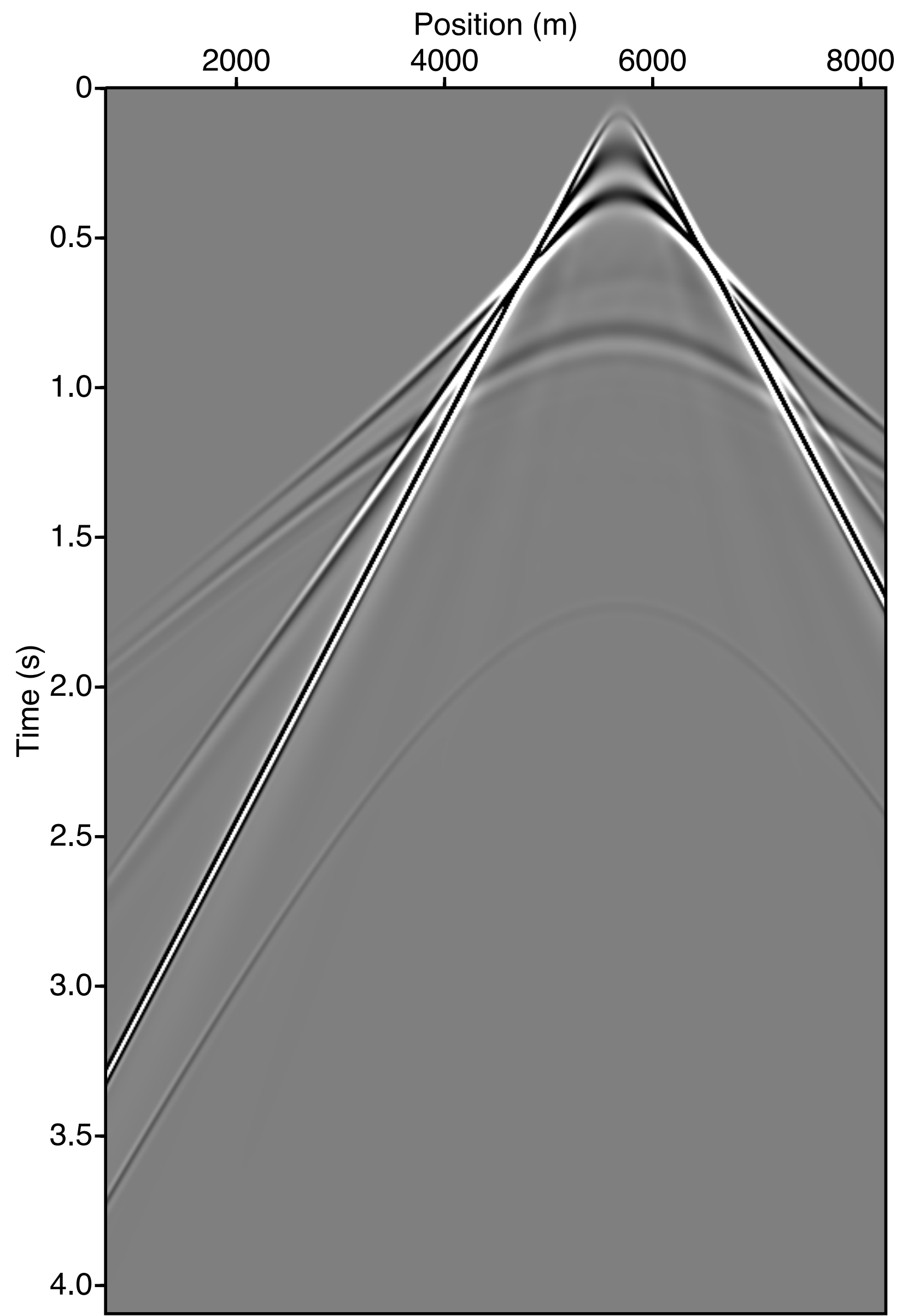
Data



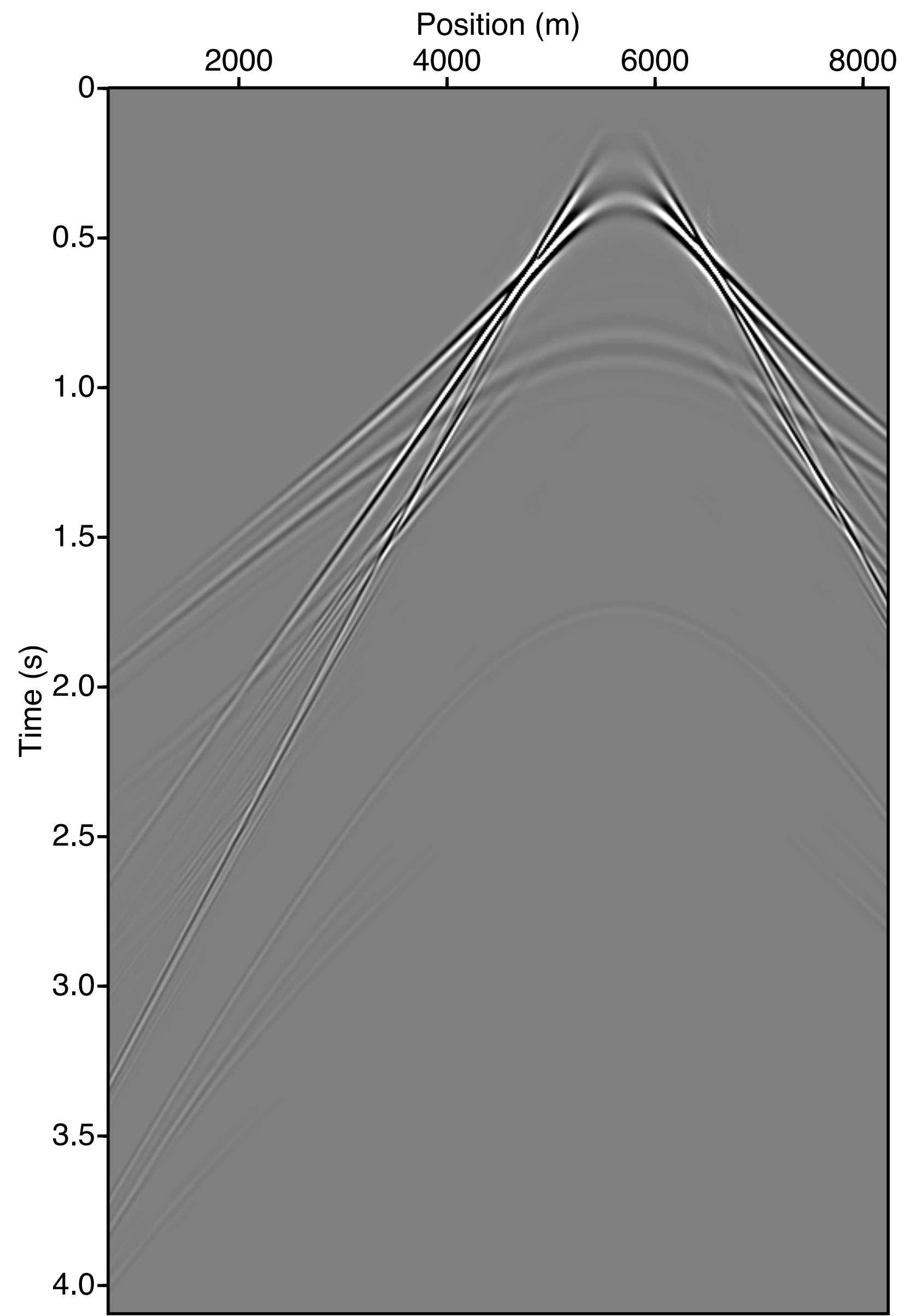
REPSI Primary

**REPSI “full band”**  
(only diving wave)  
**fixed spread of 7.5km**  
**ds=dr=15m**  
**15Hz Ricker, “full band”**  
**modeled w/iWAVE**





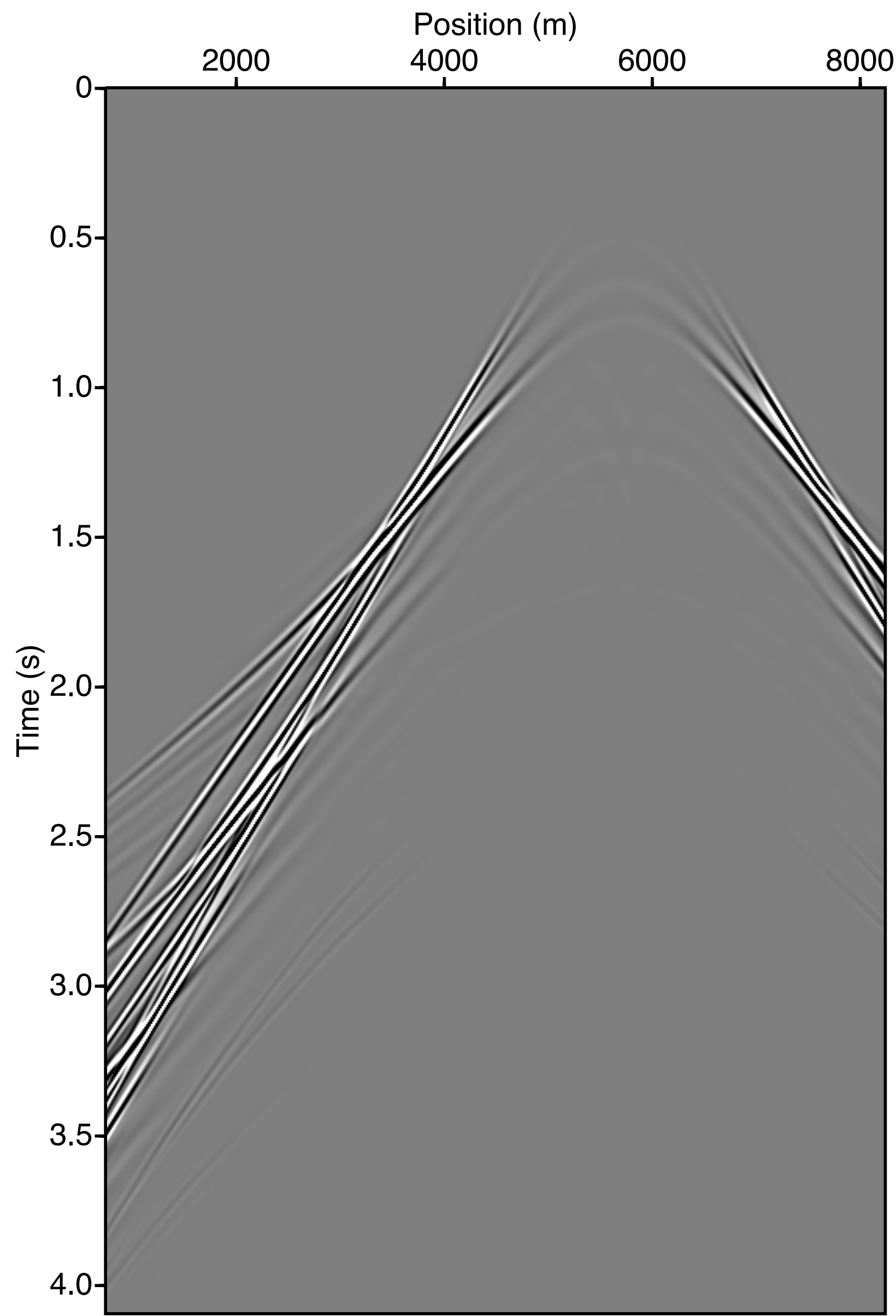
Modeled primary



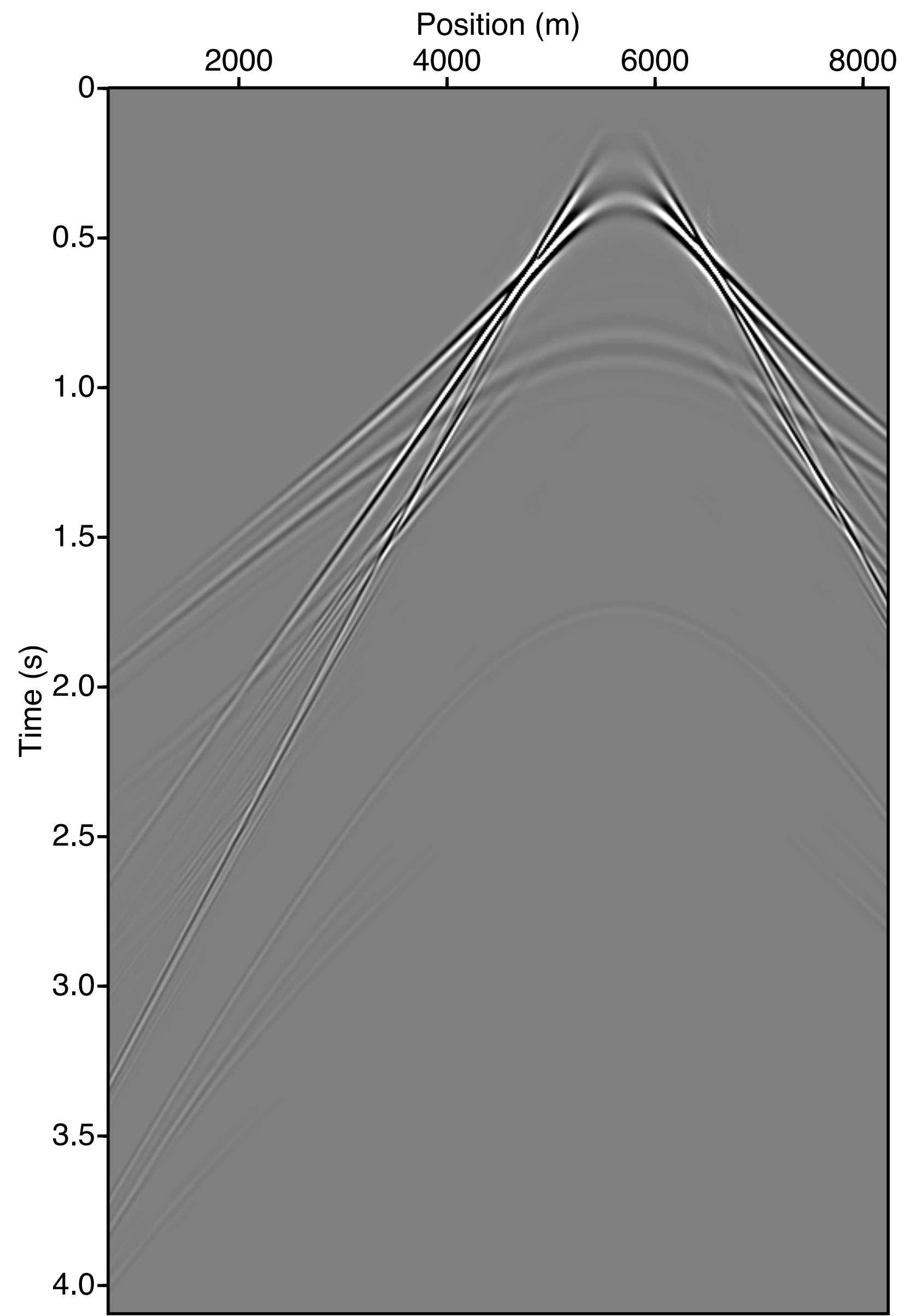
REPSI Primary

**REPSI “full band”**  
(only diving wave)  
**fixed spread of 7.5km**  
**ds=dr=15m**  
**15Hz Ricker, “full band”**  
**modeled w/iWAVE**



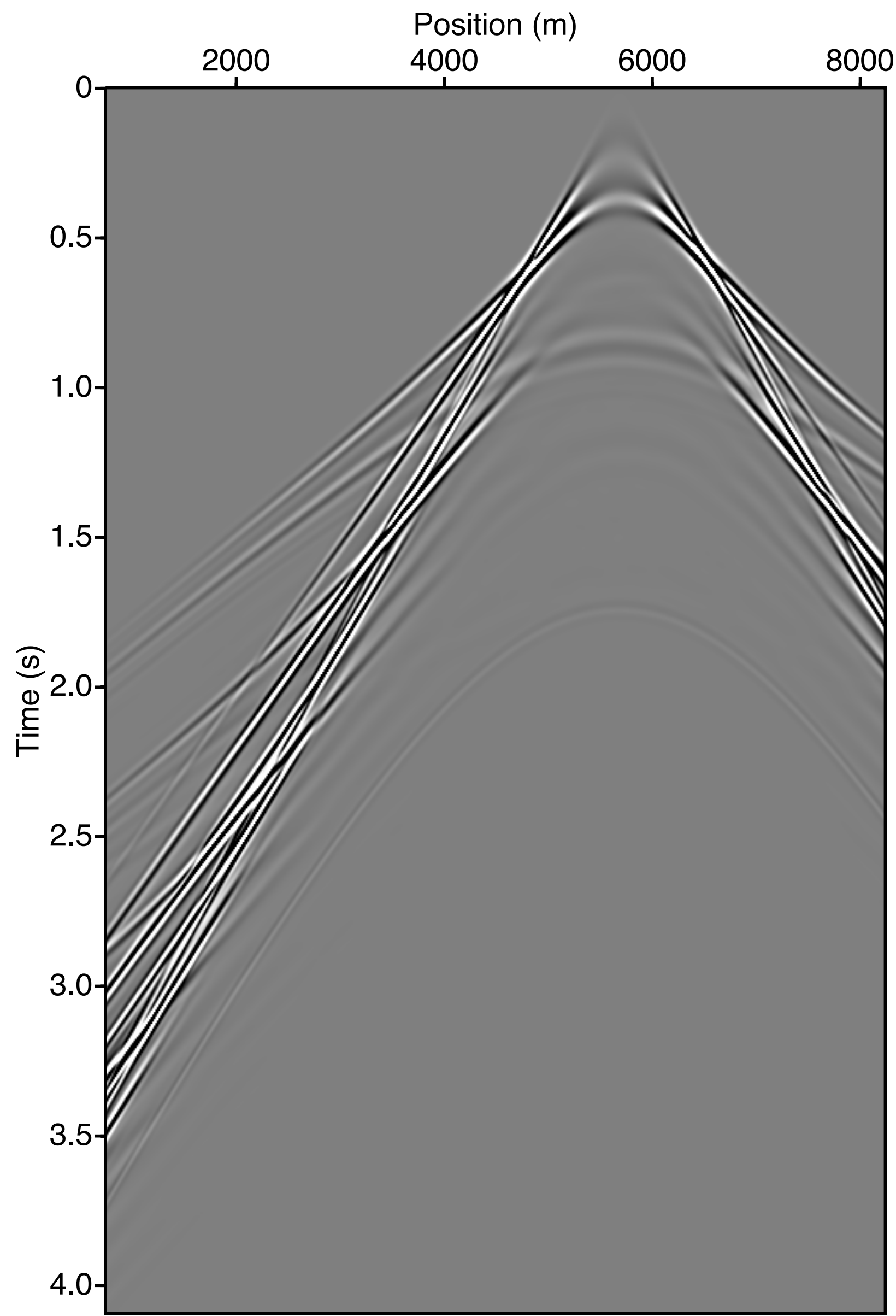


REPSI Multiples

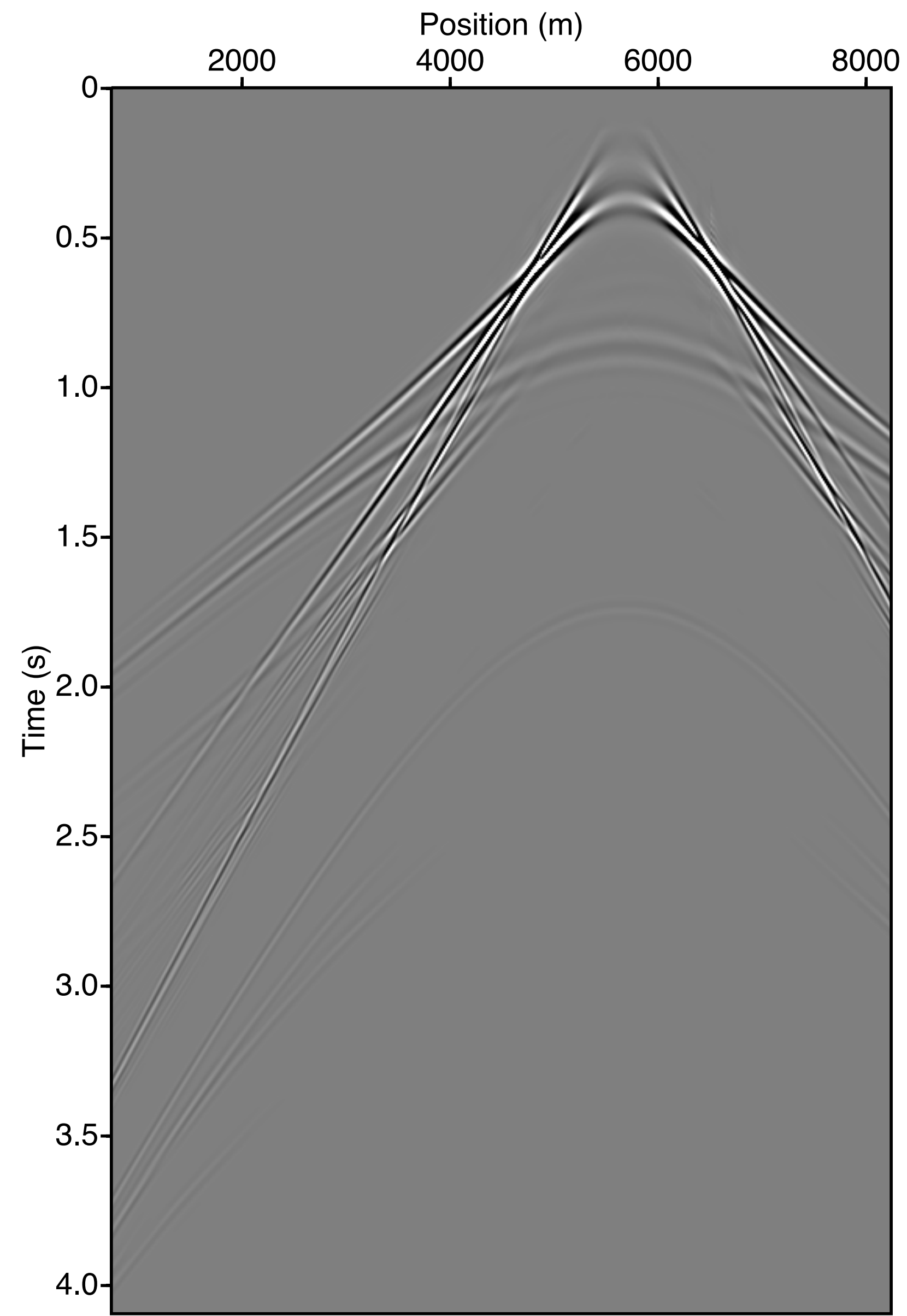


REPSI Primary

**REPSI “full band”**  
(only diving wave)  
**fixed spread of 7.5km**  
**ds=dr=15m**  
**15Hz Ricker, “full band”**  
**modeled w/iWAVE**



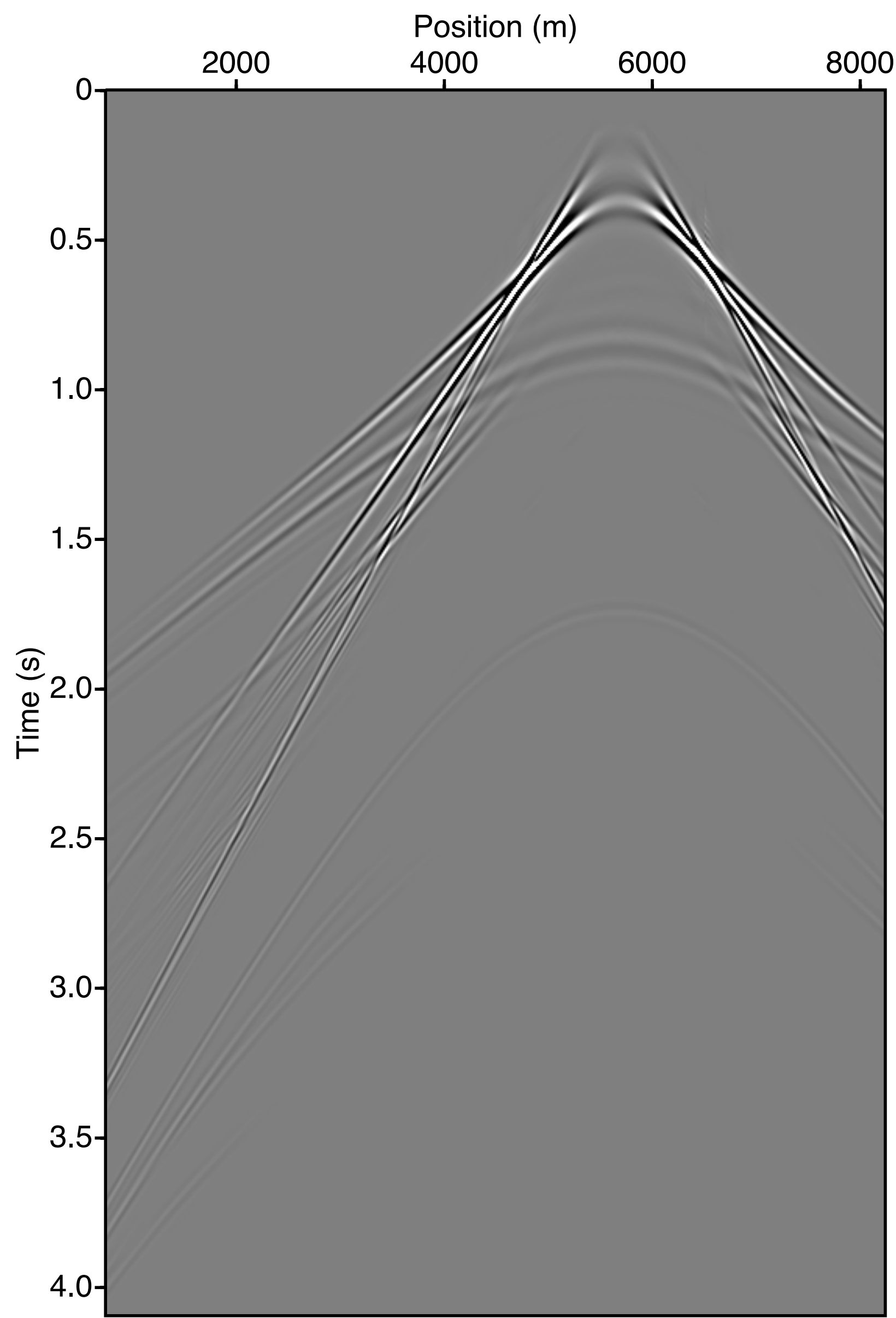
Data



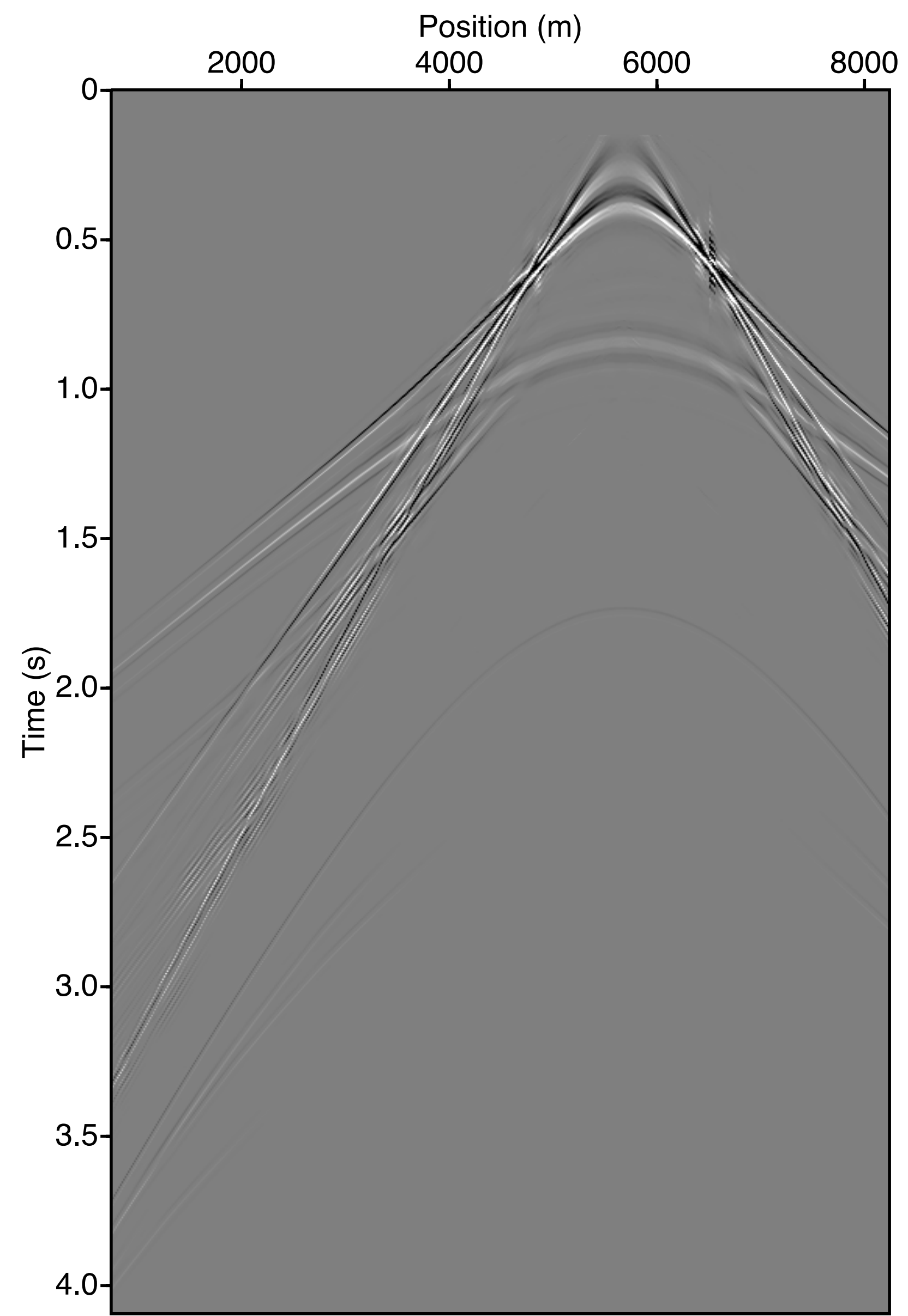
REPSI Primary

**REPSI “full band”**  
(only diving wave)  
**fixed spread of 7.5km**  
**ds=dr=15m**  
**15Hz Ricker, “full band”**  
**modeled w/iWAVE**



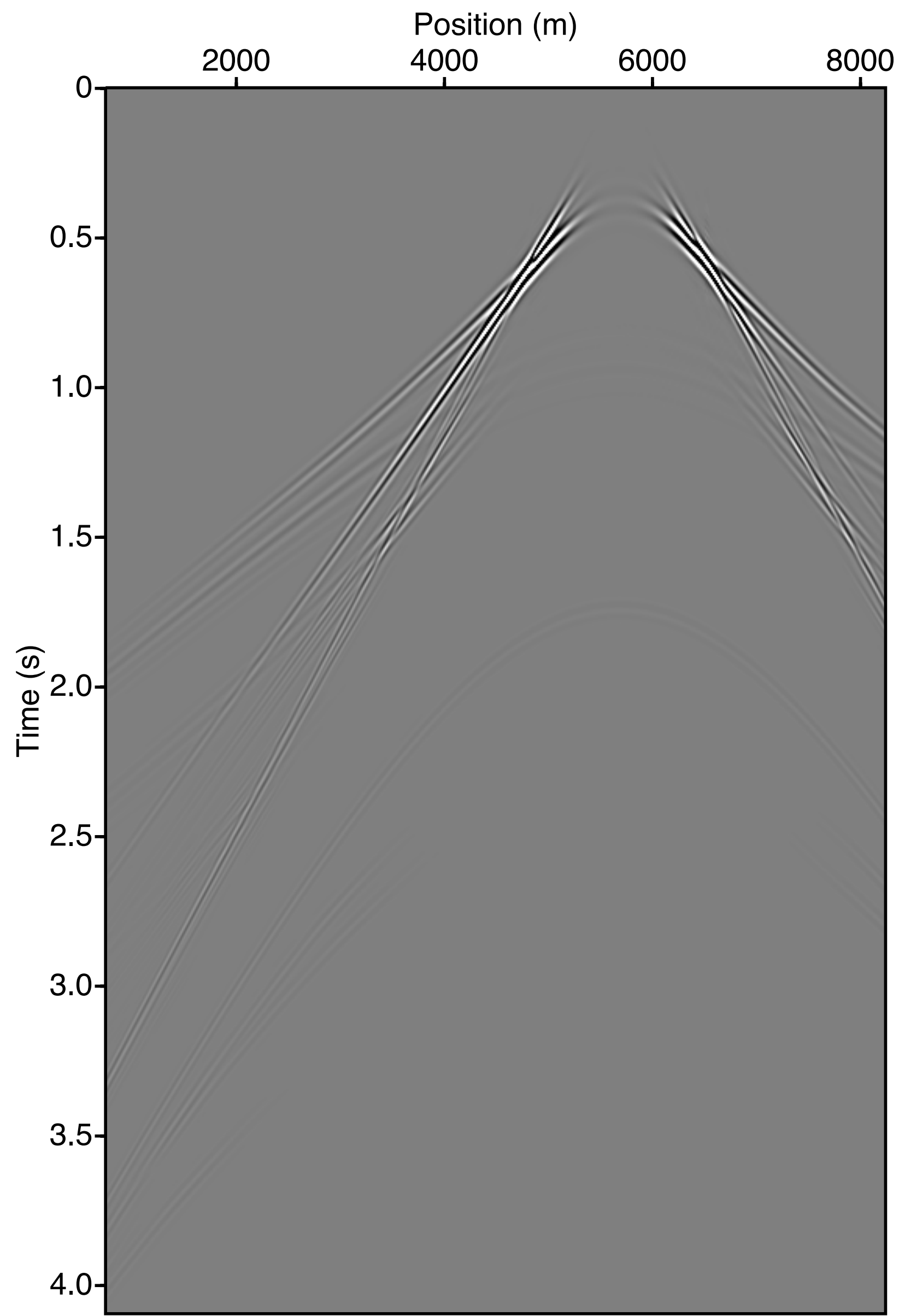


REPSI Primary

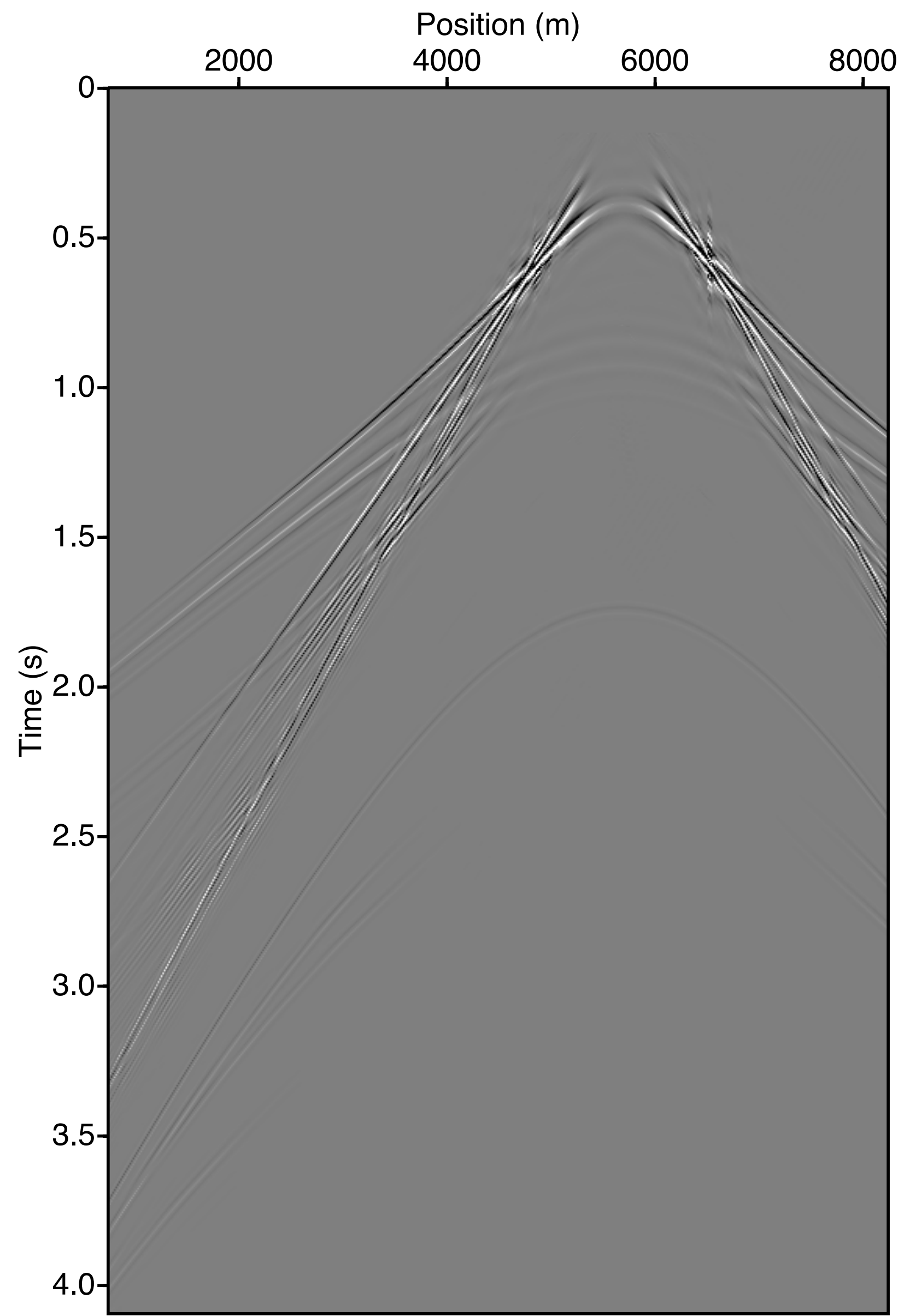


REPSI Primary IR

**REPSI “full band”**  
(only diving wave)  
**fixed spread of 7.5km**  
**ds=dr=15m**  
**15Hz Ricker, “full band”**  
**modeled w/iWAVE**



REPSI Primary

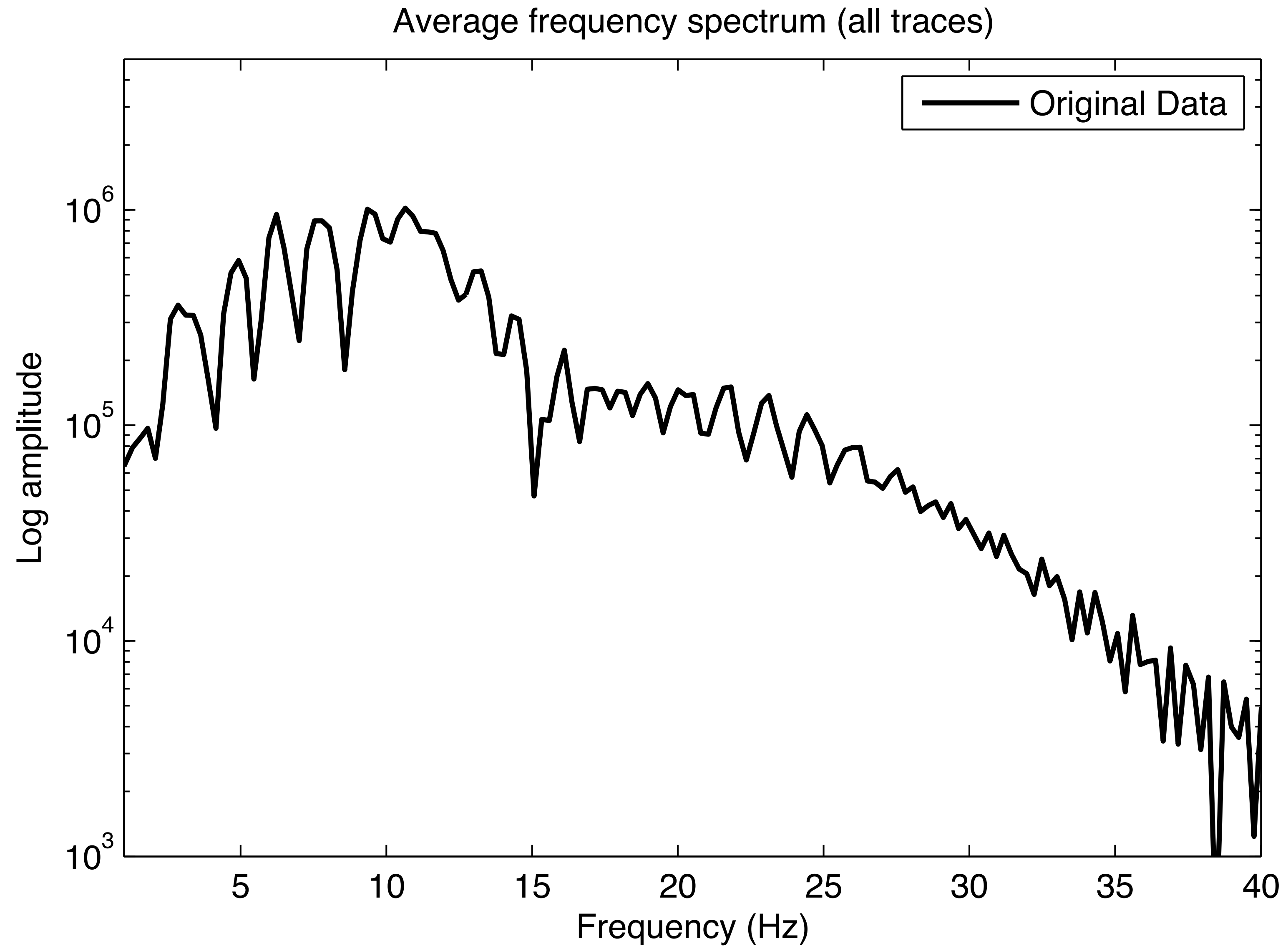


REPSI Primary IR

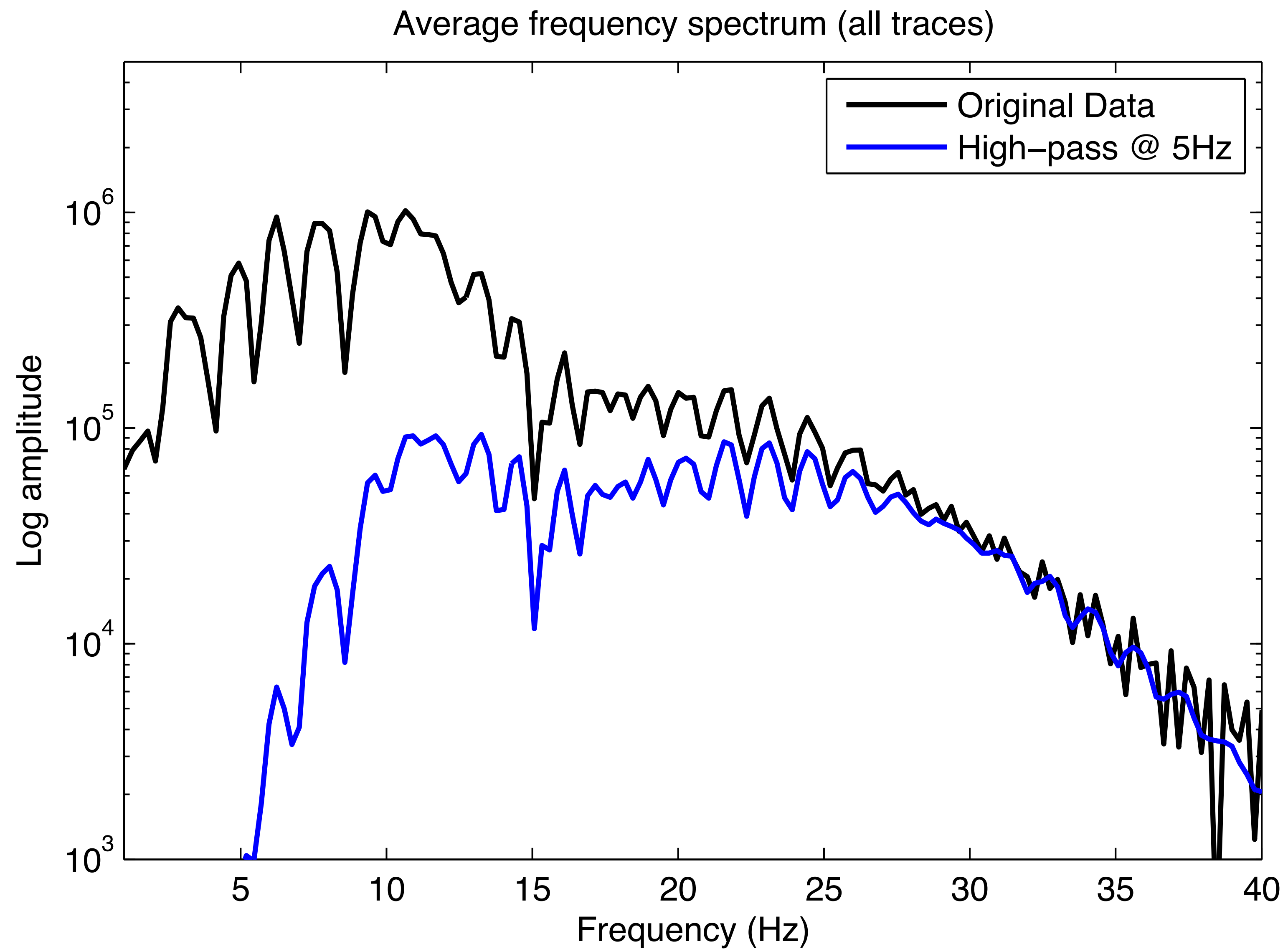
**Low cut at 5Hz**  
(only diving wave)  
**fixed spread of 7.5km**  
**ds=dr=15m**  
**15Hz Ricker, "full band"**  
**modeled w/iWAVE**  
**Low cut at 5Hz**



# Average trace spectrum



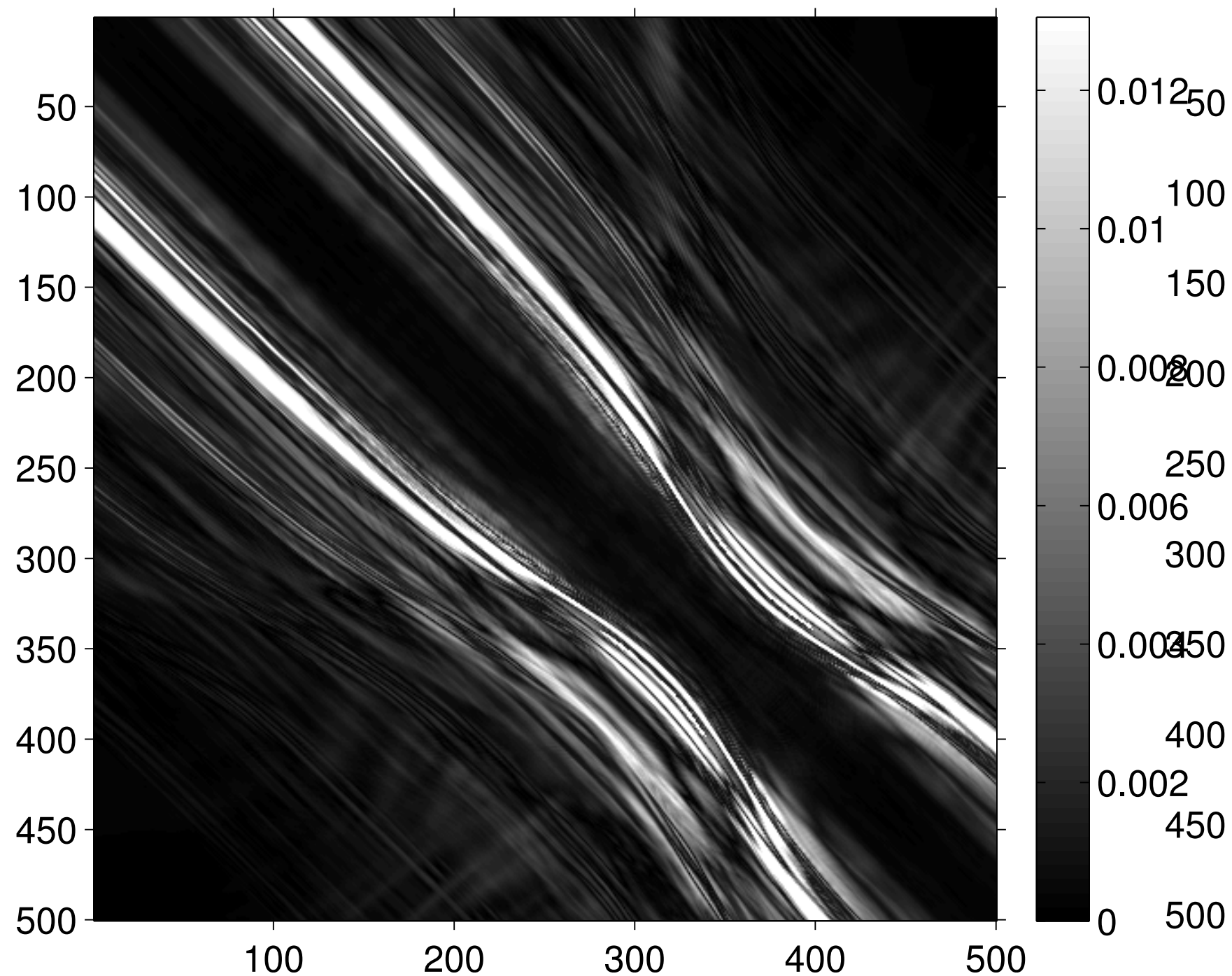
# Average trace spectrum





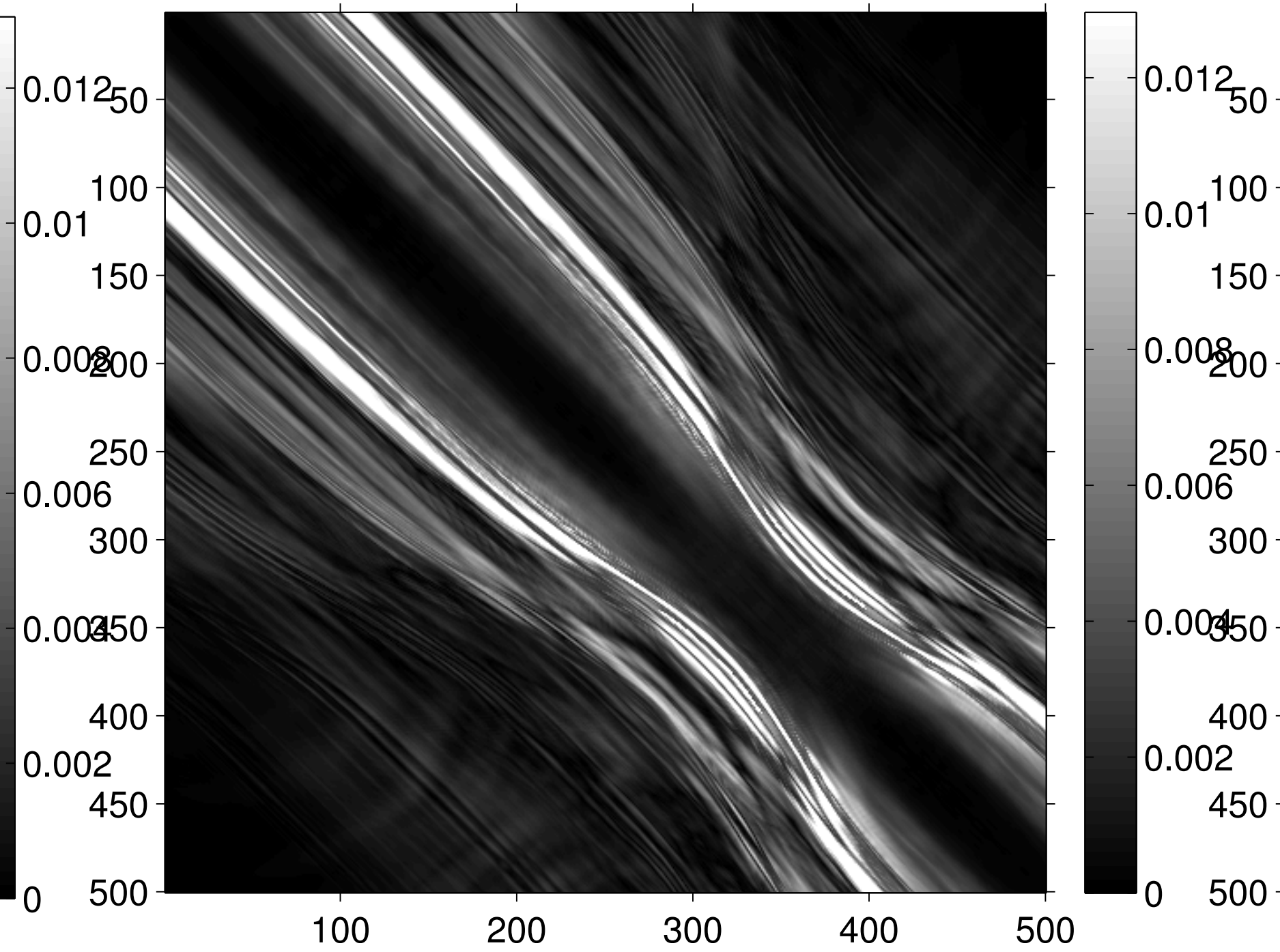
# Amplitude Reference solution

0.5Hz, Reference solution



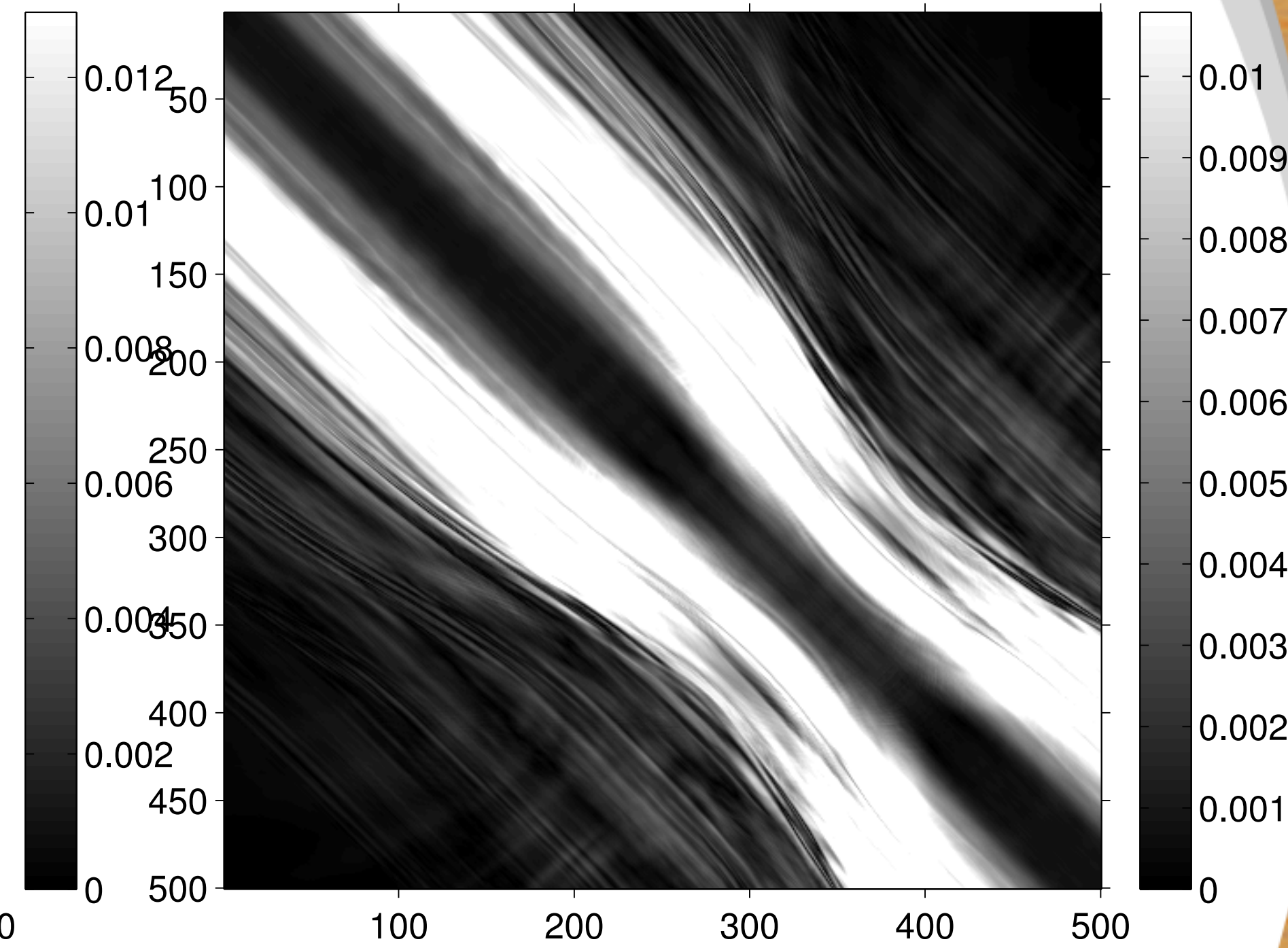
**0.5Hz**

1Hz, Reference solution



**1 Hz**

3Hz, Reference solution

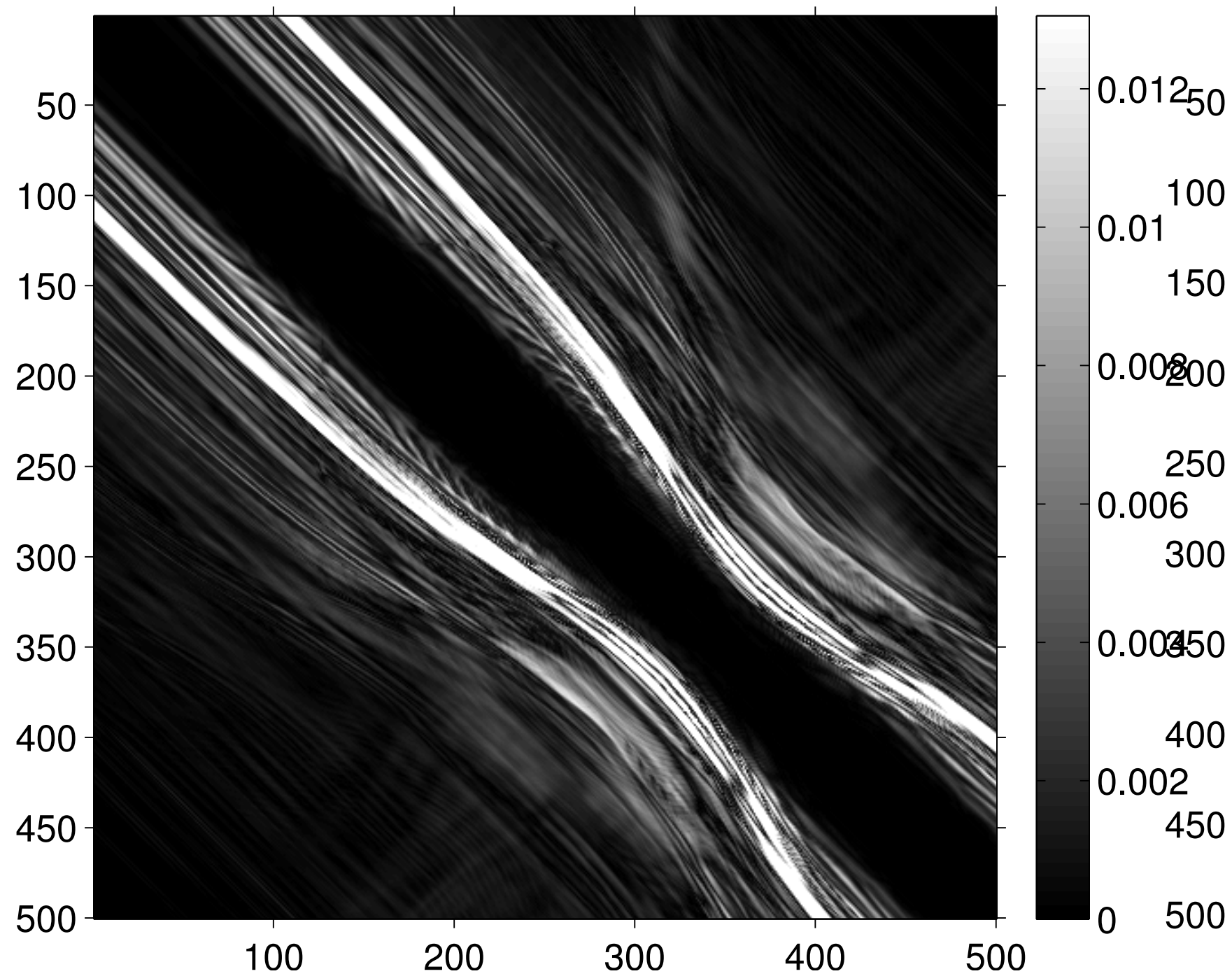


**3 Hz**



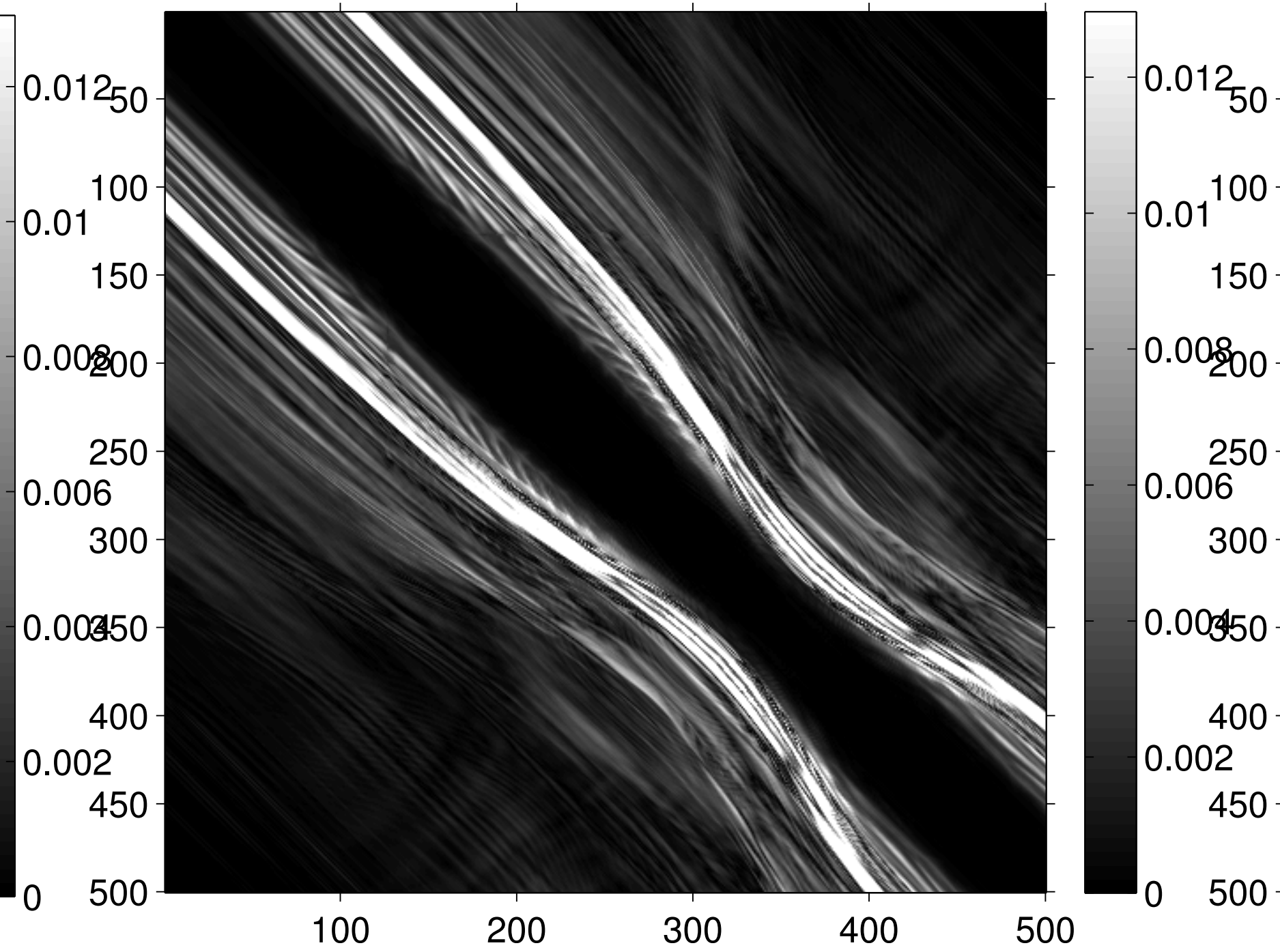
# Amplitude Low-cut at 5 Hz

0.5Hz, High-passed @ 5Hz solution



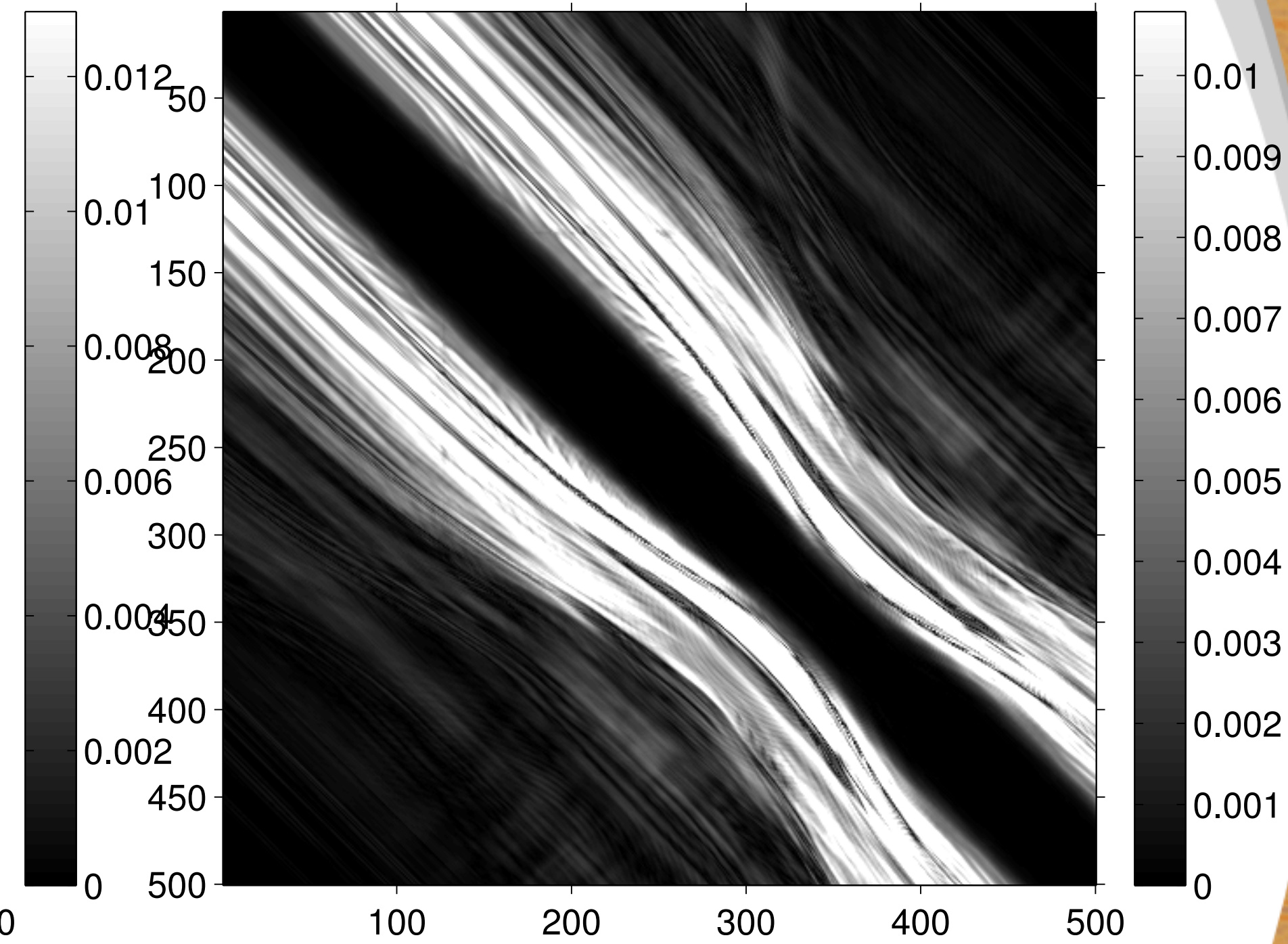
**0.5Hz**

1Hz, High-passed @ 5Hz solution



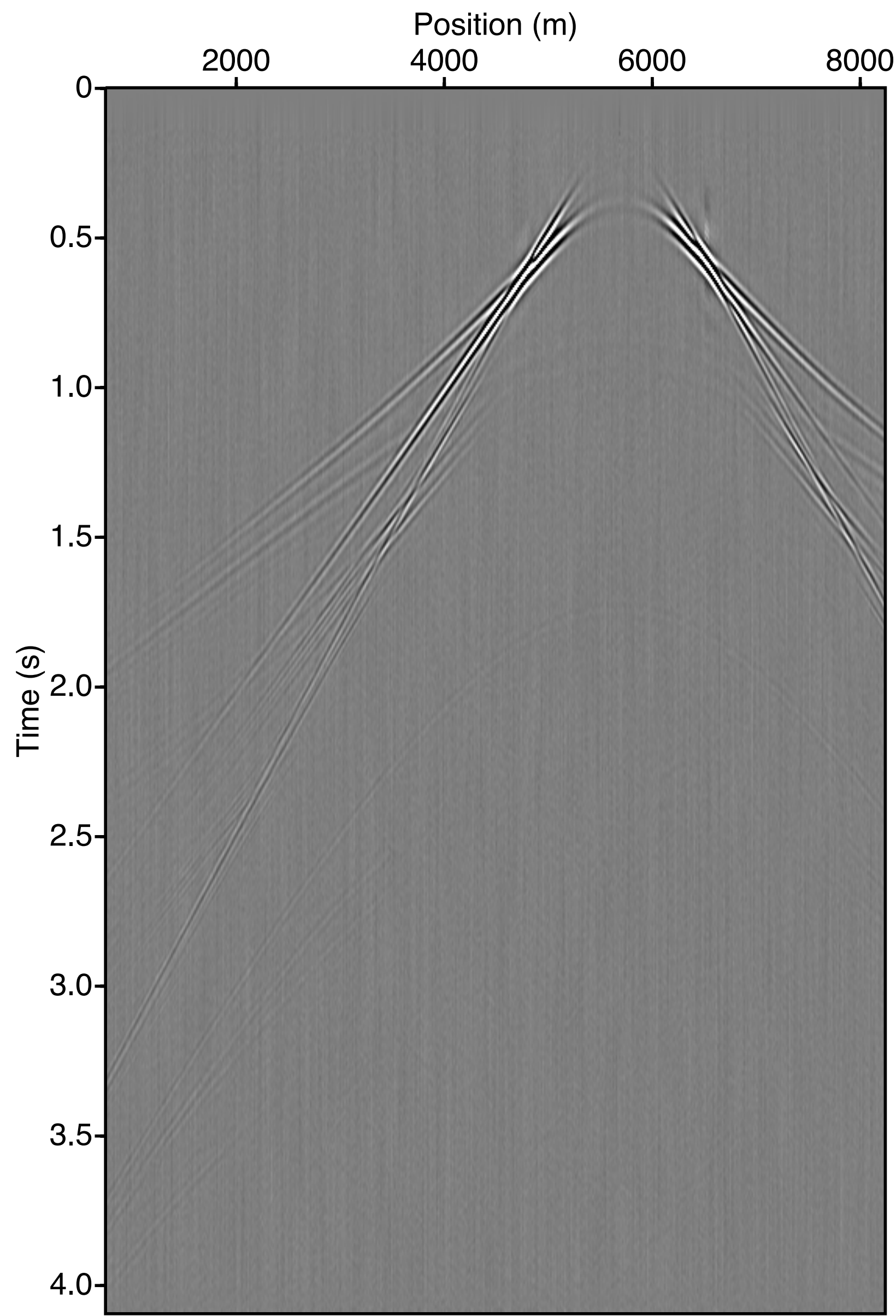
**1 Hz**

3Hz, High-passed @ 5Hz solution

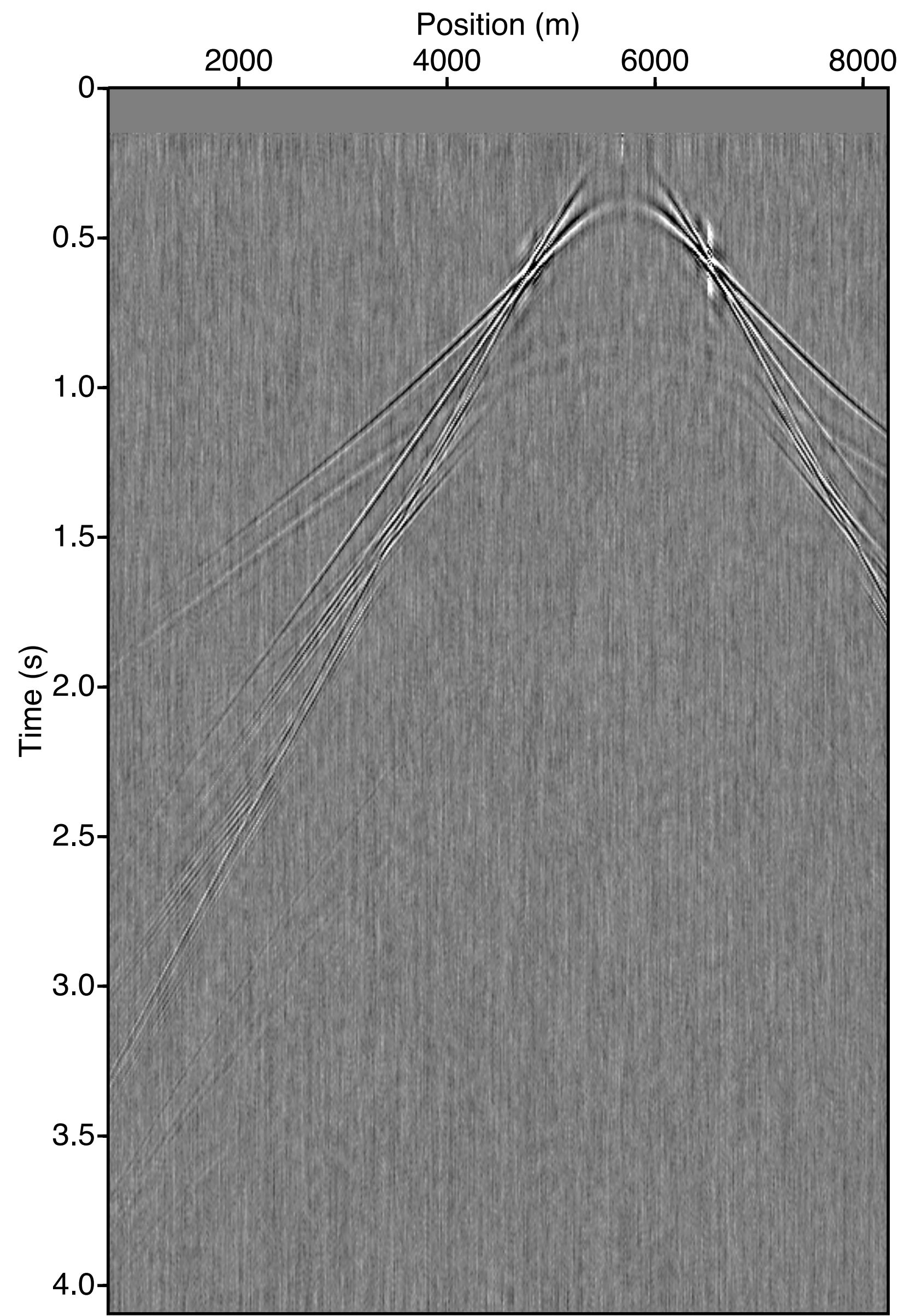


**3 Hz**





REPSI Primary

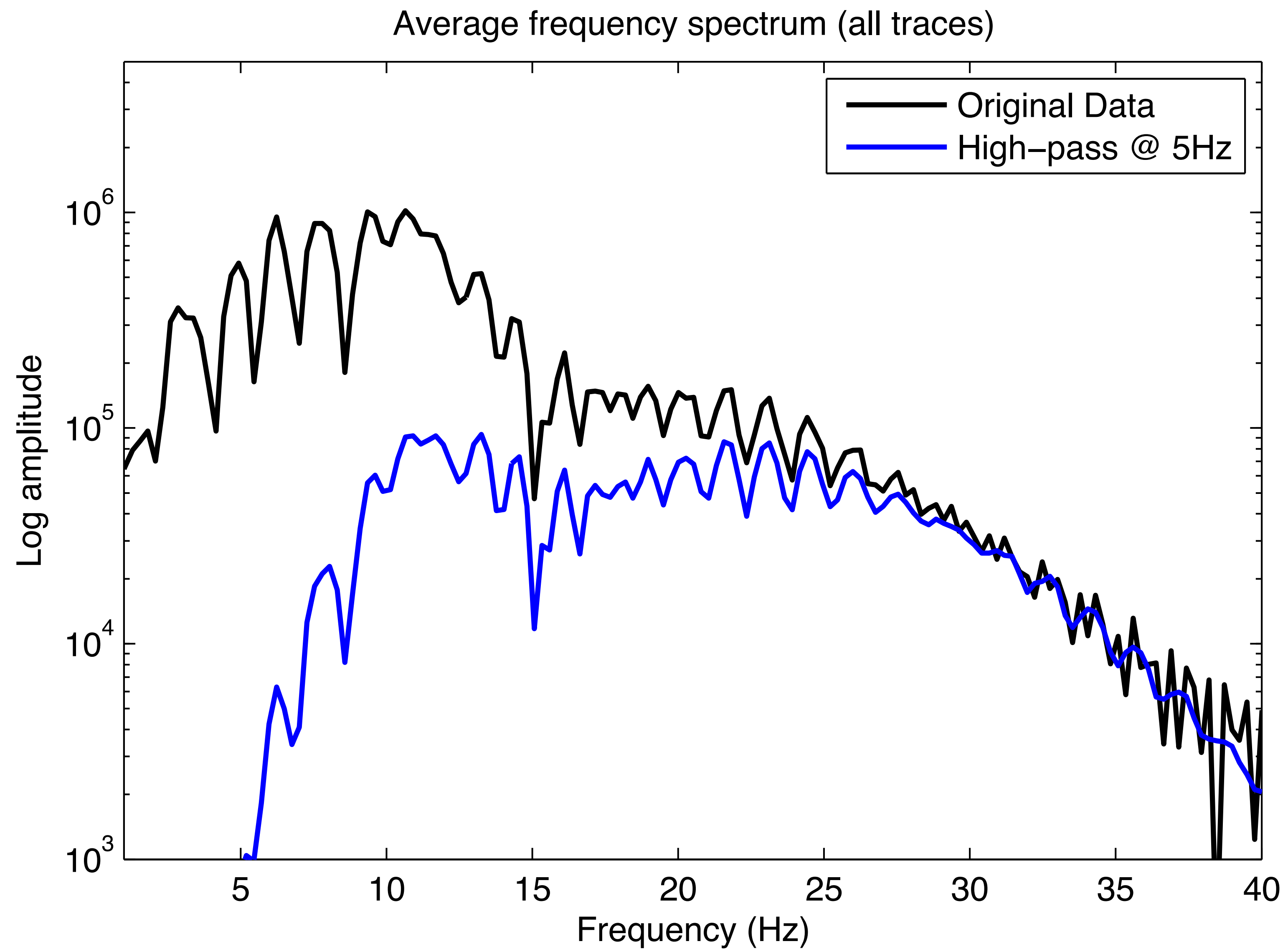


REPSI Primary IR

**Low cut at 5Hz + Noise**  
(only diving wave)  
**fixed spread of 7.5km**  
**ds=dr=15m**  
**15Hz Ricker, "full band"**  
**modeled w/iWAVE**  
**Low cut at 5Hz**  
**18dB pink noise added**  
(i.d.d. per trace)

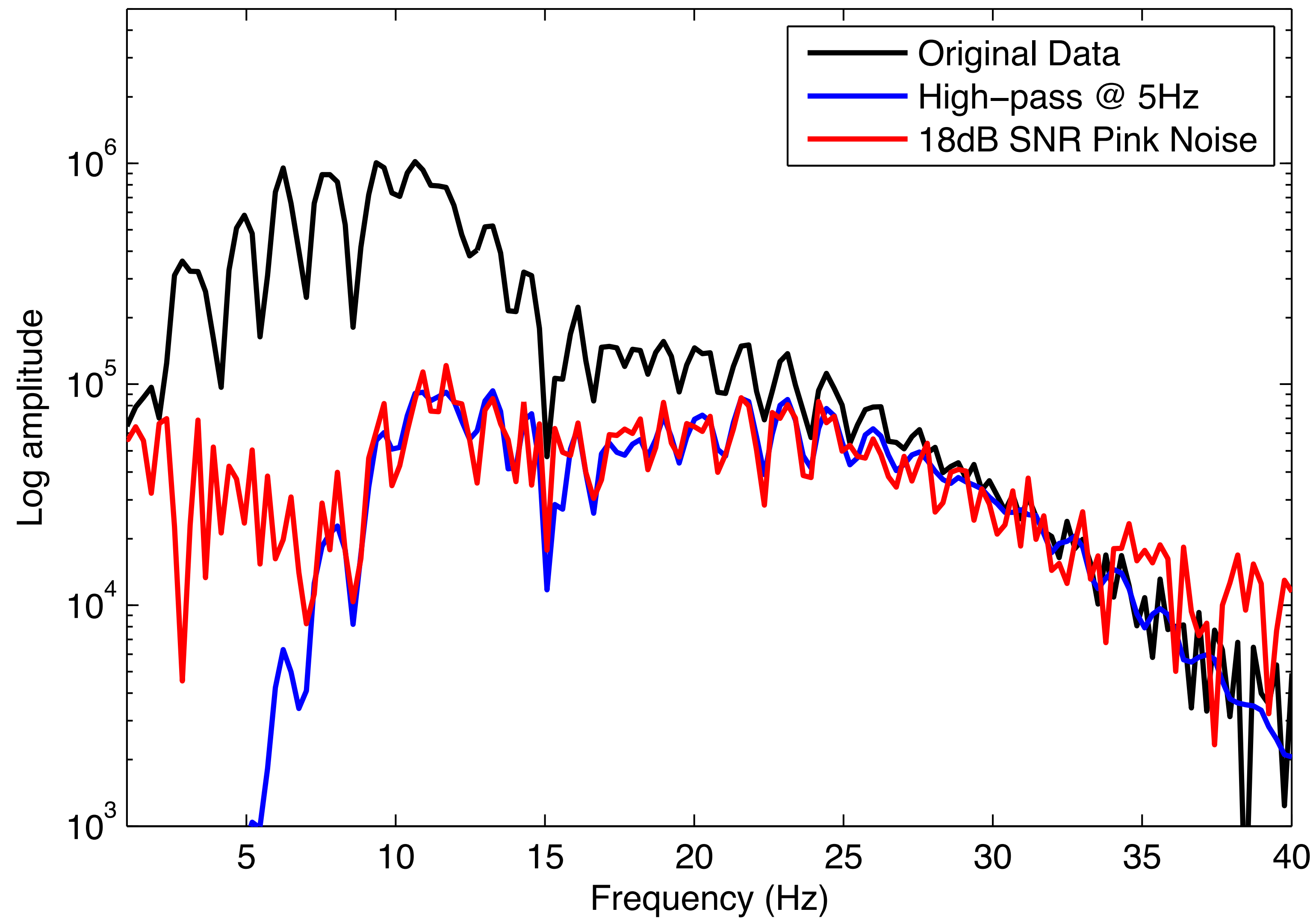


# Average trace spectrum



# Average trace spectrum

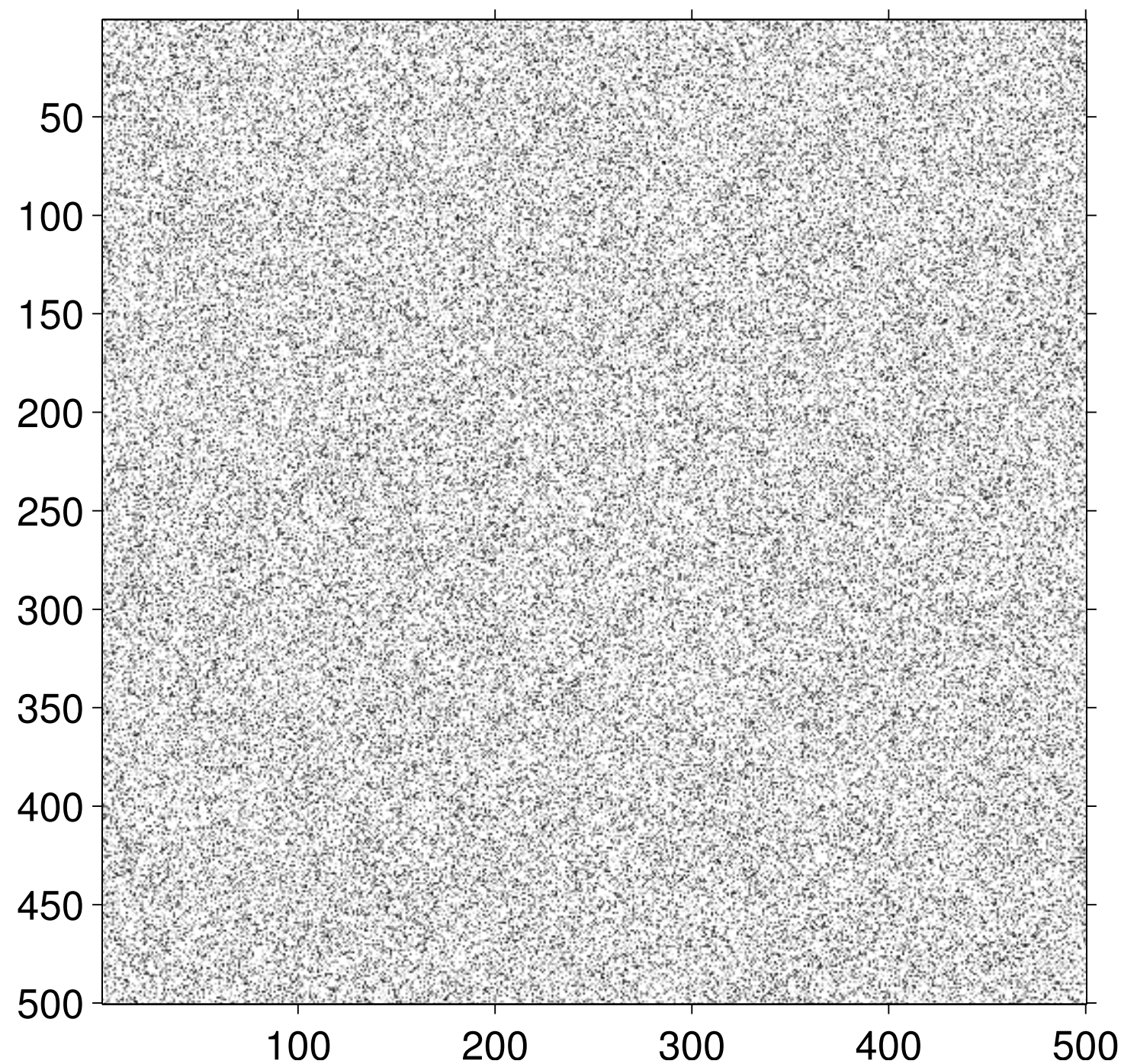
Average frequency spectrum (all traces)





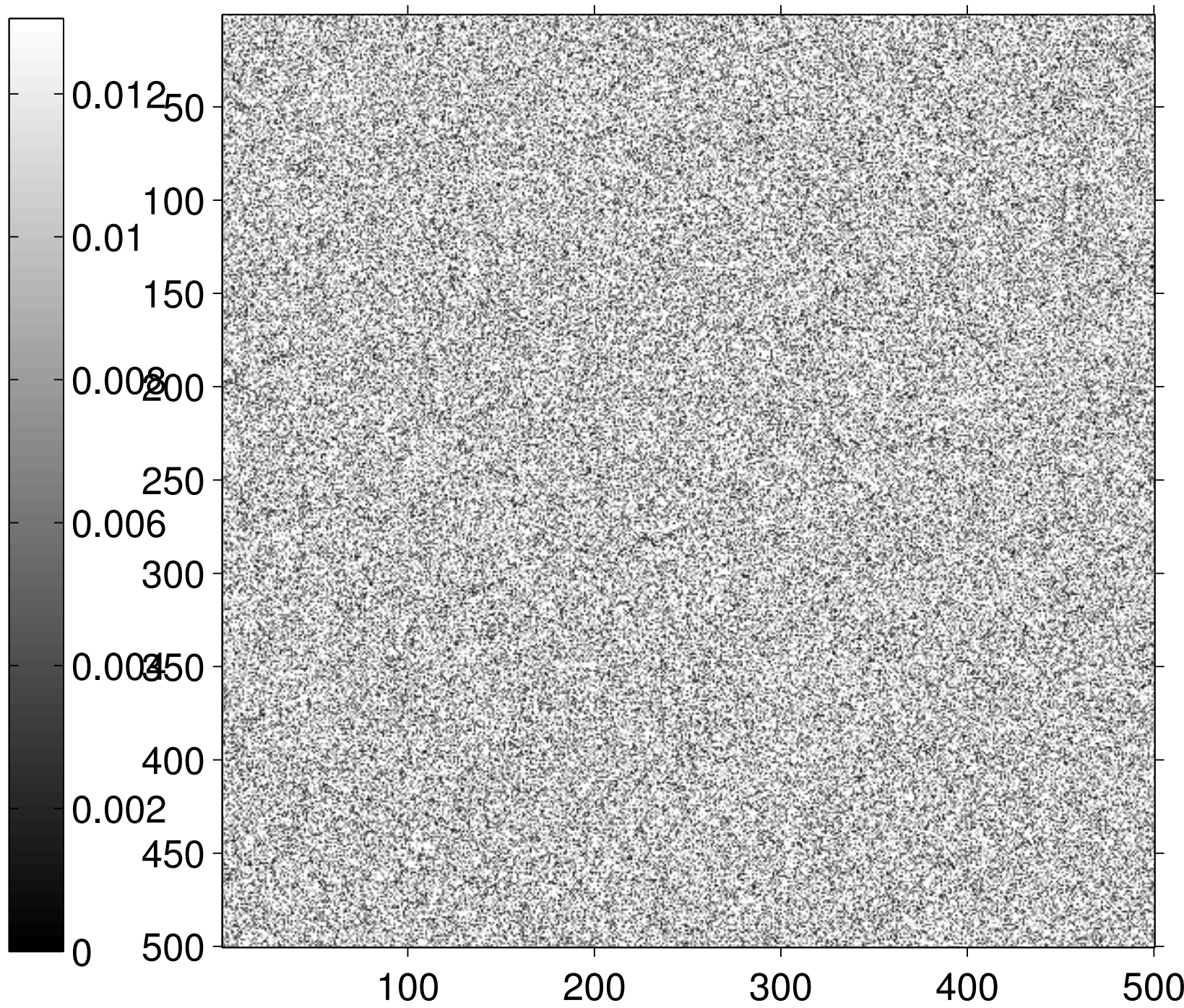
# Amplitude Low-cut at 5 Hz, 18dB Pink noise added

0.5Hz, High-passed @ 5Hz, 18dB SNR (pink noise) solution



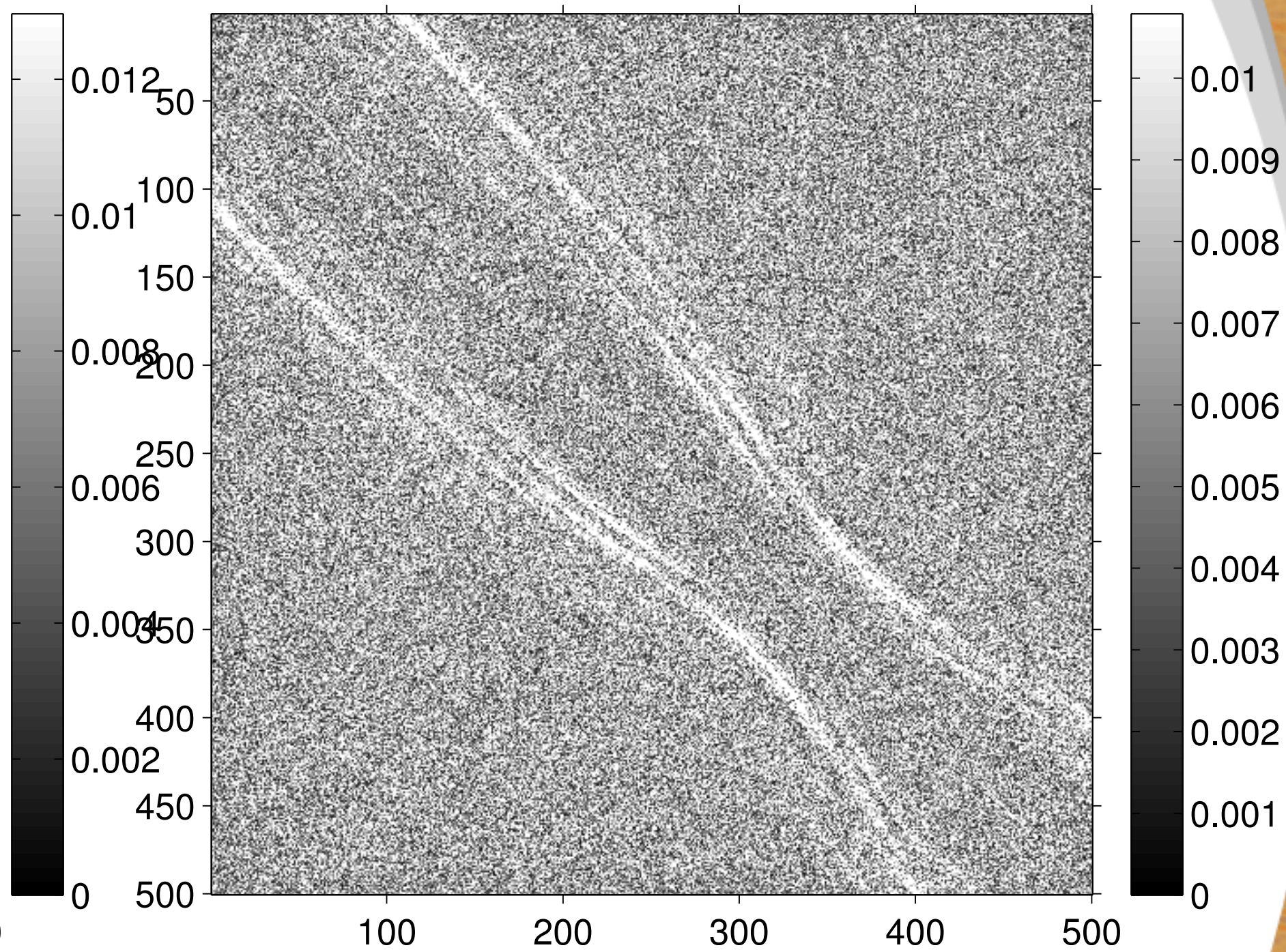
**0.5Hz**

1Hz, High-passed @ 5Hz, 18dB SNR (pink noise) solution



**1 Hz**

3Hz, High-passed @ 5Hz, 18dB SNR (pink noise) solution

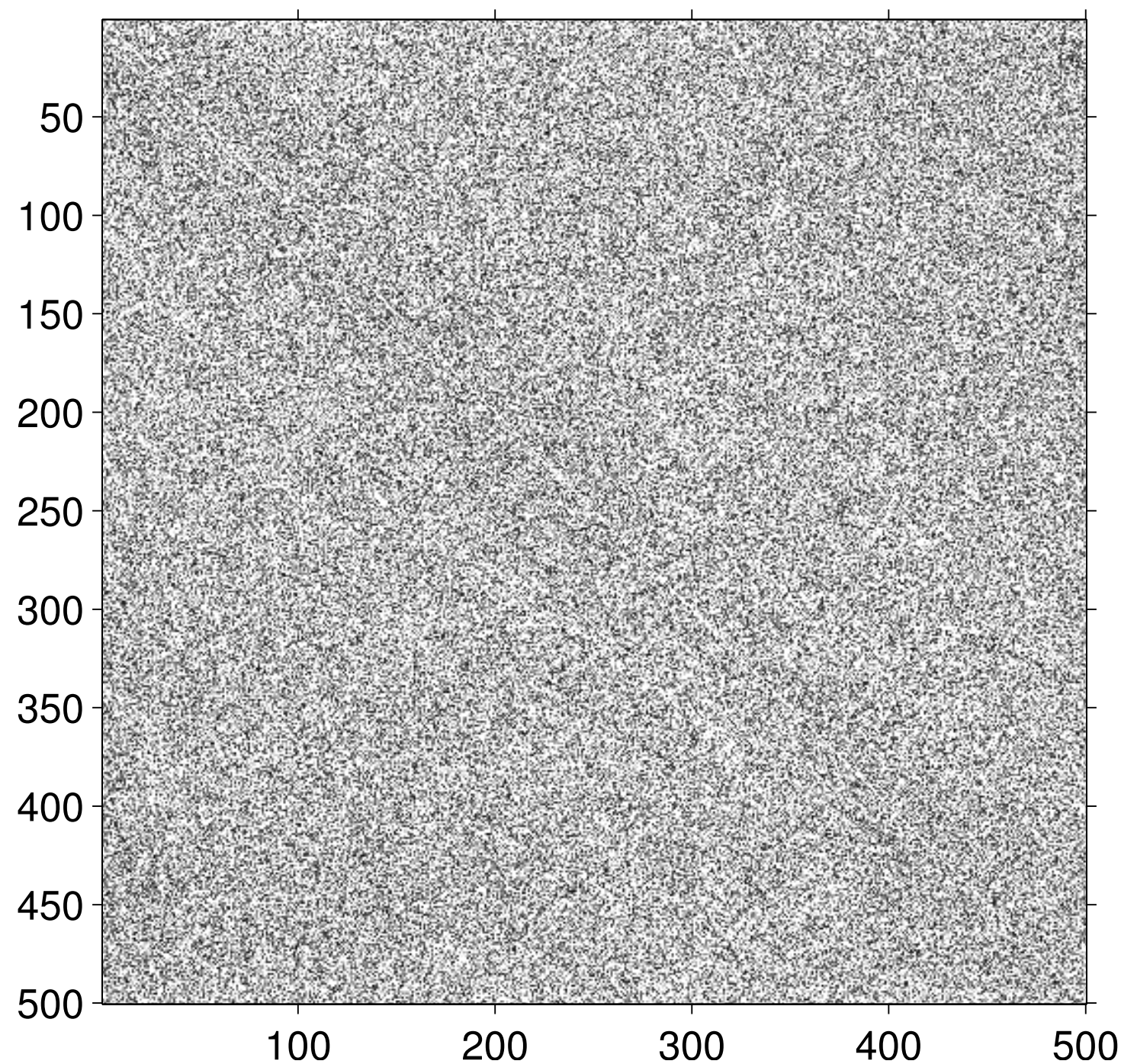


**3 Hz**

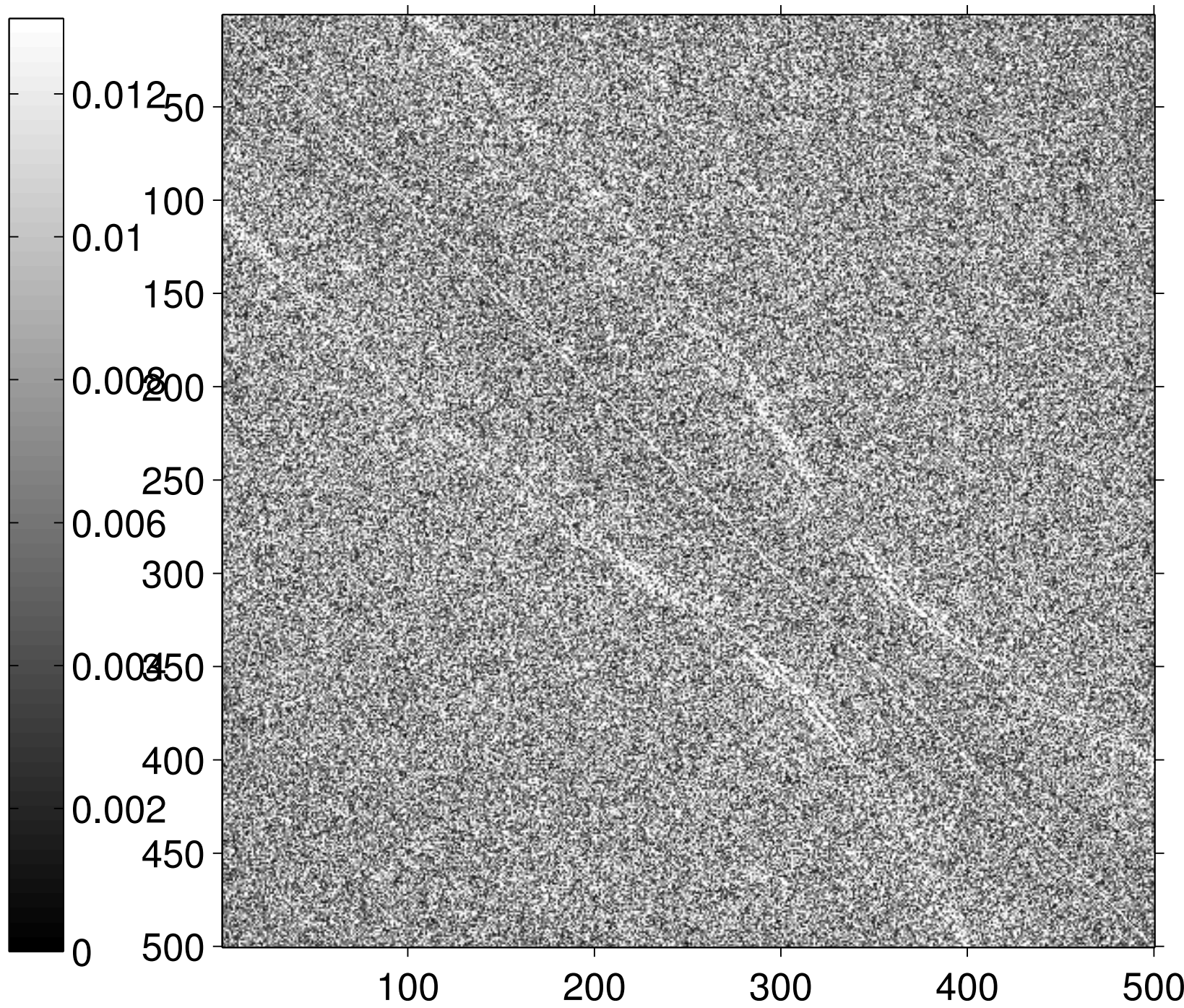


# Amplitude Low-cut at 5 Hz, 18dB Pink noise added (solved to exact sigma)

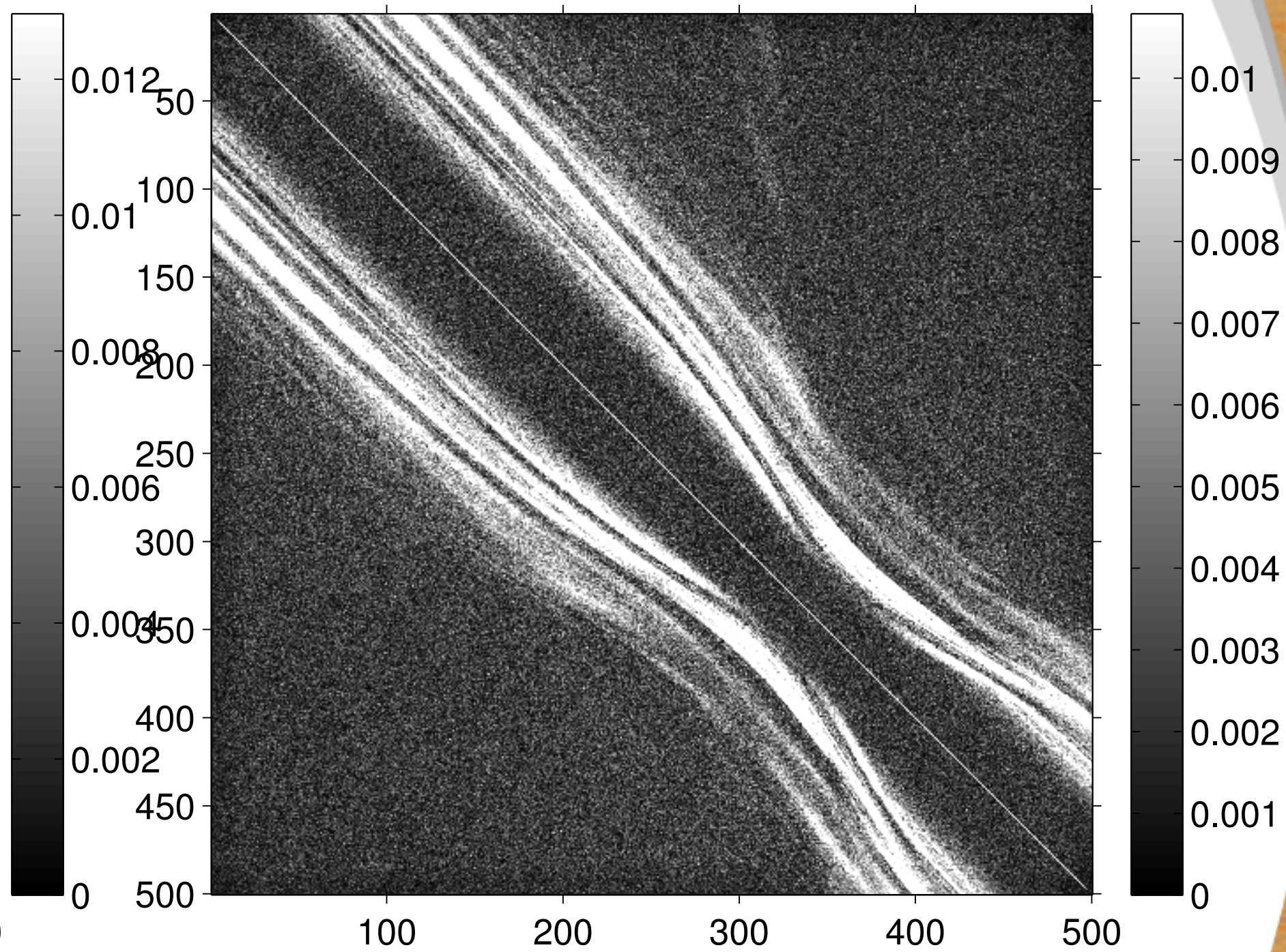
0.5 Hz, High-passed @ 5Hz, 18dB SNR (pink noise), exact sigma solution    1 Hz, High-passed @ 5Hz, 18dB SNR (pink noise), exact sigma solution    3 Hz, High-passed @ 5Hz, 18dB SNR (pink noise), exact sigma solution



**0.5 Hz**

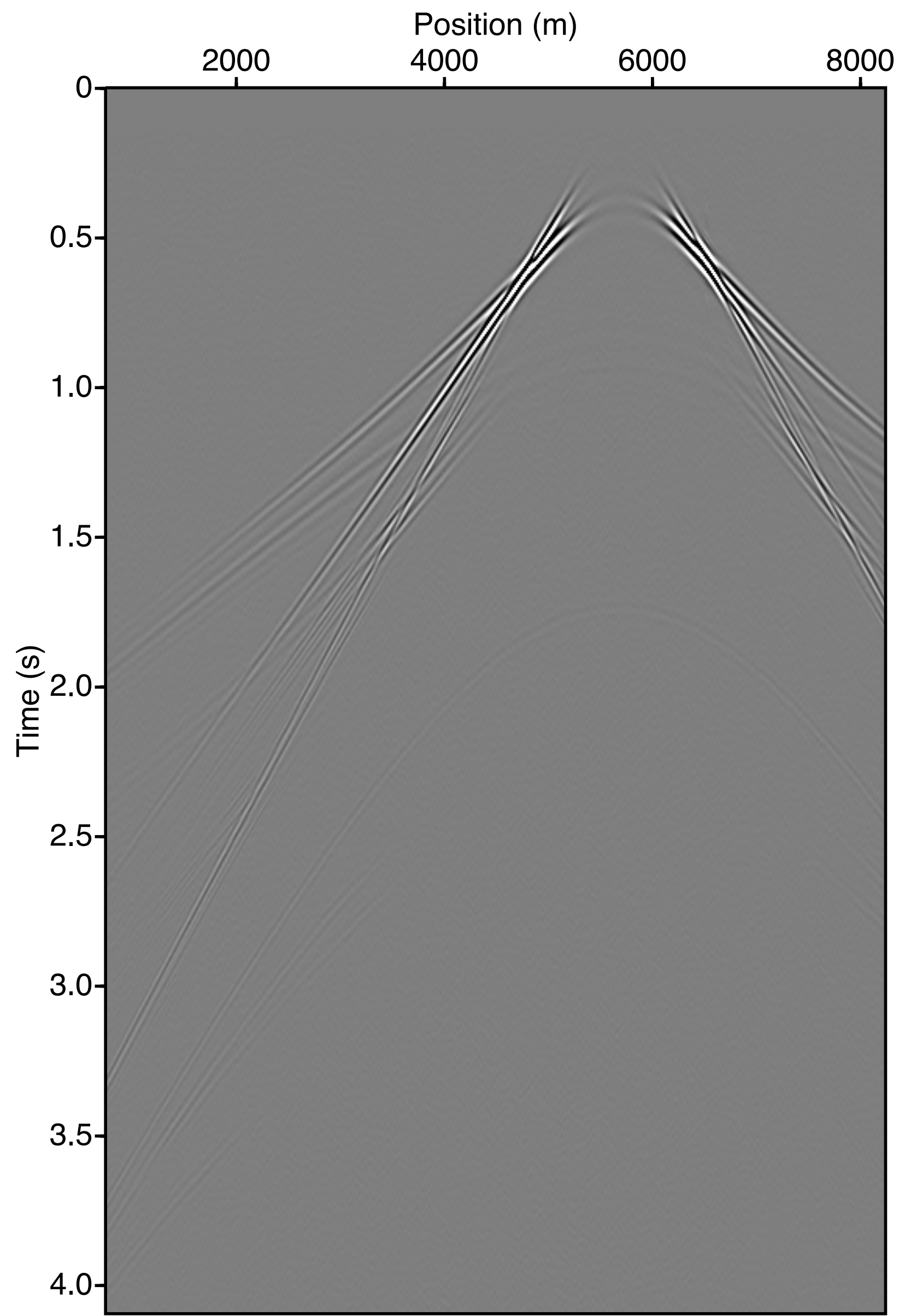


**1 Hz**

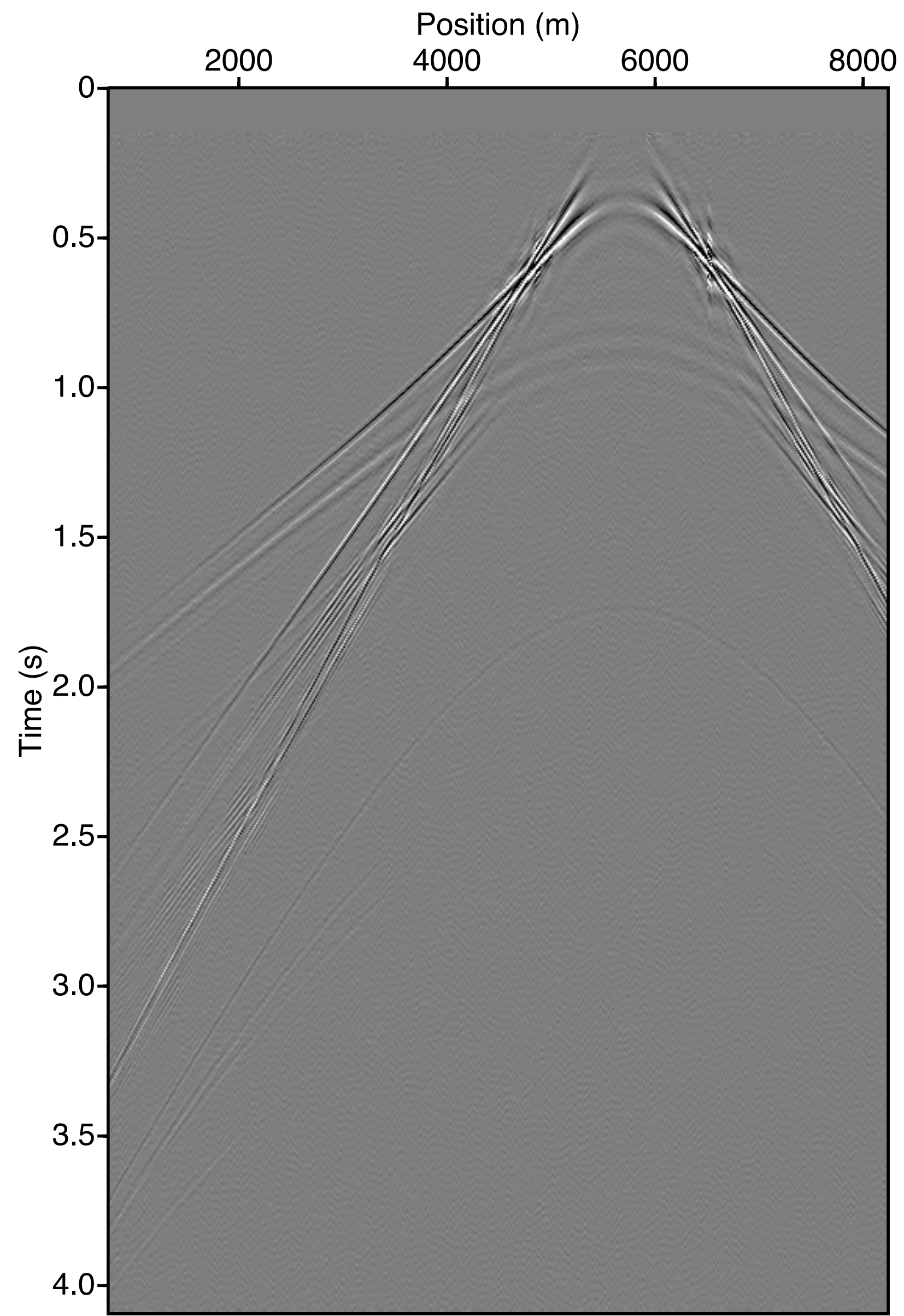


**3 Hz**





REPSI Primary



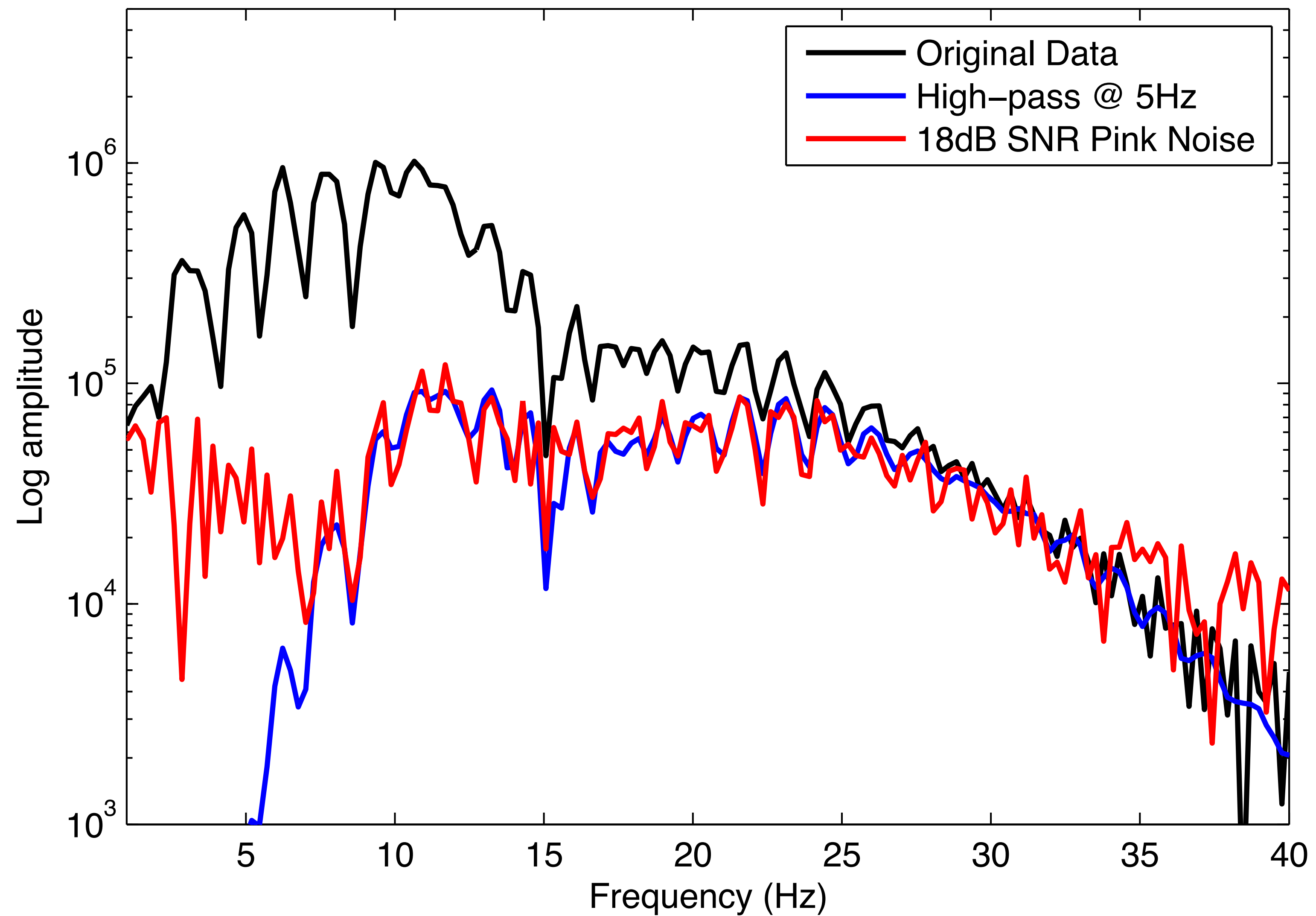
REPSI Primary IR

**Noise + Low cut at 5Hz**  
(only diving wave)  
**fixed spread of 7.5km**  
**ds=dr=15m**  
**15Hz Ricker, "full band"**  
**modeled w/iWAVE**  
**18dB pink noise added**  
**(i.d.d. per trace)**  
**Low cut at 5Hz**  
(low-cut *after* noise is added)



# Average trace spectrum

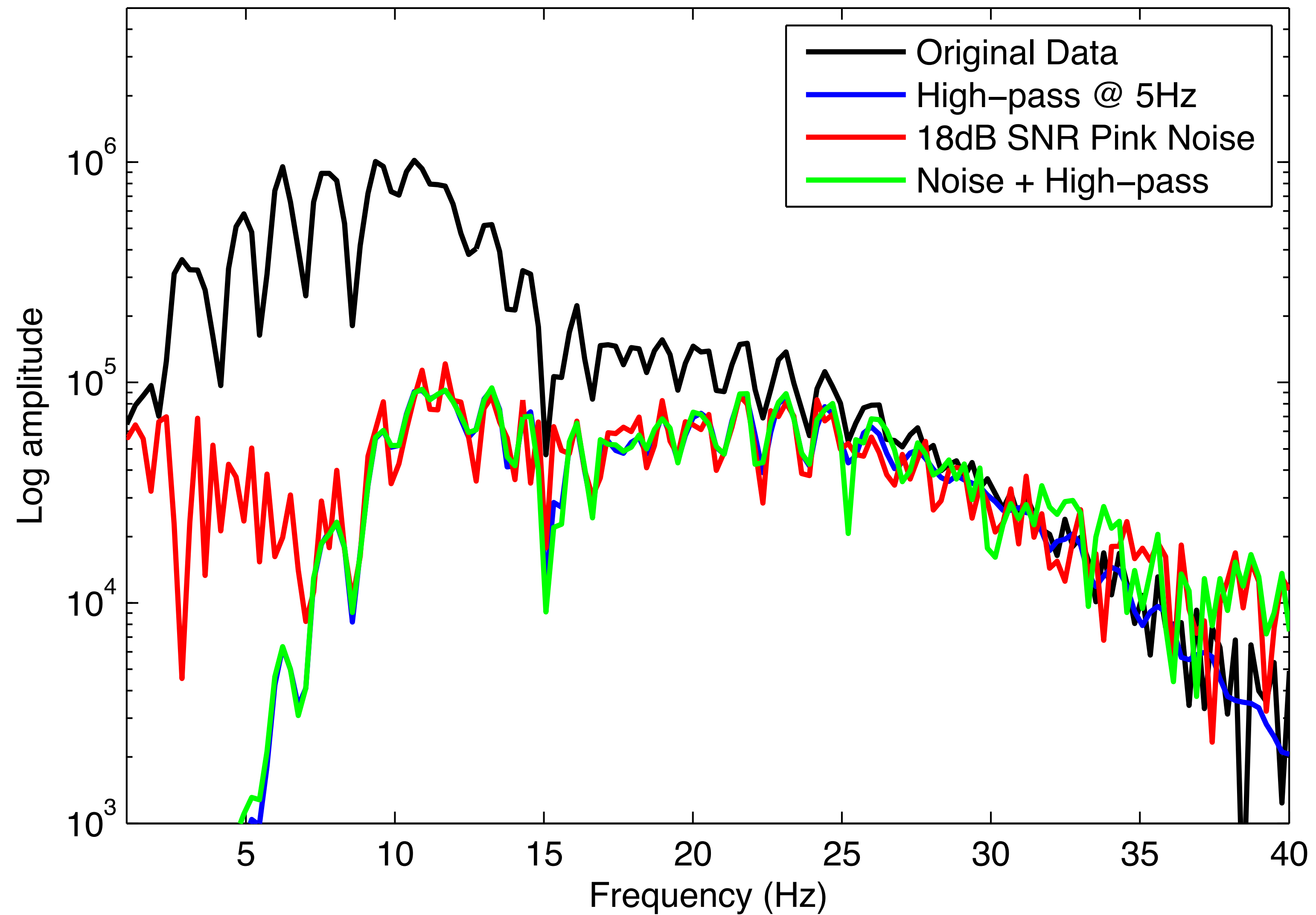
Average frequency spectrum (all traces)





# Average trace spectrum

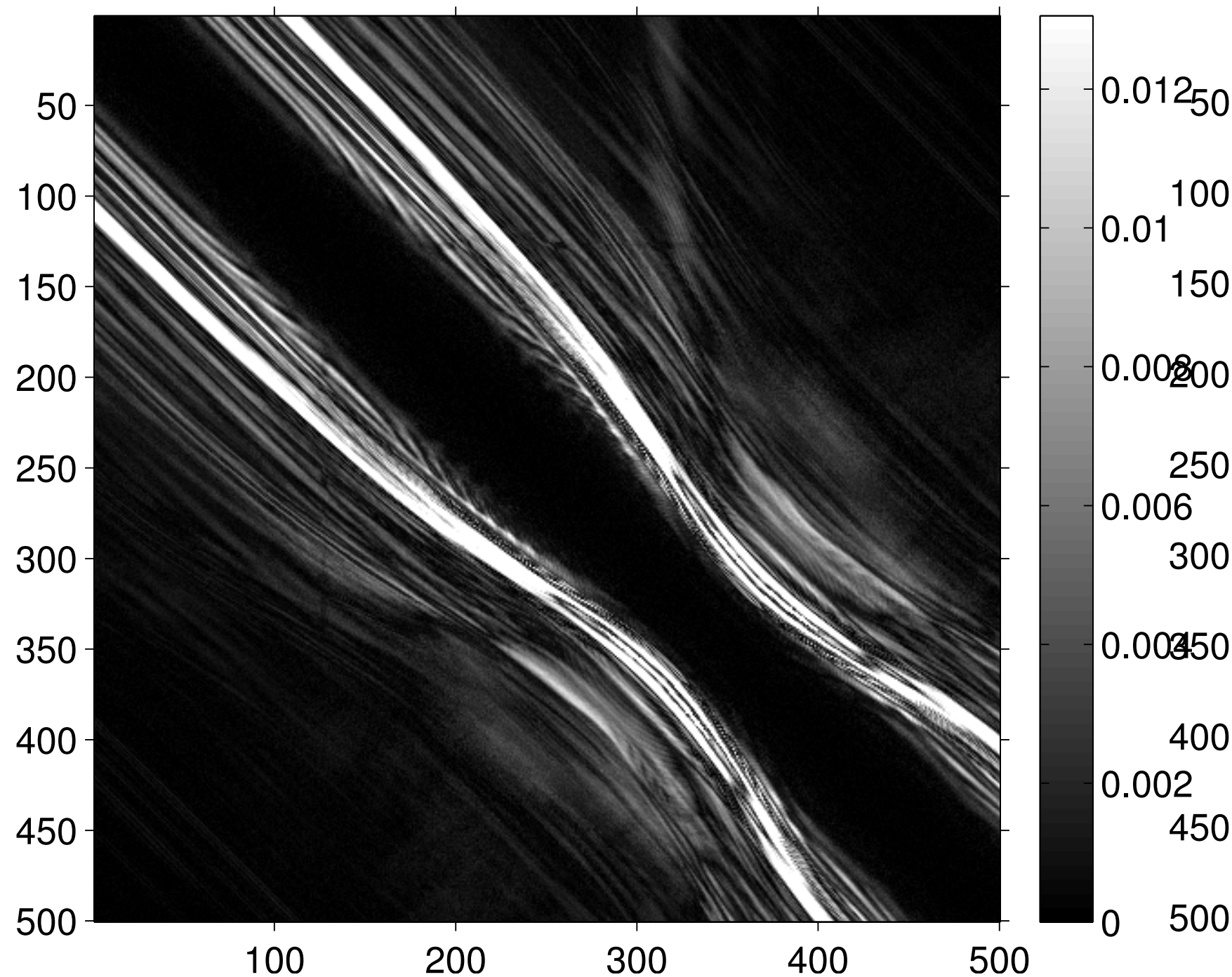
Average frequency spectrum (all traces)





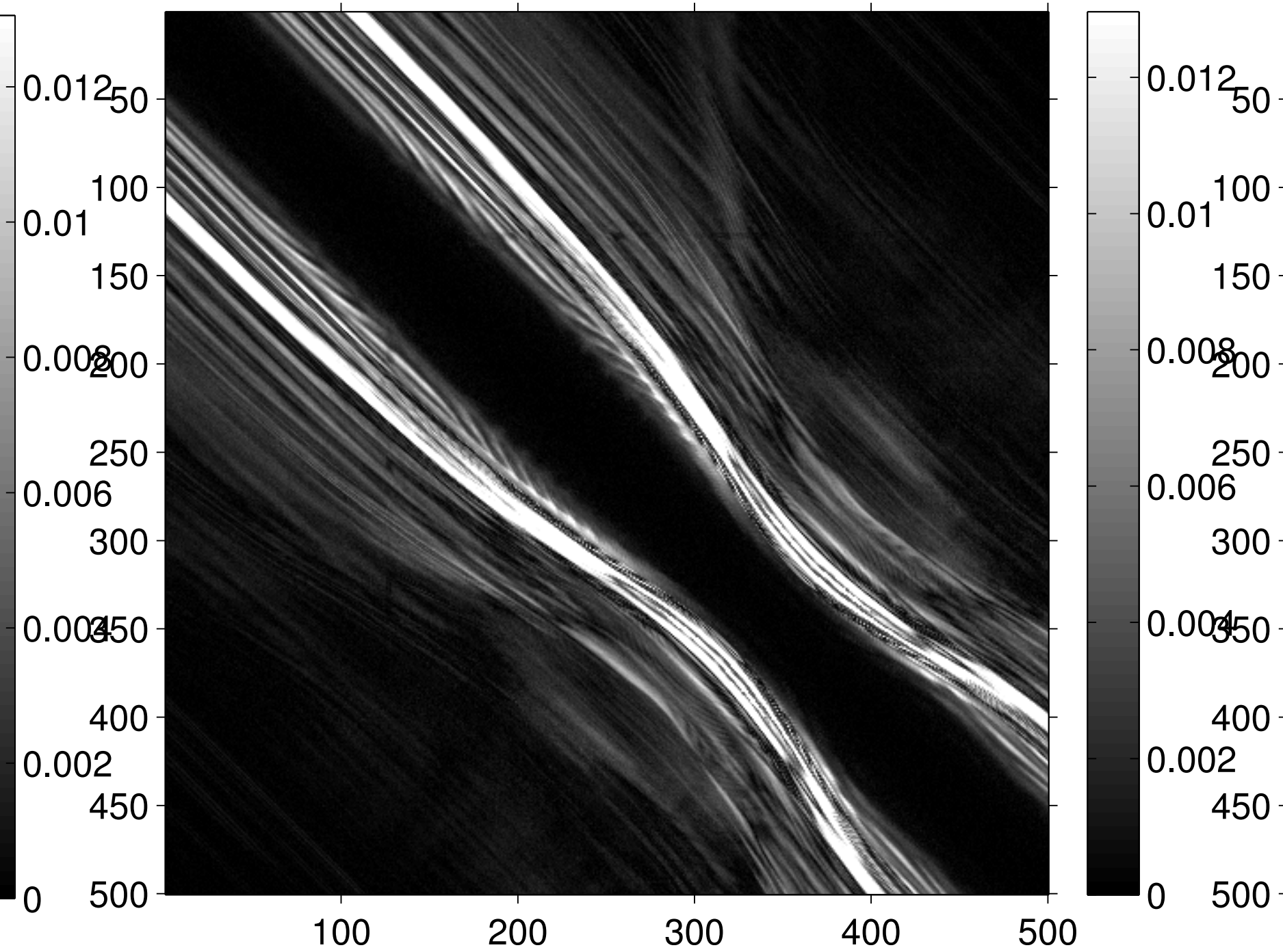
# Amplitude 18dB Pink noise added, Low-cut at 5 Hz

0.5Hz, 18dB SNR (pink noise), High-passed @ 5Hz solution



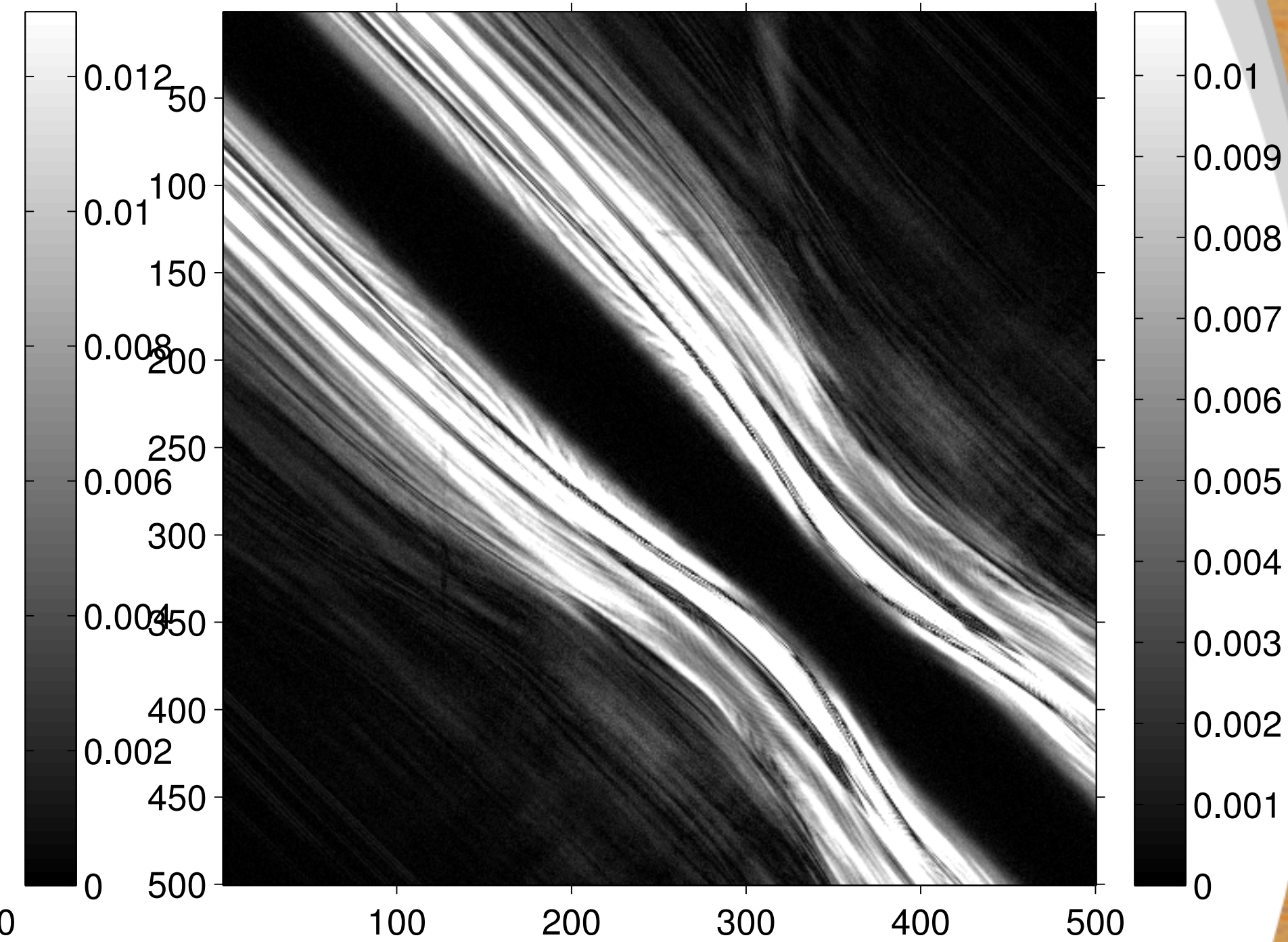
**0.5Hz**

1Hz, 18dB SNR (pink noise), High-passed @ 5Hz solution



**1 Hz**

3Hz, 18dB SNR (pink noise), High-passed @ 5Hz solution

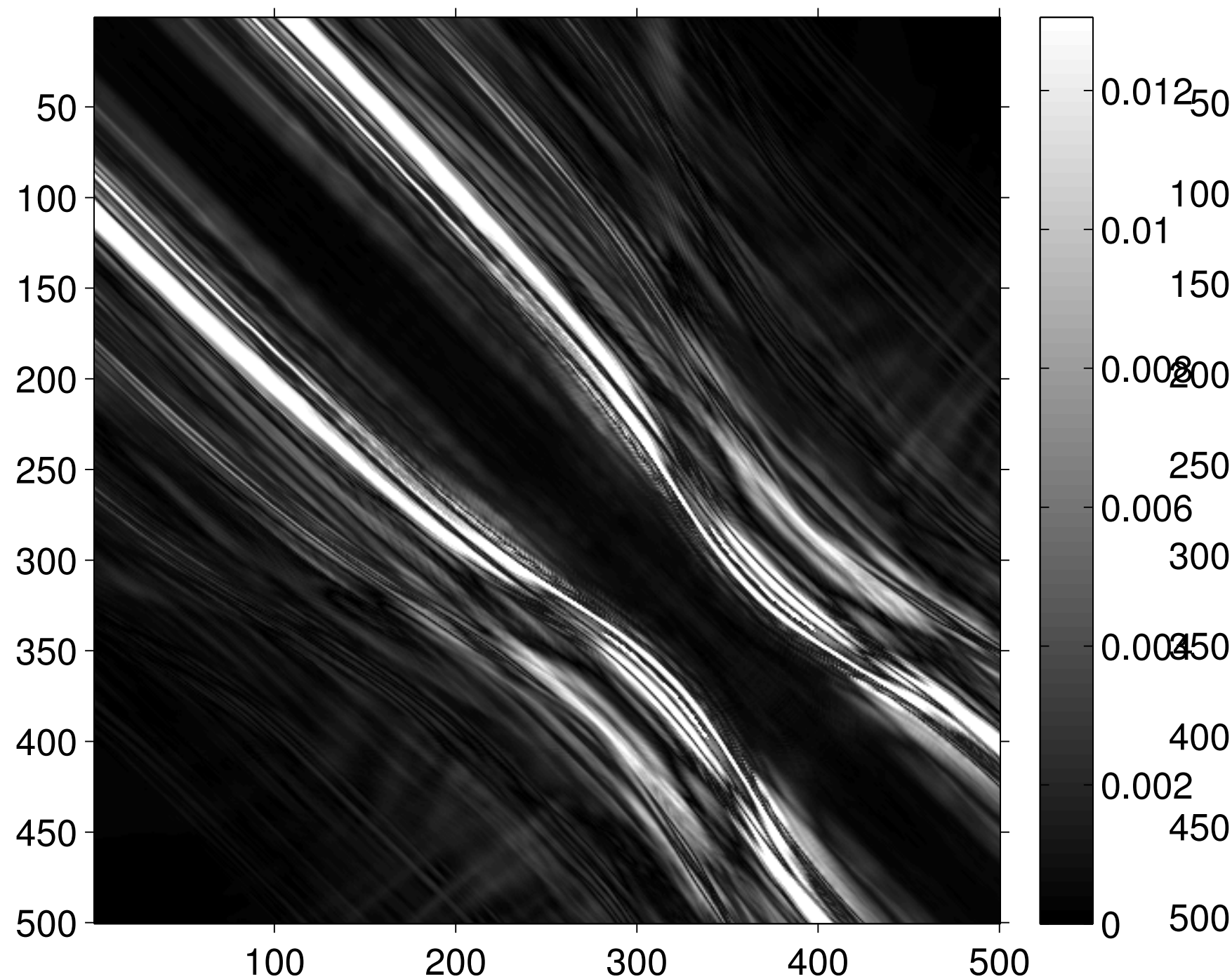


**3 Hz**



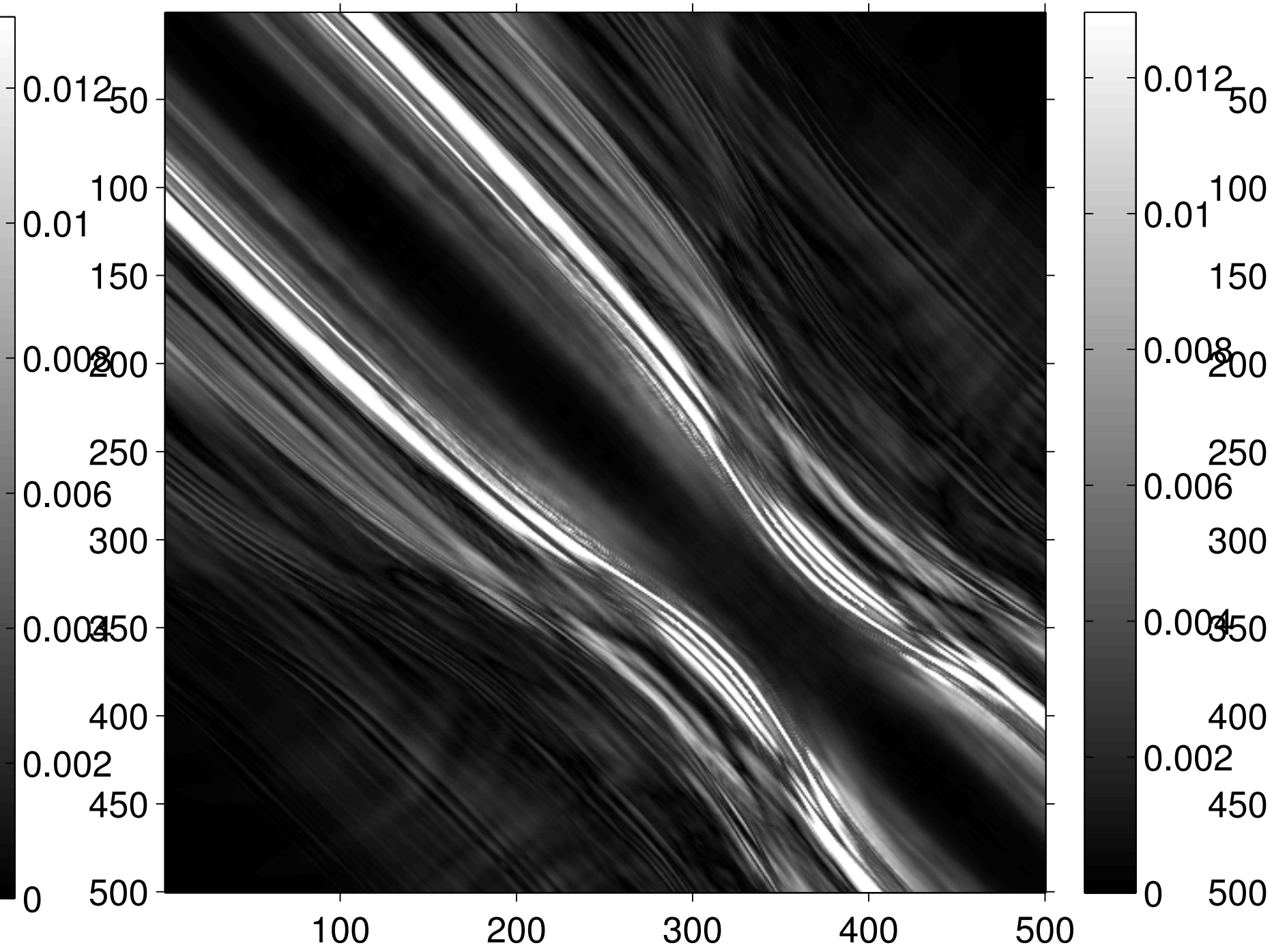
# Amplitude Reference solution

0.5Hz, Reference solution



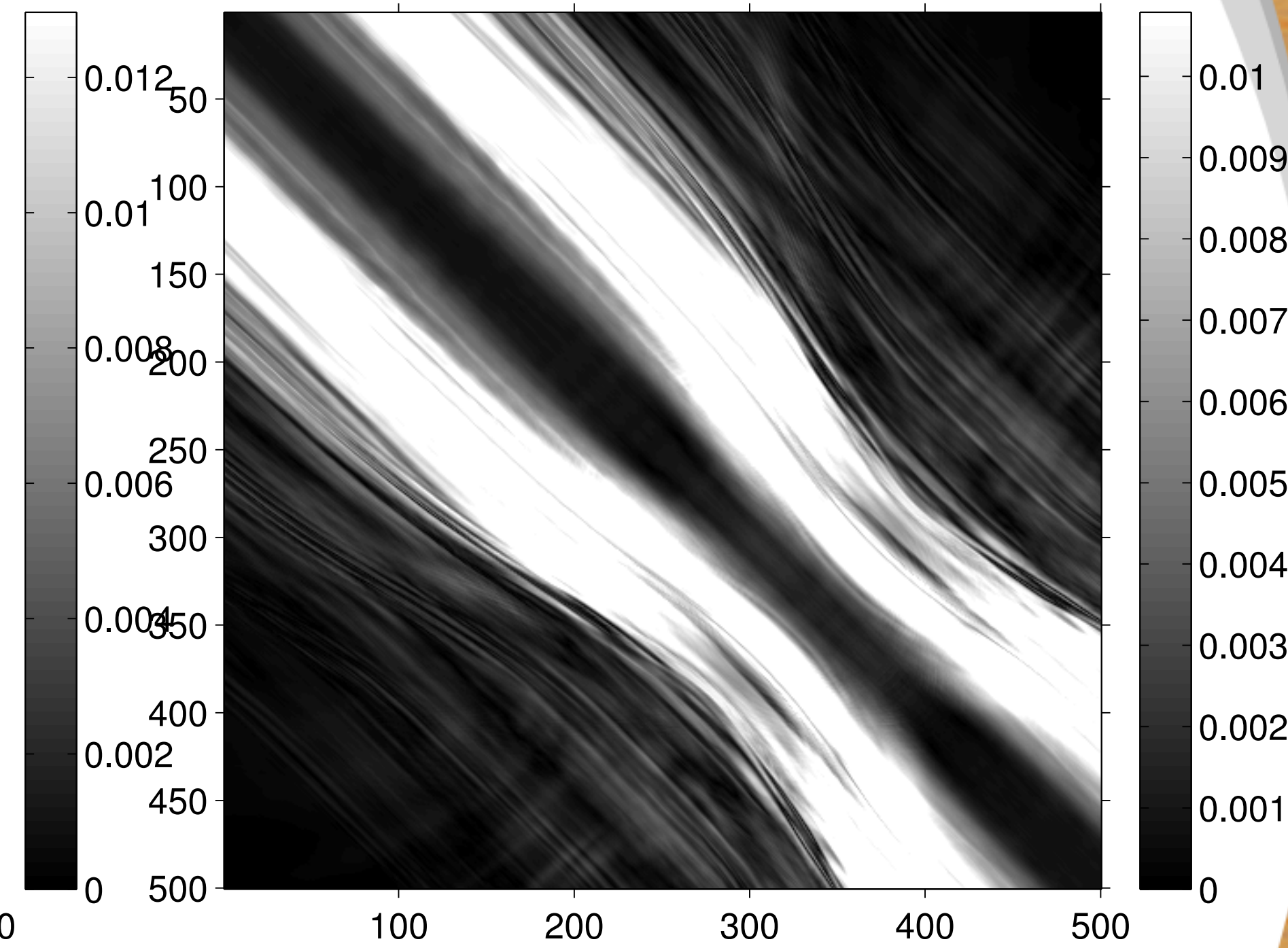
**0.5Hz**

1Hz, Reference solution



**1 Hz**

3Hz, Reference solution

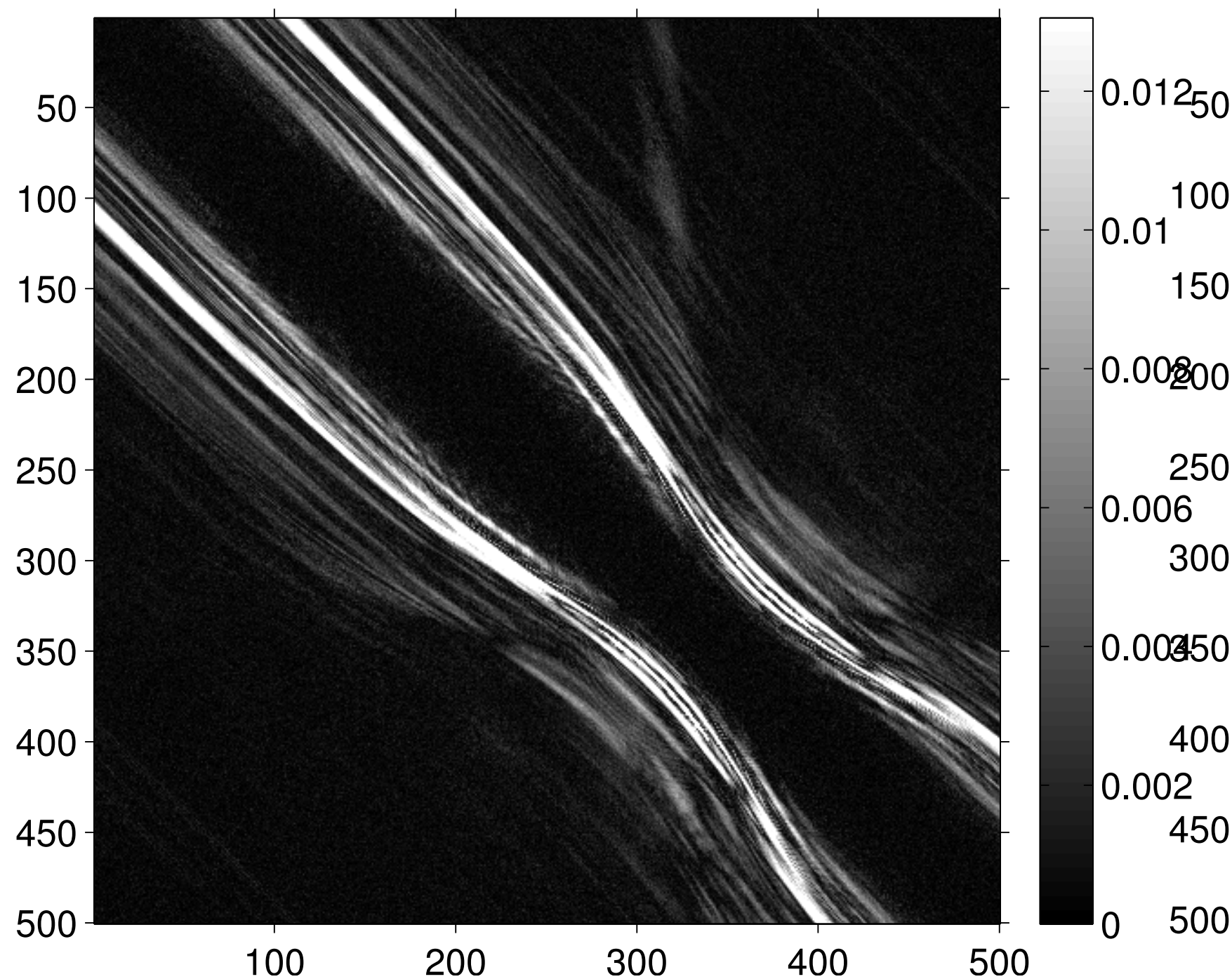


**3 Hz**



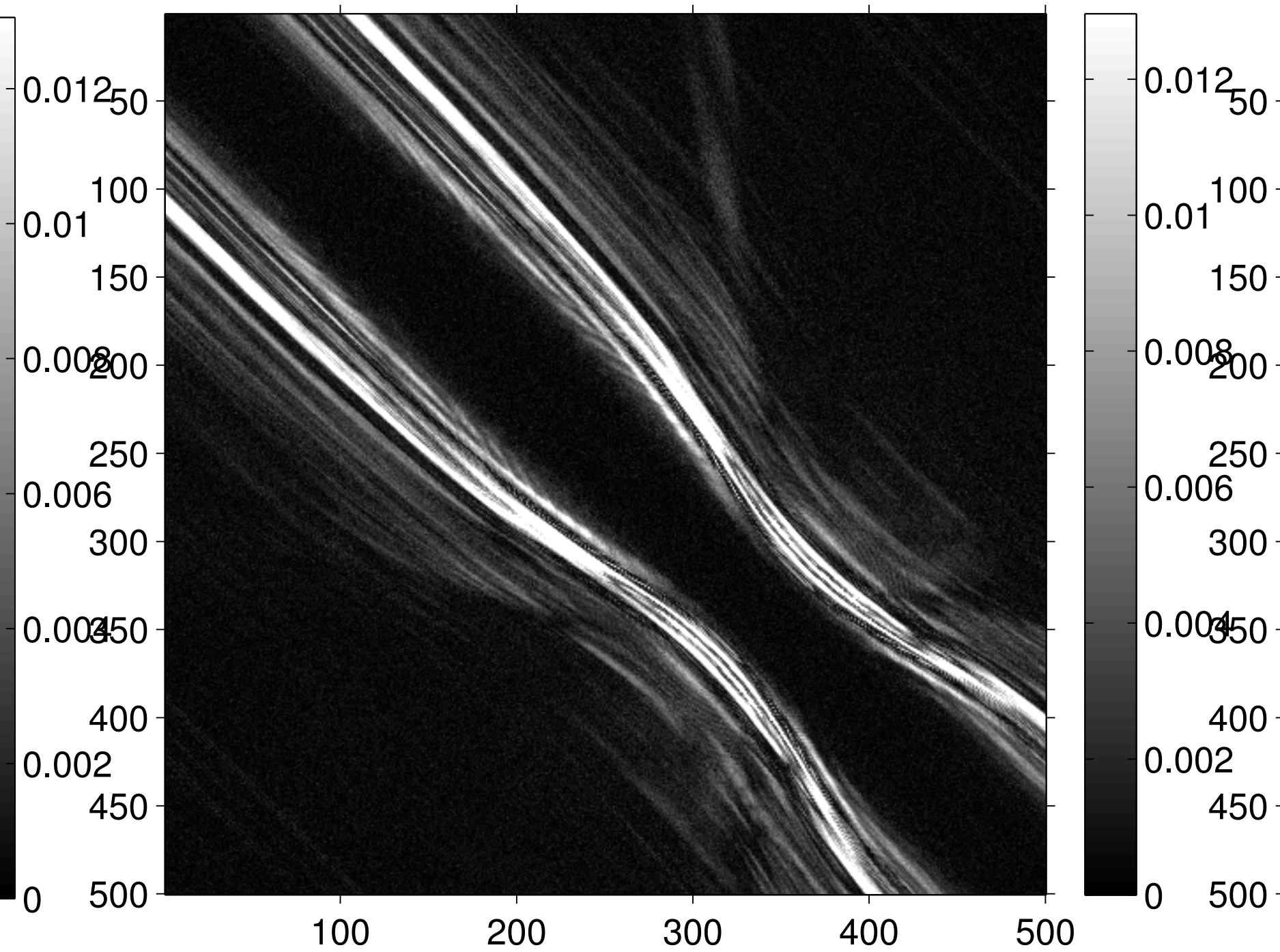
# Amplitude 8dB Pink noise added, Low-cut at 5 Hz

0.5Hz, 8dB SNR (pink noise), High-passed @ 5Hz solution



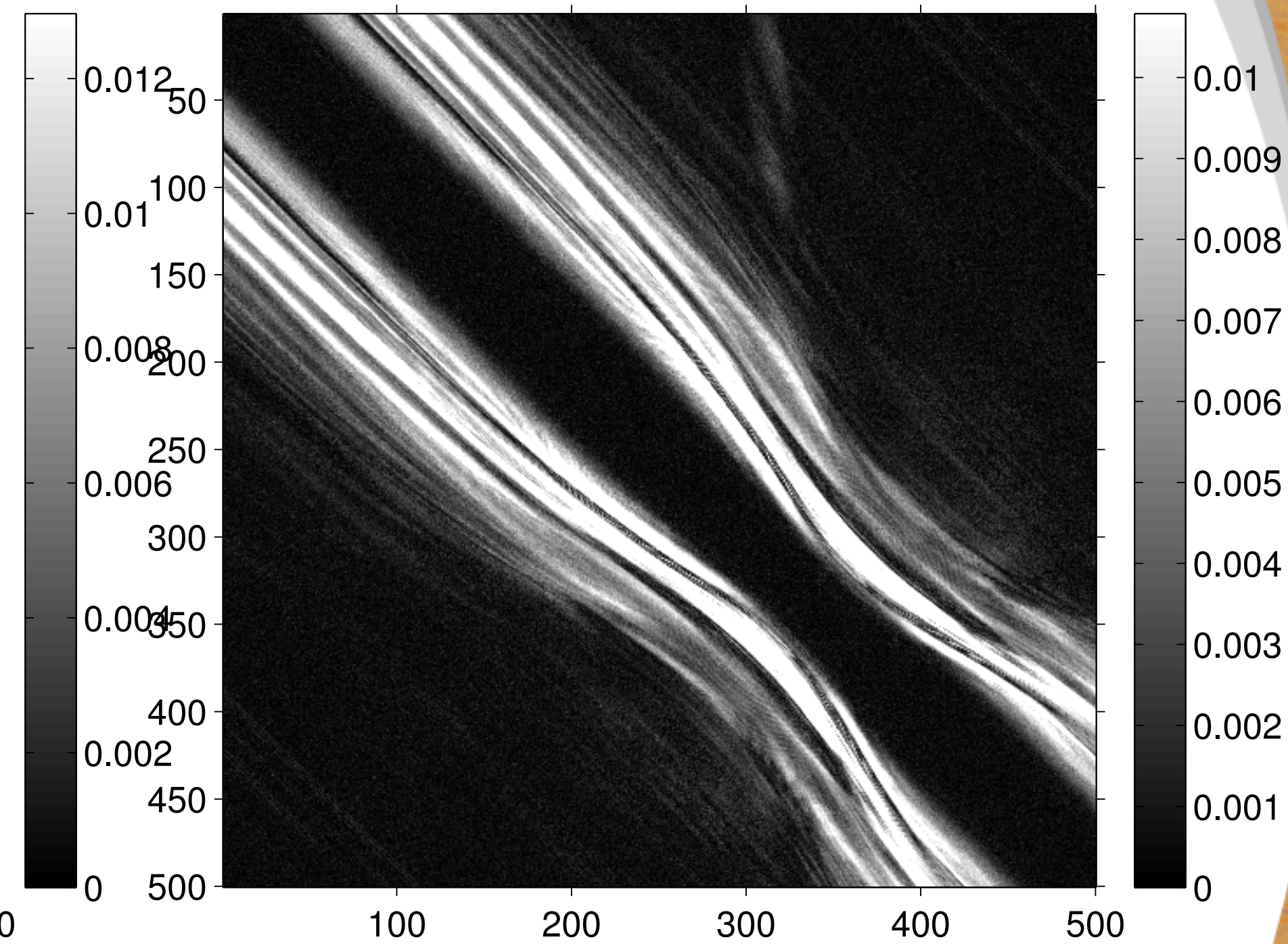
**0.5Hz**

1Hz, 8dB SNR (pink noise), High-passed @ 5Hz solution



**1 Hz**

3Hz, 8dB SNR (pink noise), High-passed @ 5Hz solution

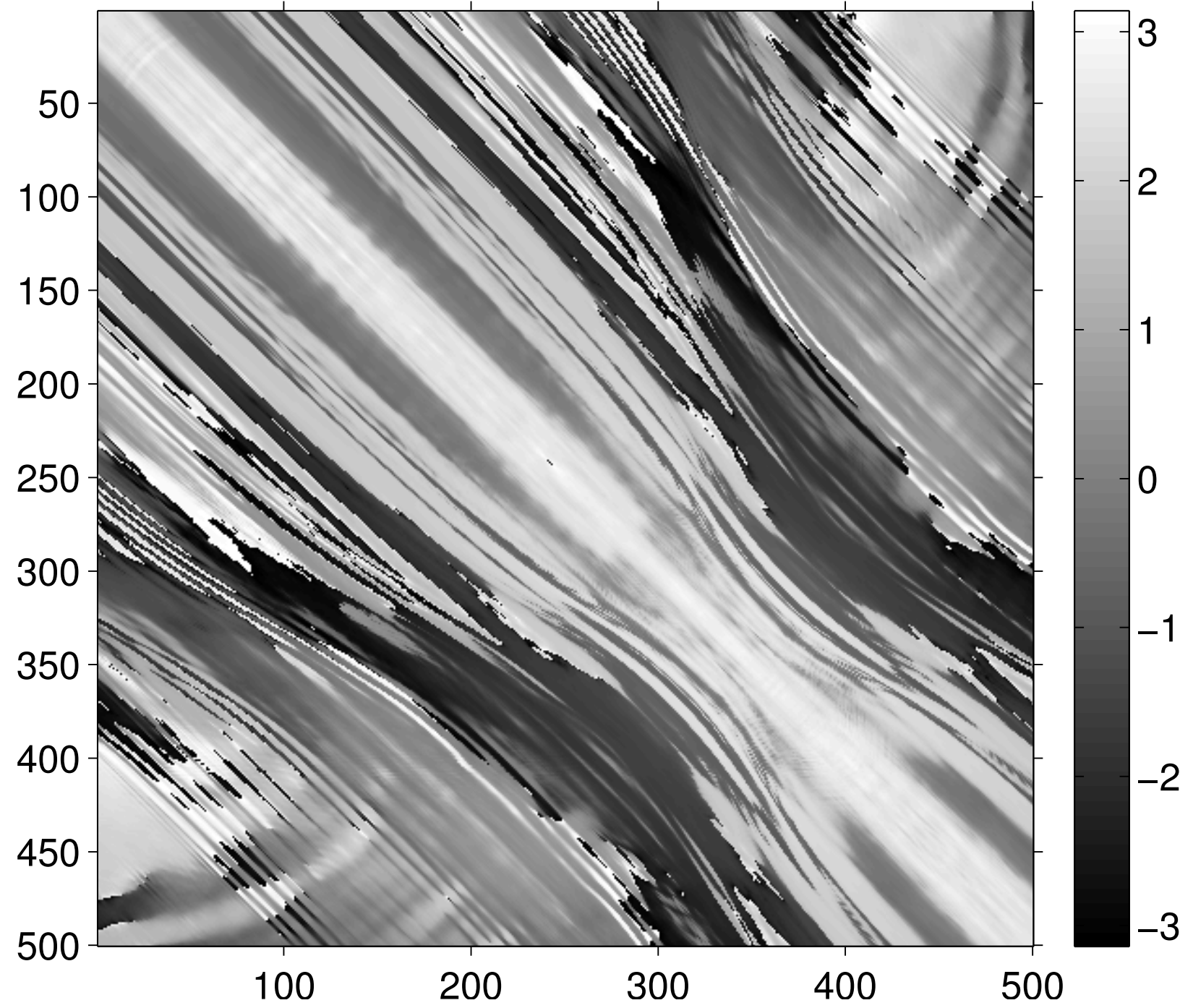


**3 Hz**



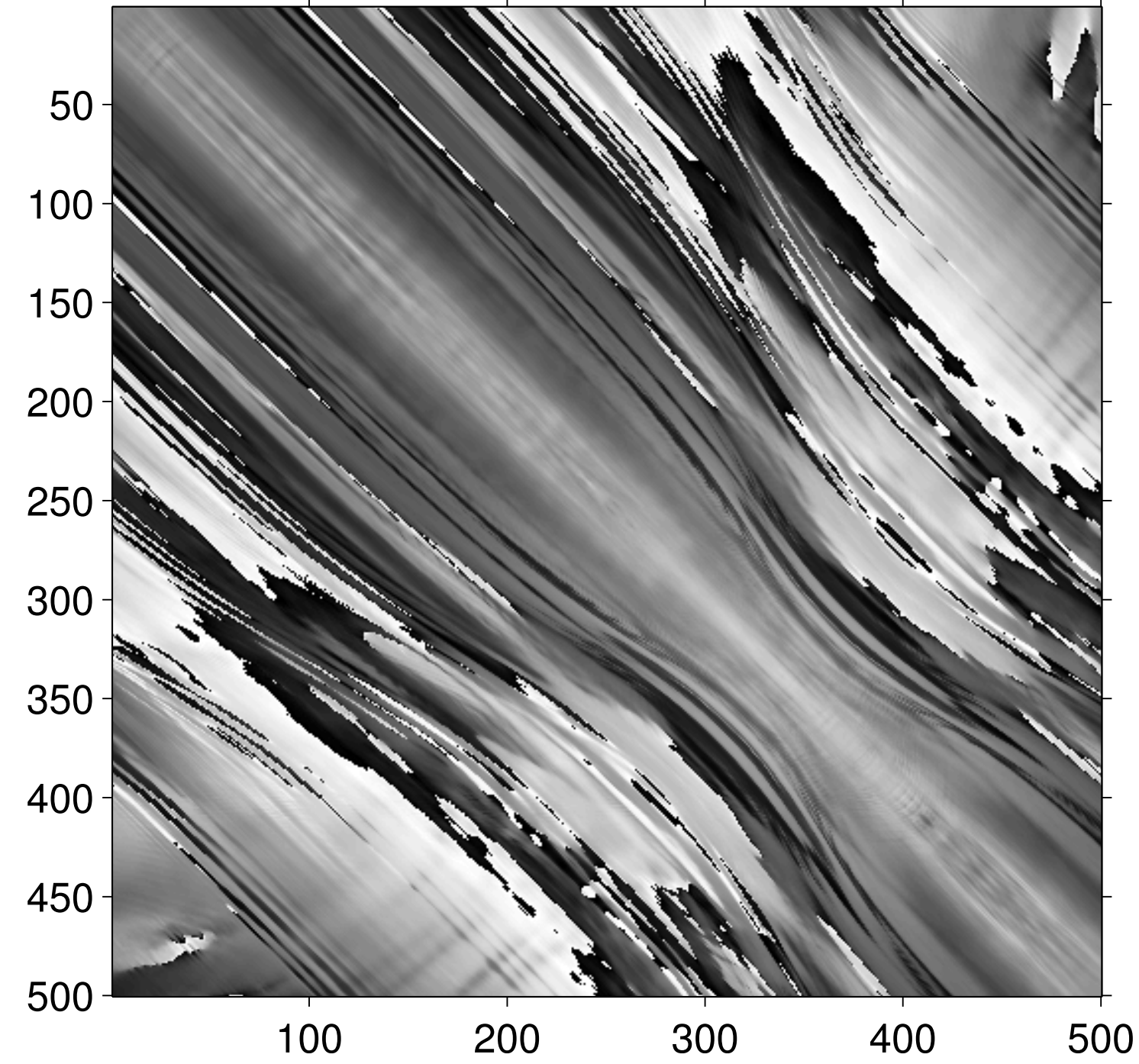
# Phase Reference solution

0.5Hz, Reference solution



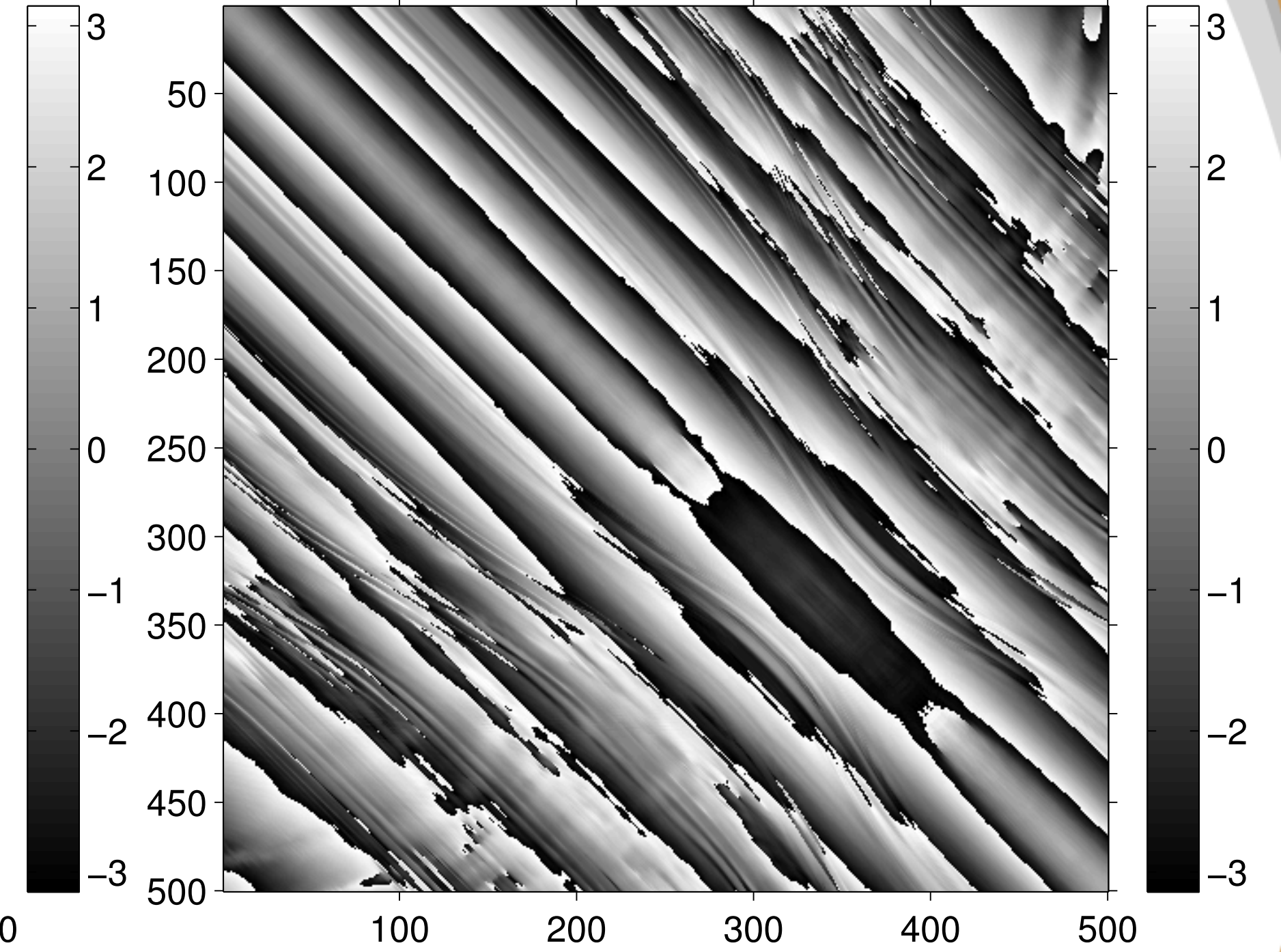
**0.5Hz**

1Hz, Reference solution



**1 Hz**

3Hz, Reference solution

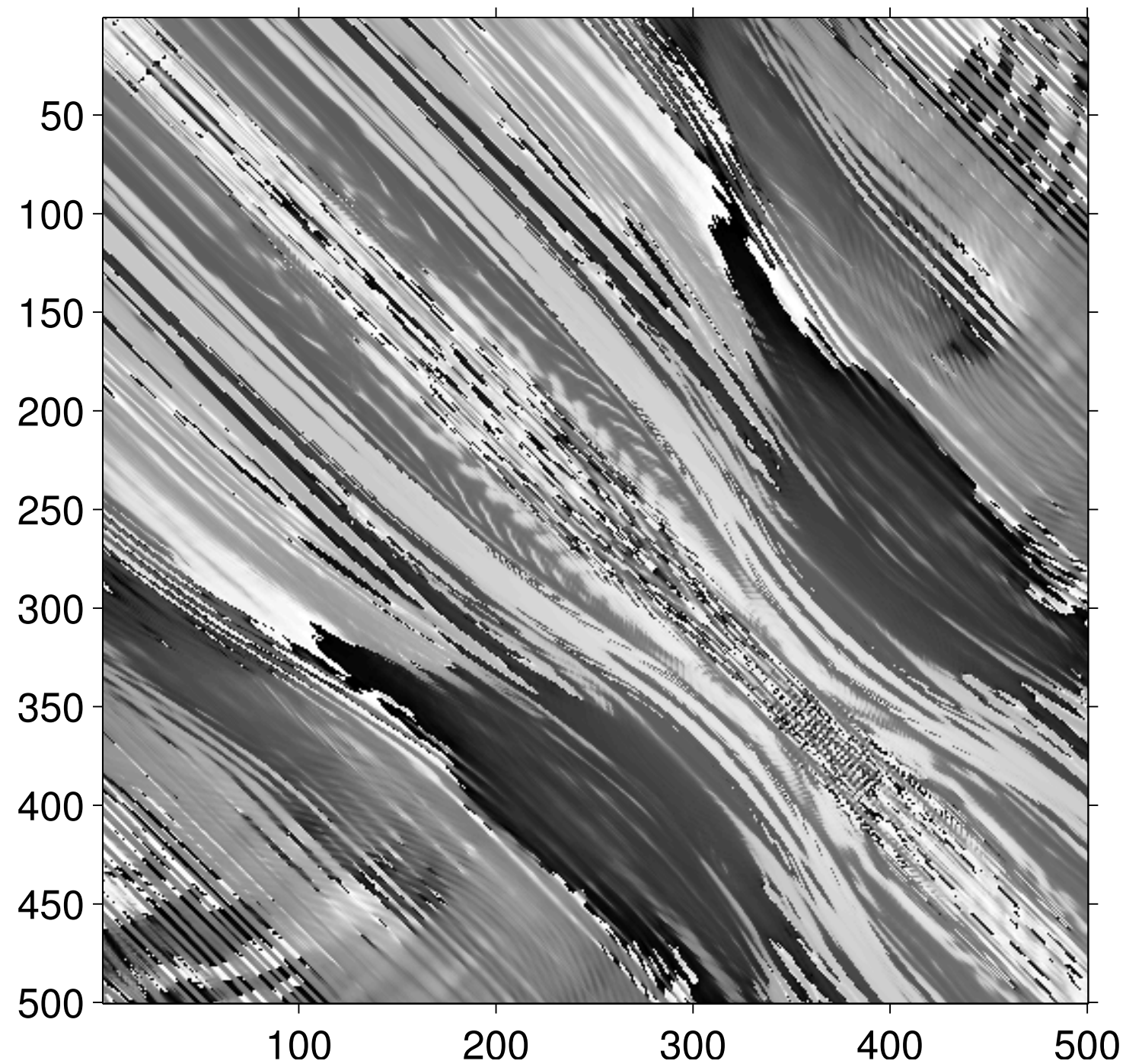


**3 Hz**



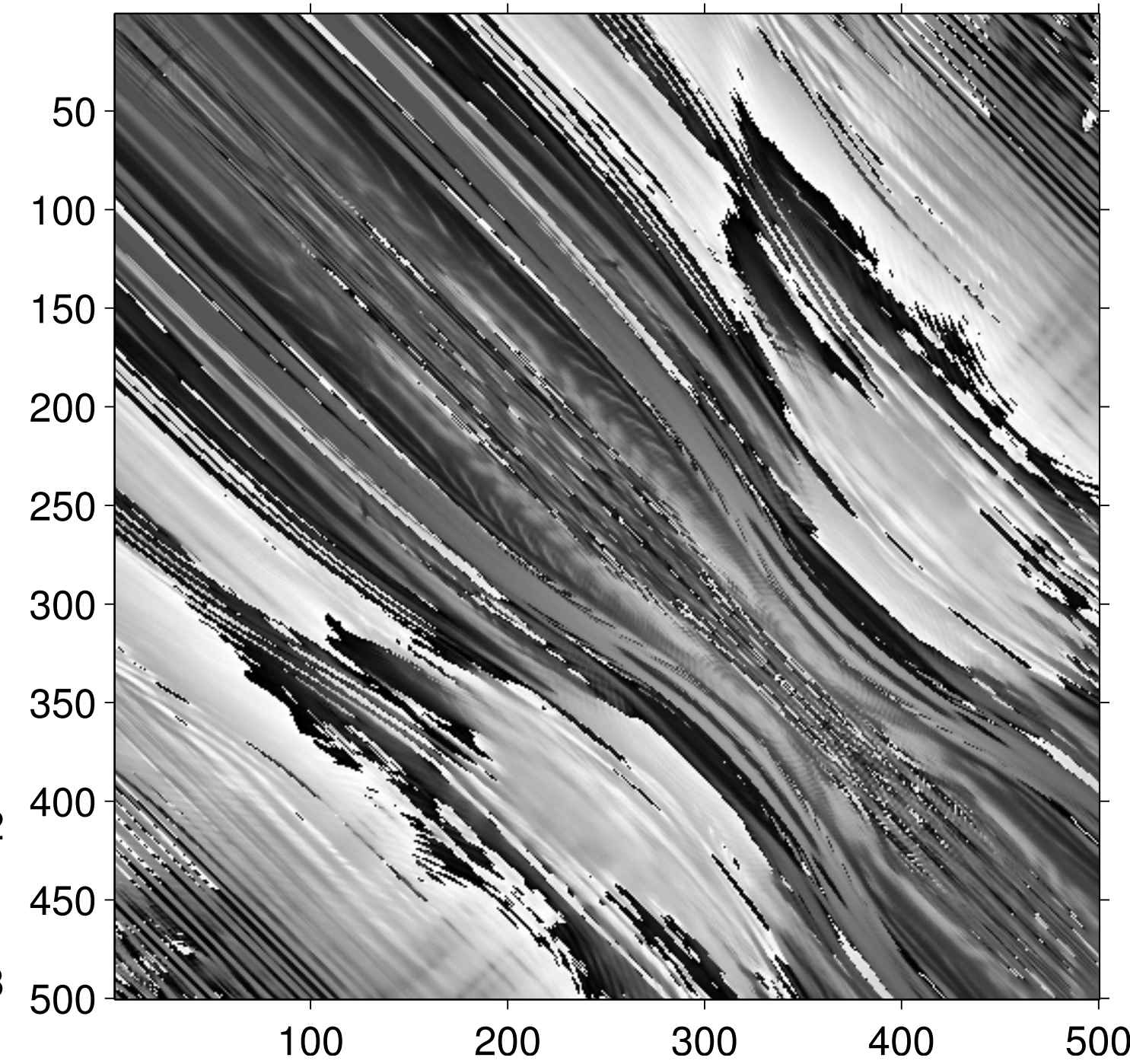
# Phase Low-cut at 5 Hz

0.5Hz, High-passed @ 5Hz solution



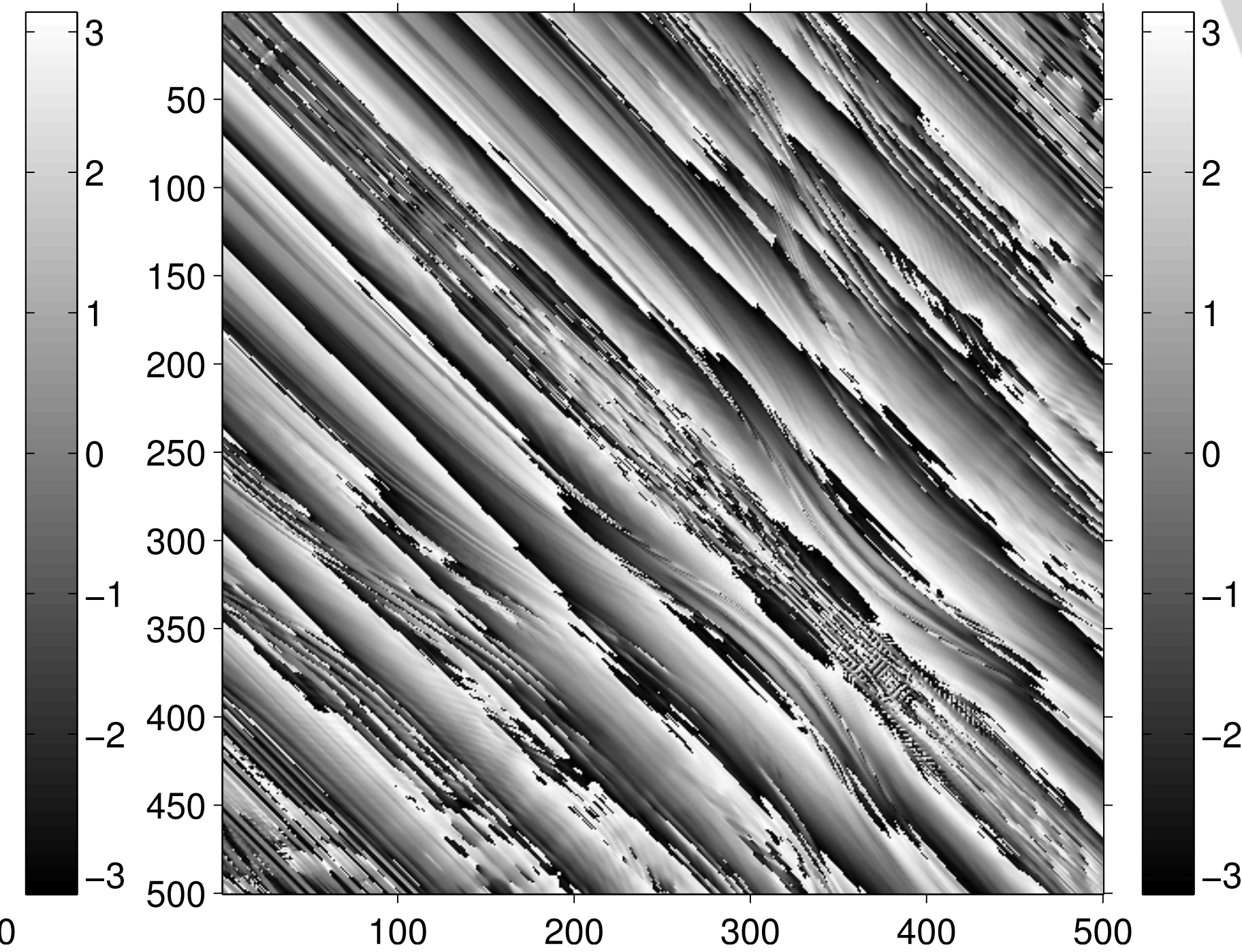
**0.5Hz**

1Hz, High-passed @ 5Hz solution



**1 Hz**

3Hz, High-passed @ 5Hz solution

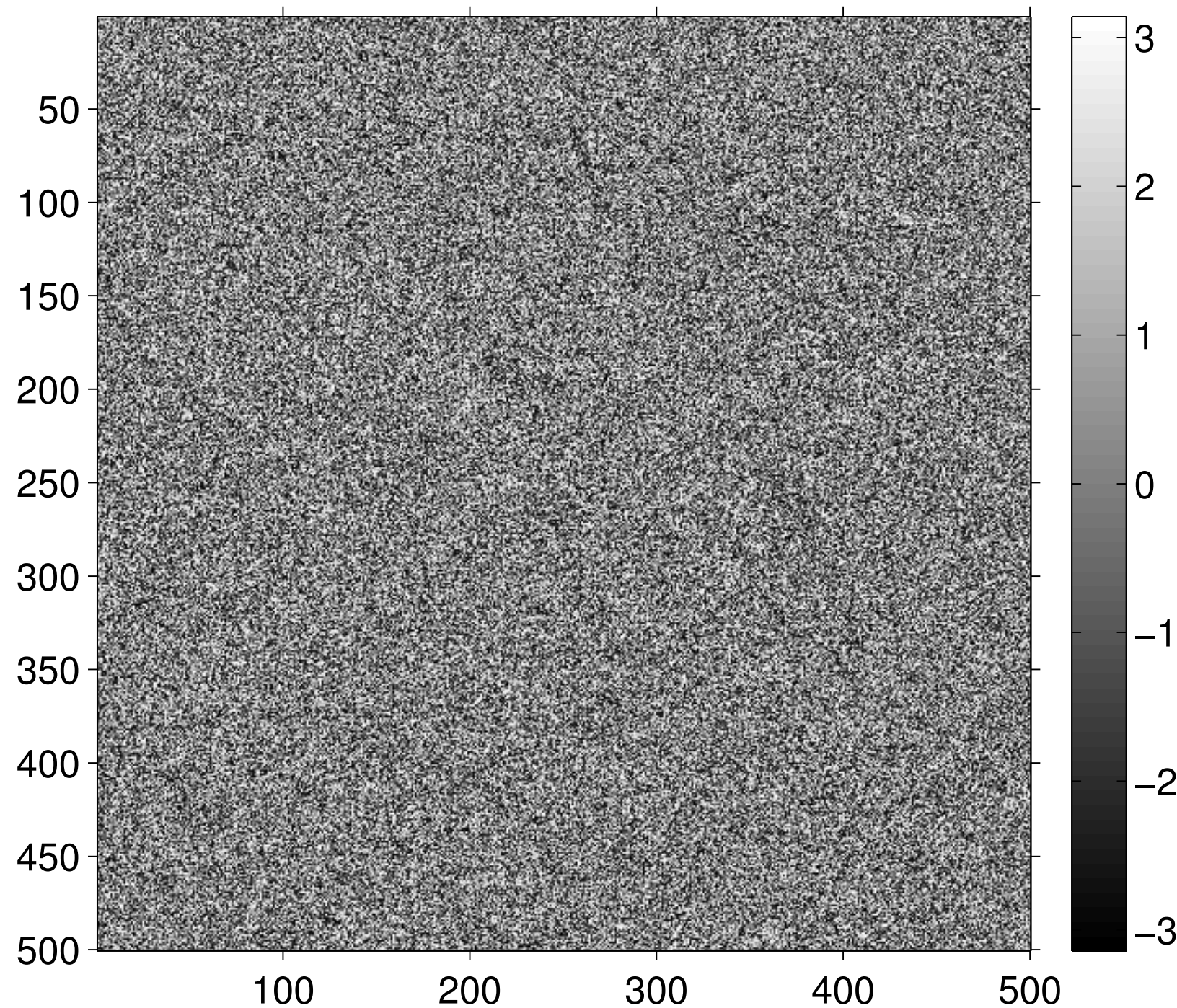


**3 Hz**



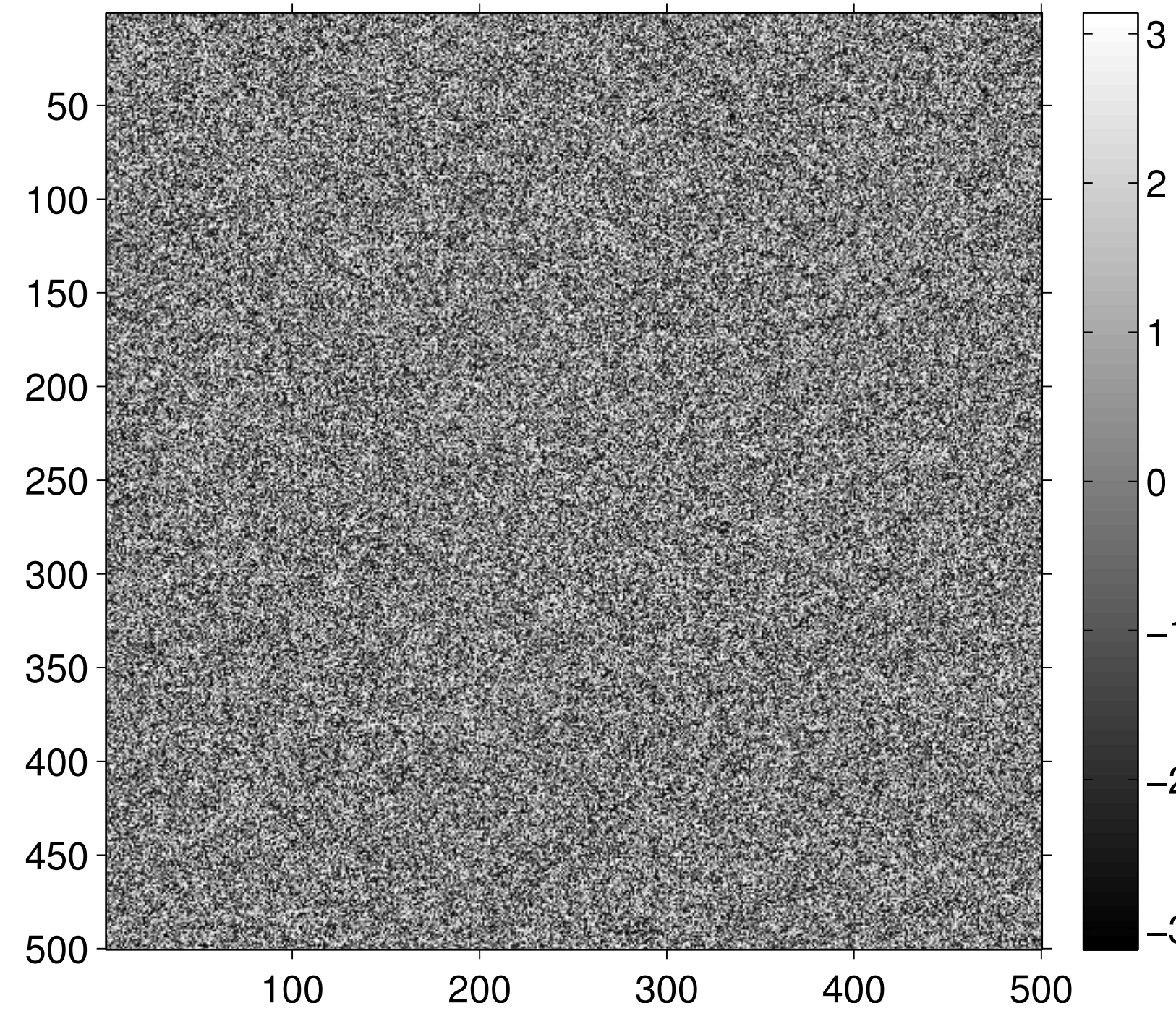
# Phase Low-cut at 5 Hz, 18dB Pink noise added

0.5Hz, High-passed @ 5Hz, 18dB SNR (pink noise) solution



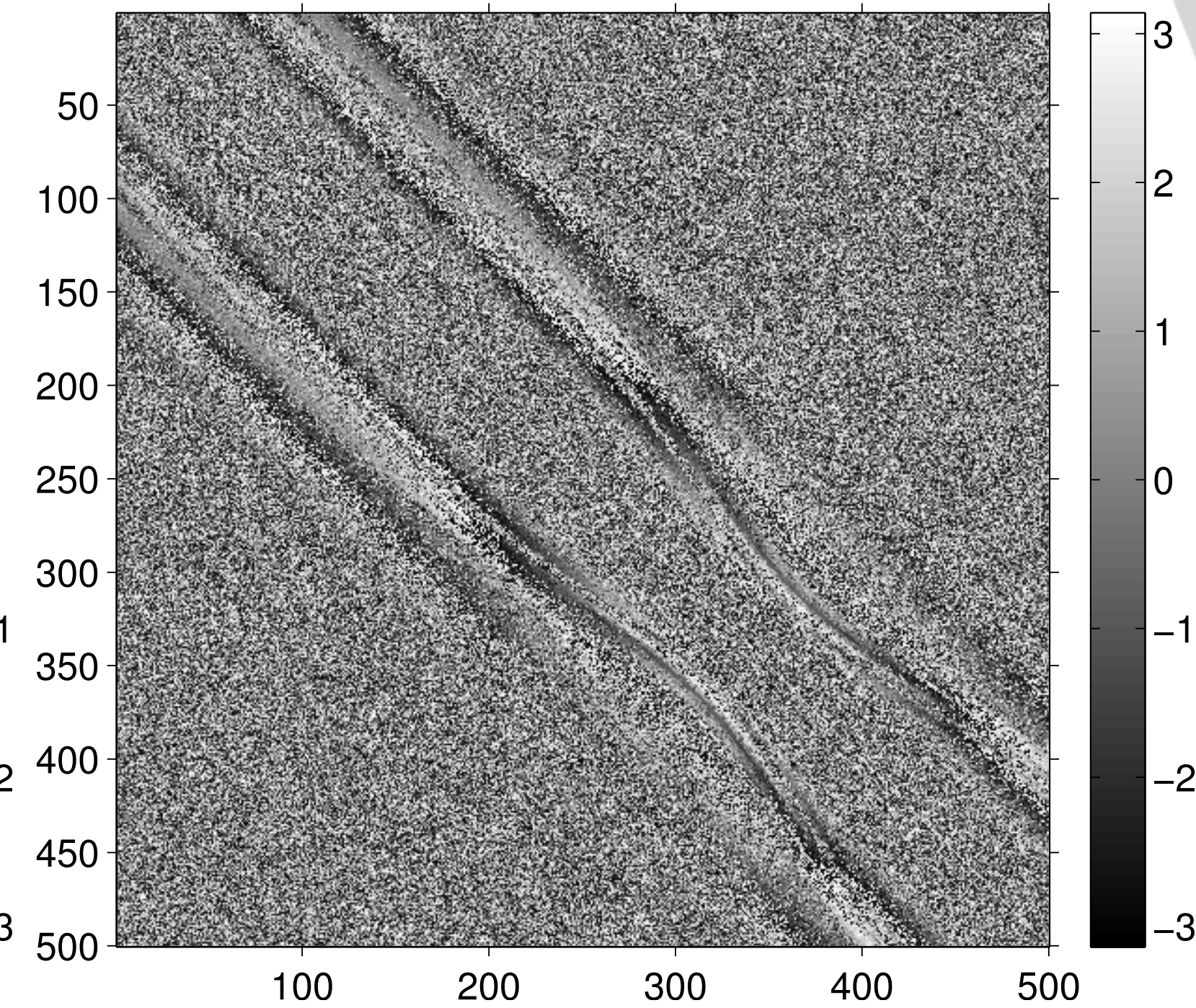
**0.5Hz**

1Hz, High-passed @ 5Hz, 18dB SNR (pink noise) solution



**1 Hz**

3Hz, High-passed @ 5Hz, 18dB SNR (pink noise) solution

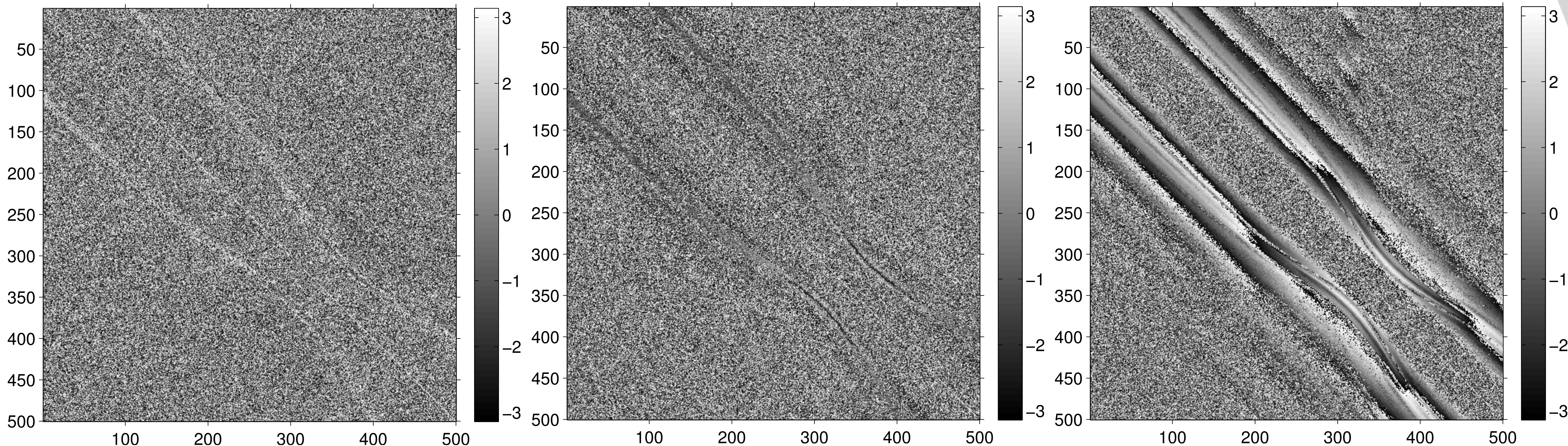


**3 Hz**



# Phase Low-cut at 5 Hz, 18dB Pink noise added (solved to exact sigma)

500 Hz, High-passed @ 5Hz, 18dB SNR (pink noise), exact sigma solution



**0.5Hz**

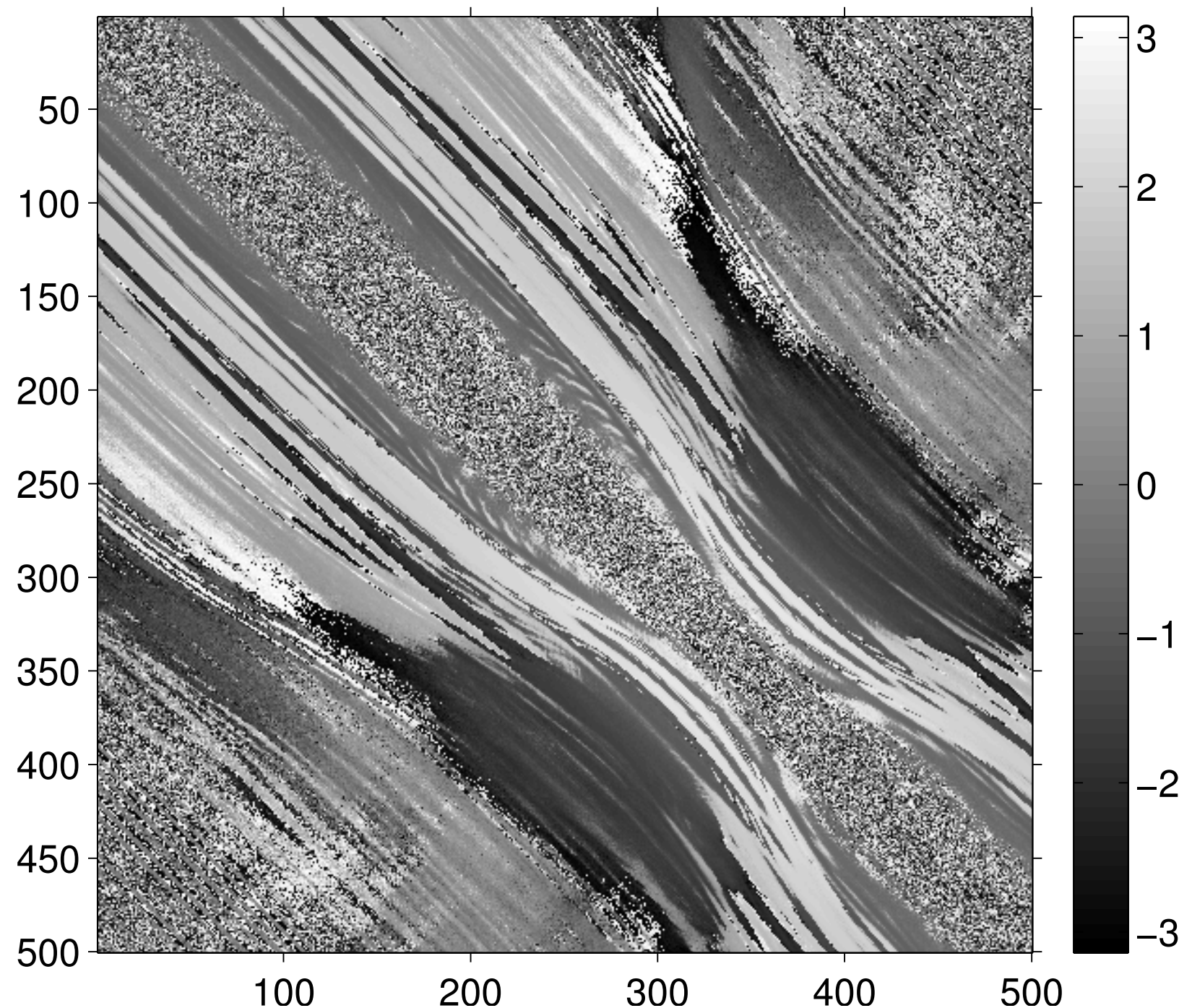
**1 Hz**

**3 Hz**



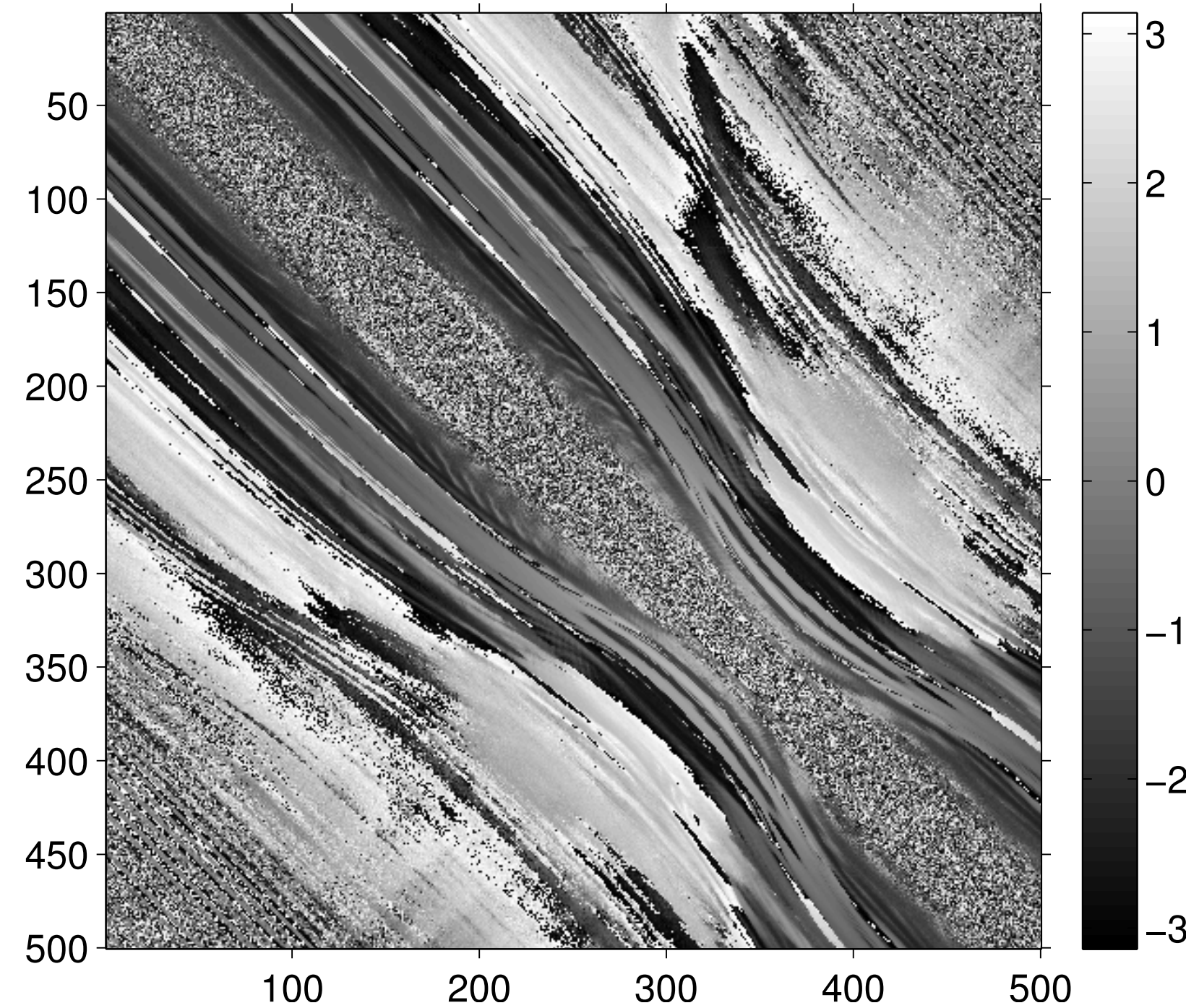
# Phase 18dB Pink noise added, Low-cut at 5 Hz

0.5Hz, 18dB SNR (pink noise), High-passed @ 5Hz solution



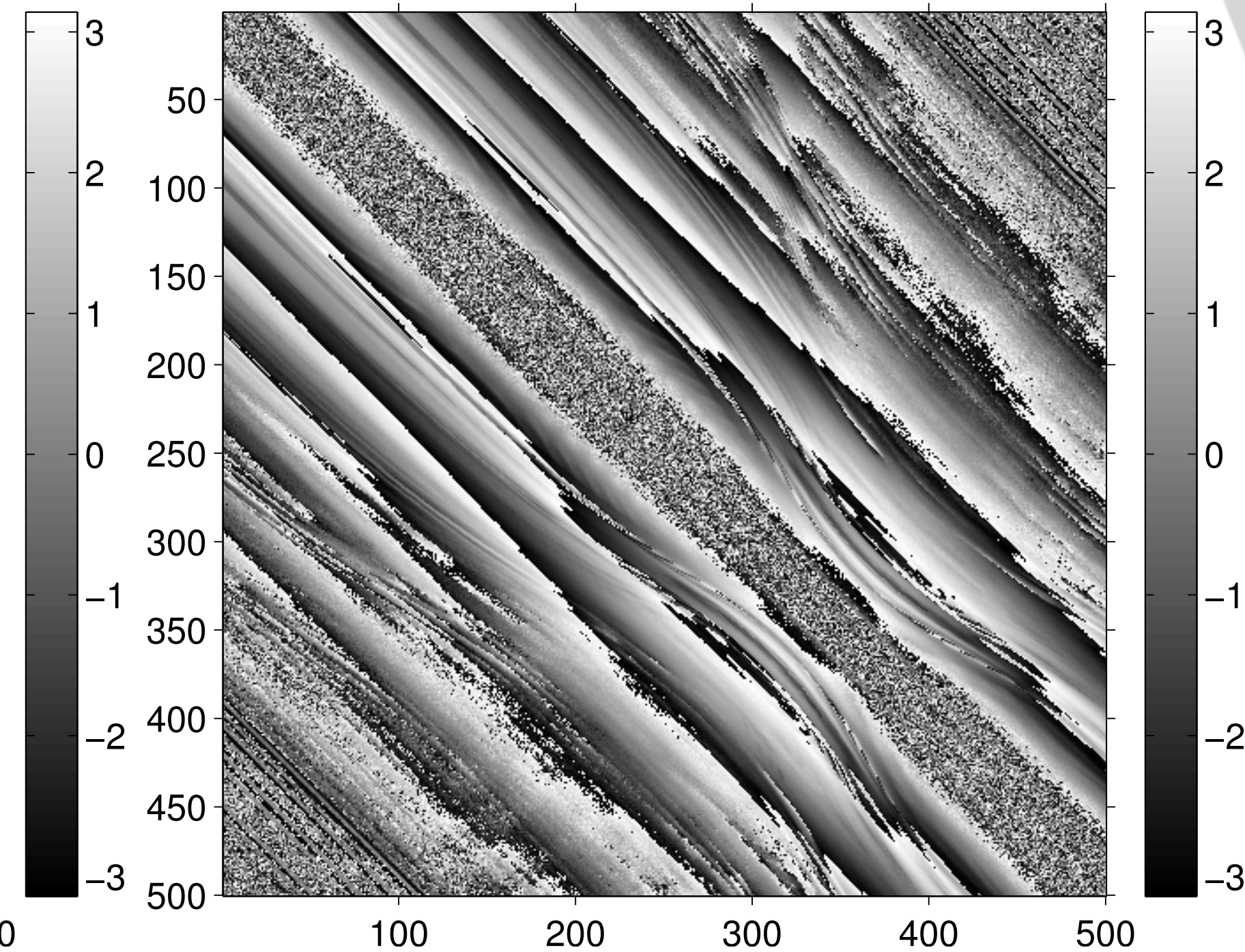
**0.5Hz**

1Hz, 18dB SNR (pink noise), High-passed @ 5Hz solution



**1 Hz**

3Hz, 18dB SNR (pink noise), High-passed @ 5Hz solution

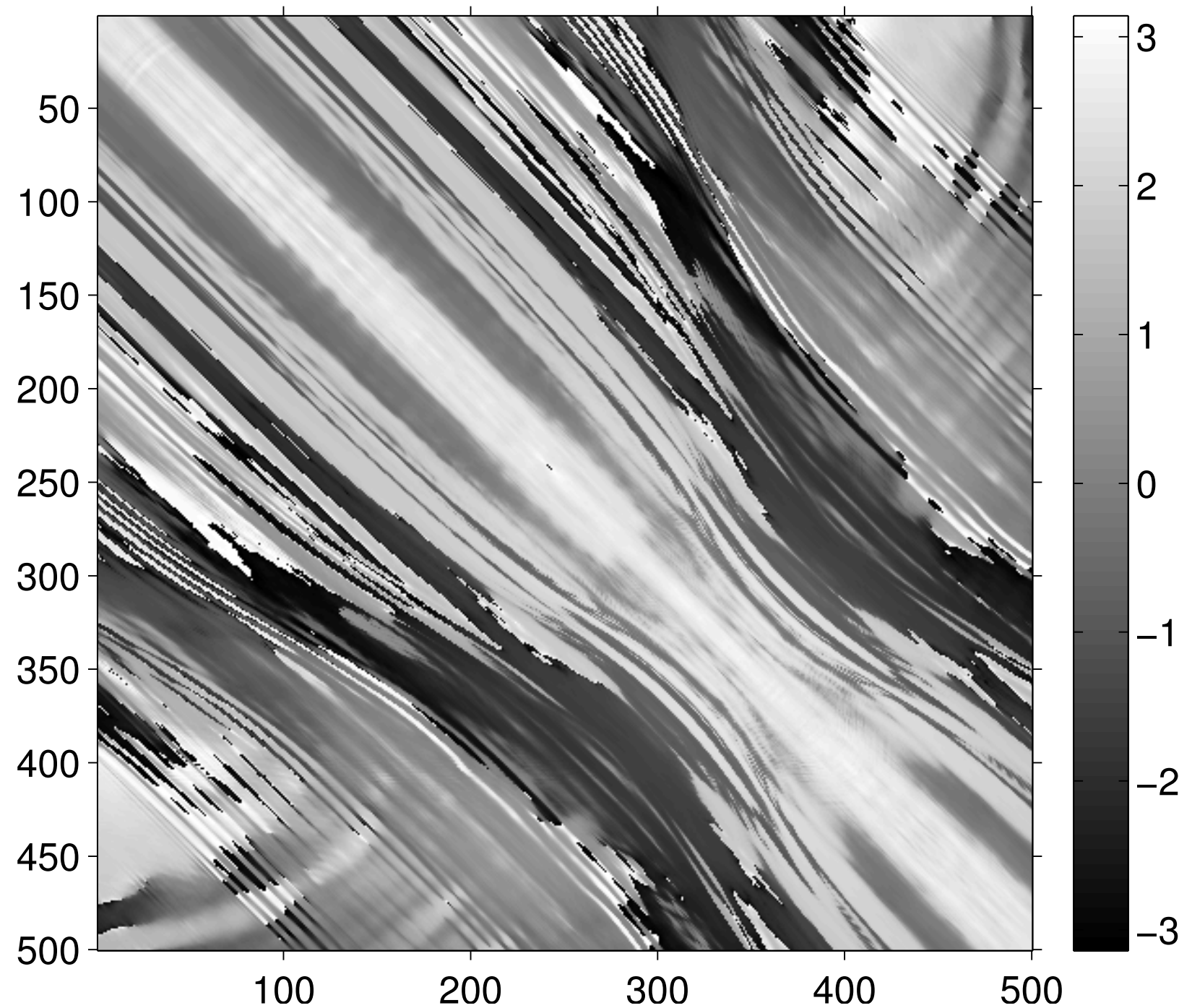


**3 Hz**



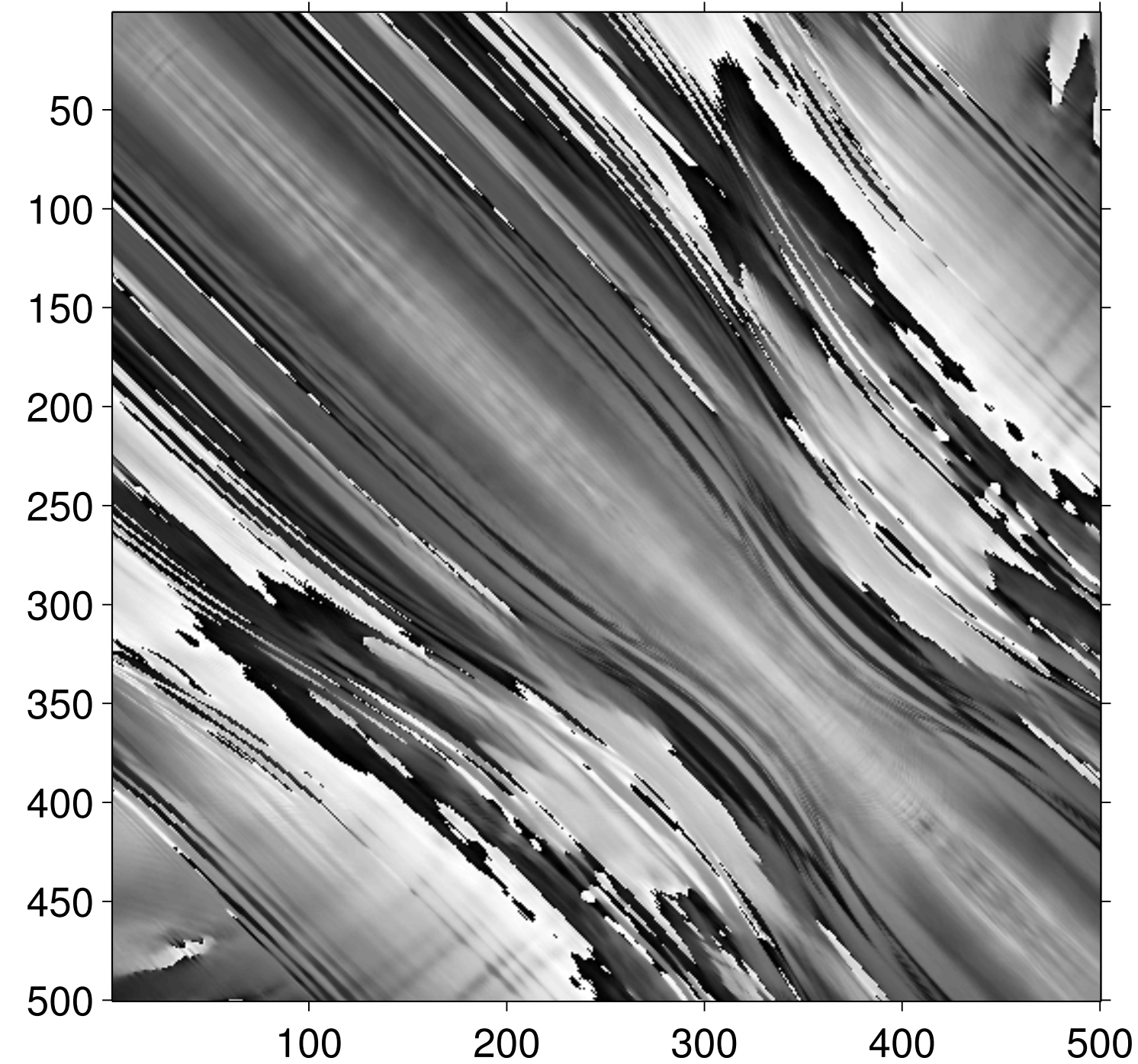
# Phase Reference solution

0.5Hz, Reference solution



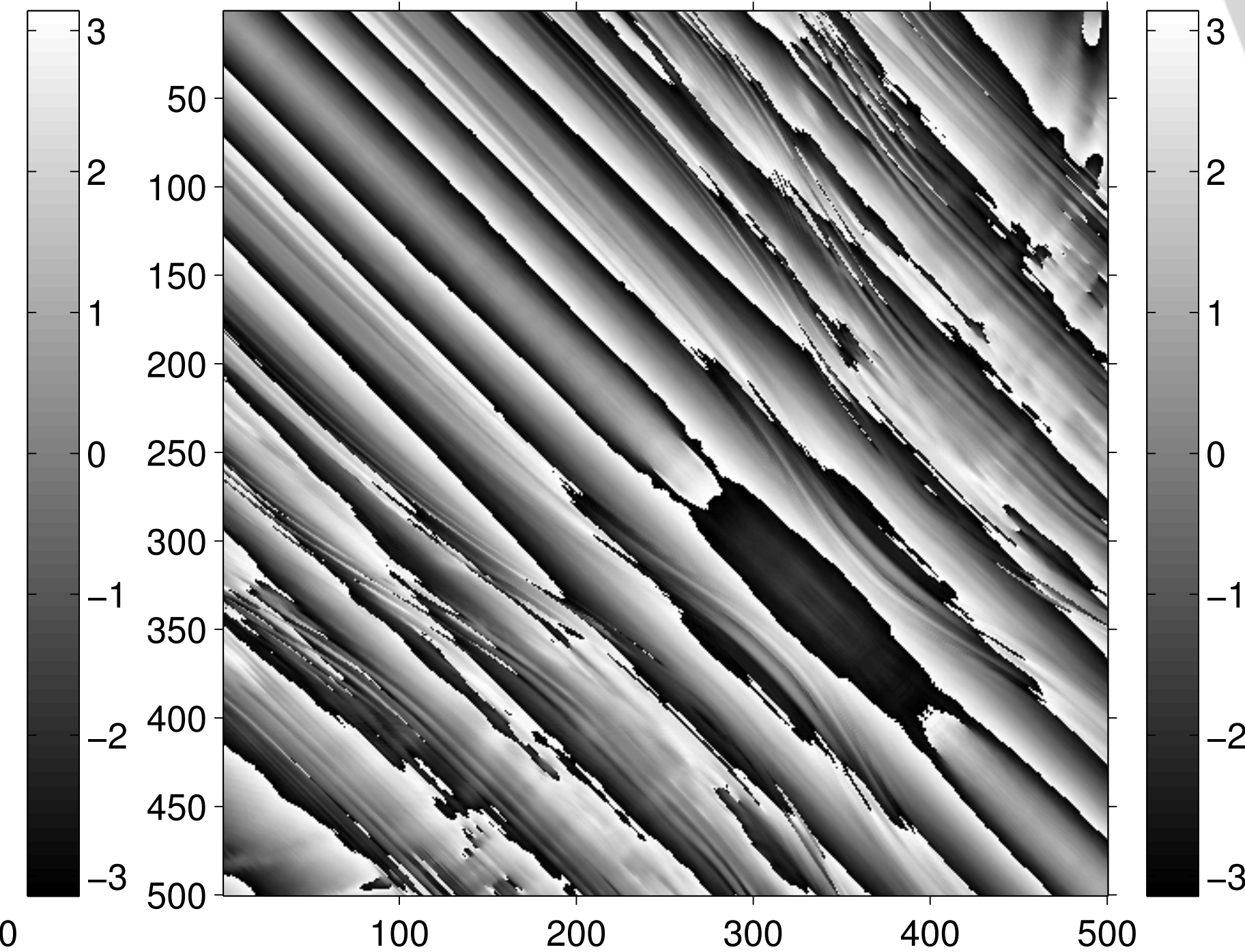
**0.5Hz**

1Hz, Reference solution



**1 Hz**

3Hz, Reference solution

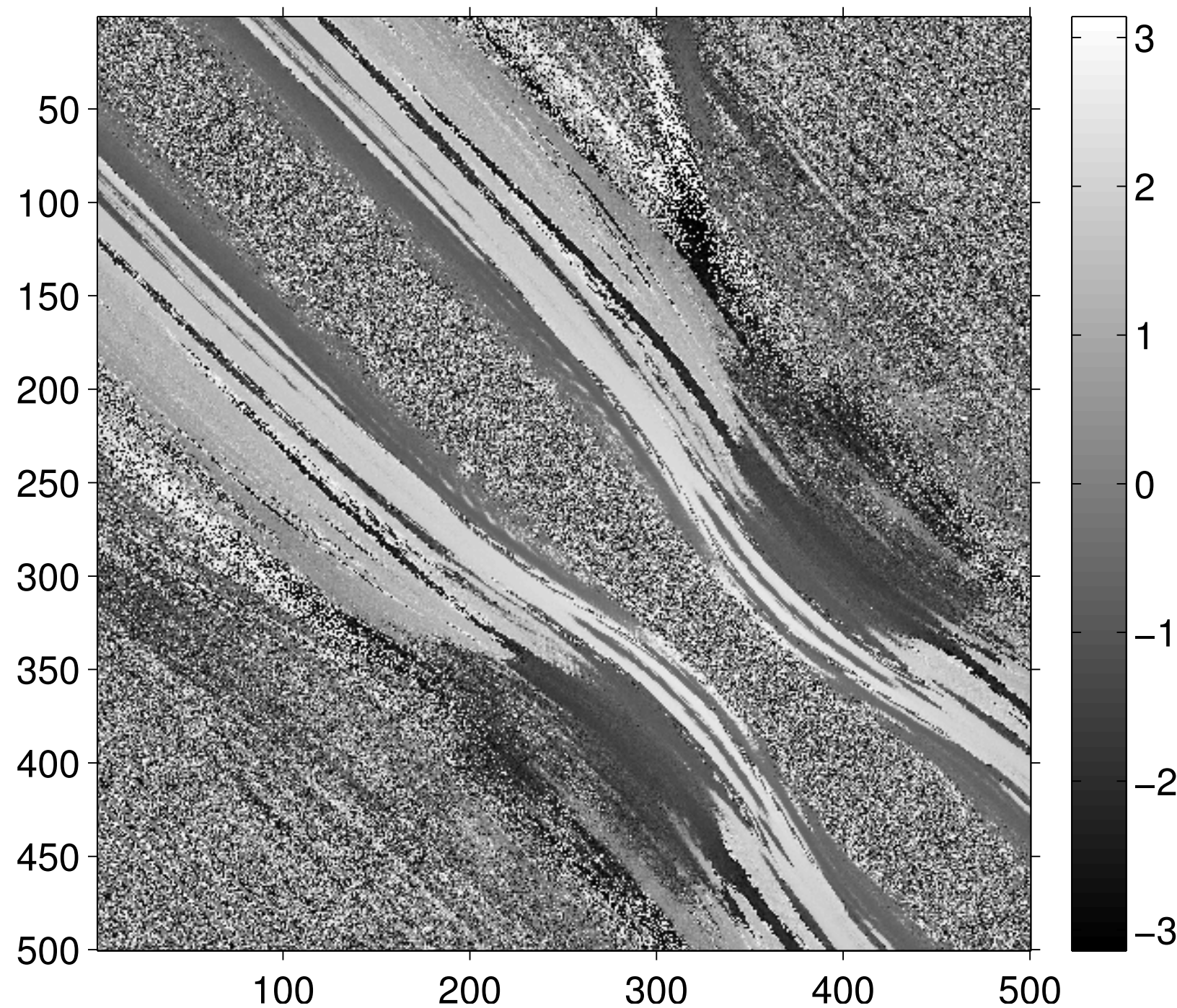


**3 Hz**



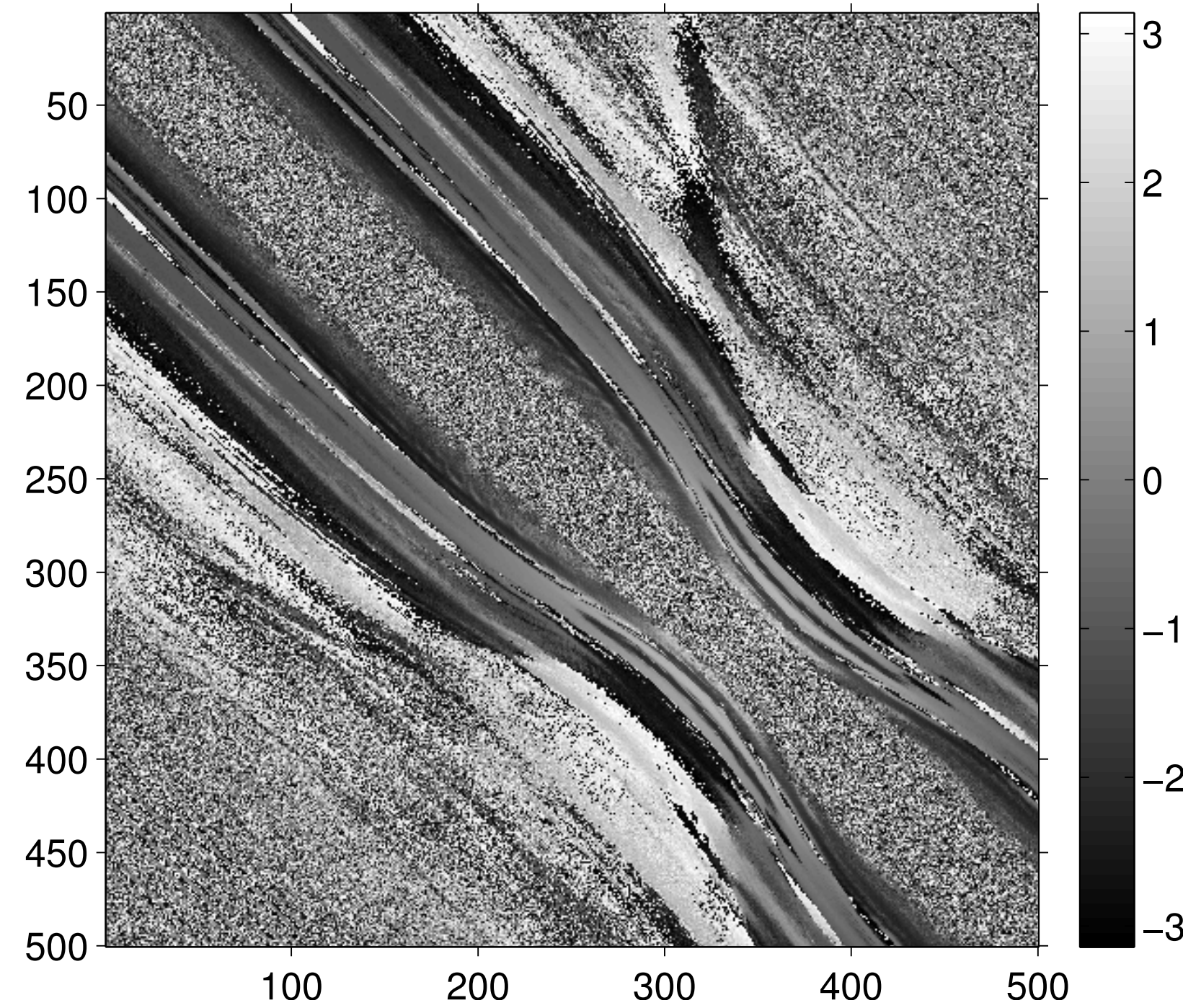
# Phase 8dB Pink noise added, Low-cut at 5 Hz

0.5Hz, 8dB SNR (pink noise), High-passed @ 5Hz solution



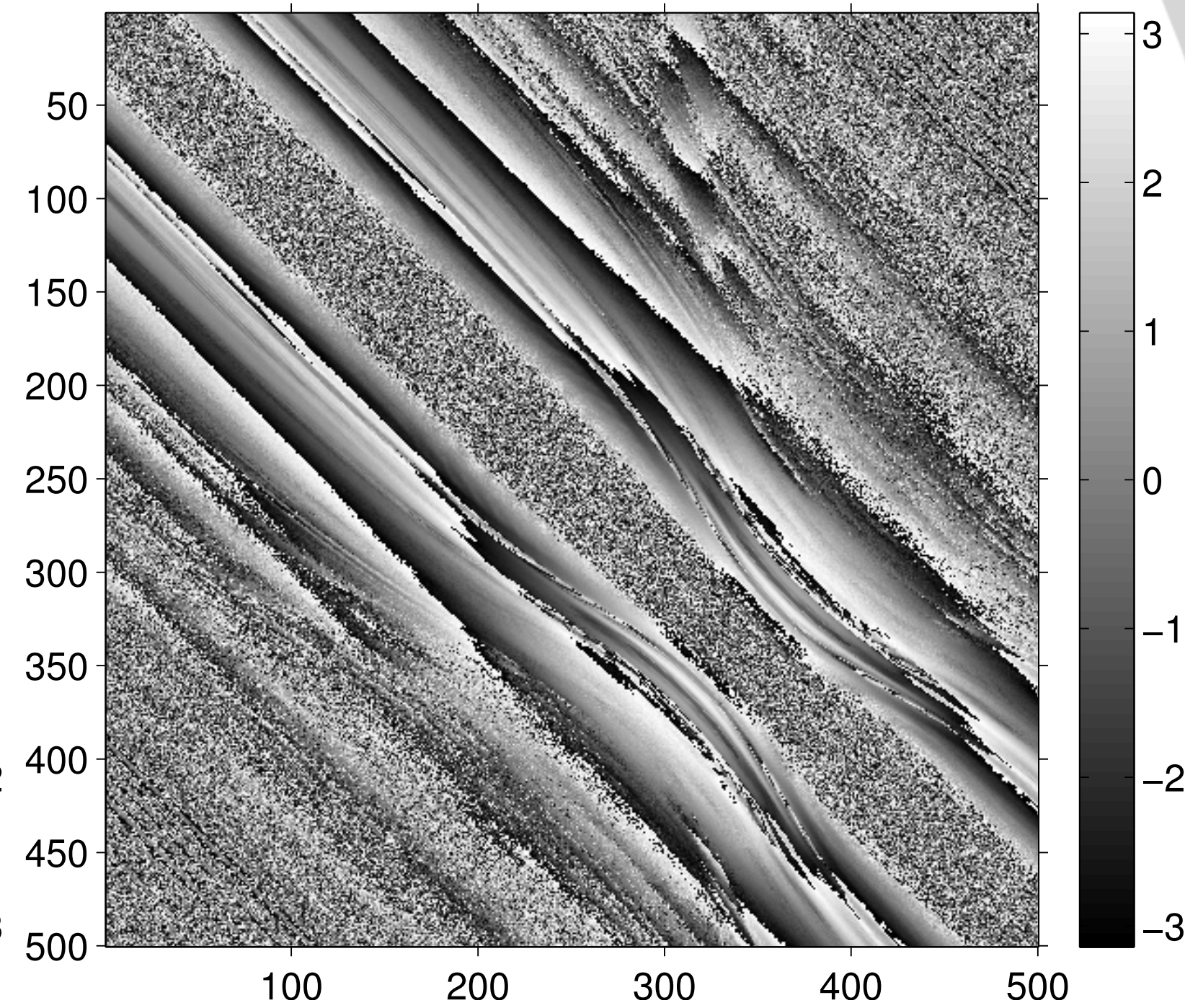
**0.5Hz**

1Hz, 8dB SNR (pink noise), High-passed @ 5Hz solution



**1 Hz**

3Hz, 8dB SNR (pink noise), High-passed @ 5Hz solution



**3 Hz**



## Remaining areas of investigation for REPSI

At coarsest levels, use more advanced/costly sparsifying methods? Go grid-free in the time domain? (*i.e., super-resolution methods*)

More sophisticated data-update method, less prone to local minima (use correlation between P and G, etc)

Incorporate up/down decomposition operator to work on P & Vz data

Potentially extract low-frequency information from G for diving-wave full-waveform inversion

# Acknowledgements

Gratitude to everyone who have given me with advice on multiple removal over the years

Thank YOU! This talk had **117** slides!



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